

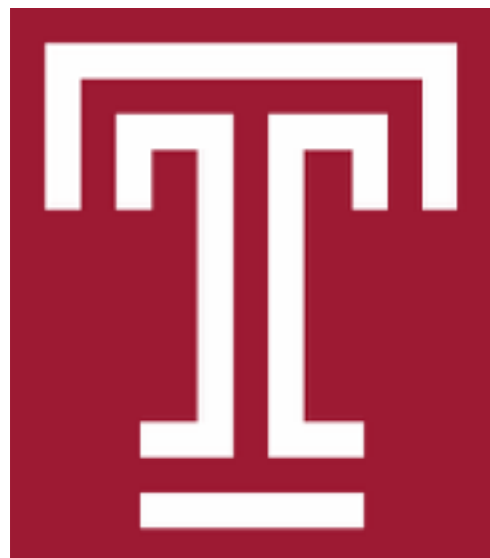


Precision



Electroweak Experiments at GeV Energies

*Parity Violating Electron Scattering
at Jefferson Lab*



Jim Napolitano
Temple University
Yale Physics Club
1 Nov 2021



What is an “Elementary Particle?”



Today's Players

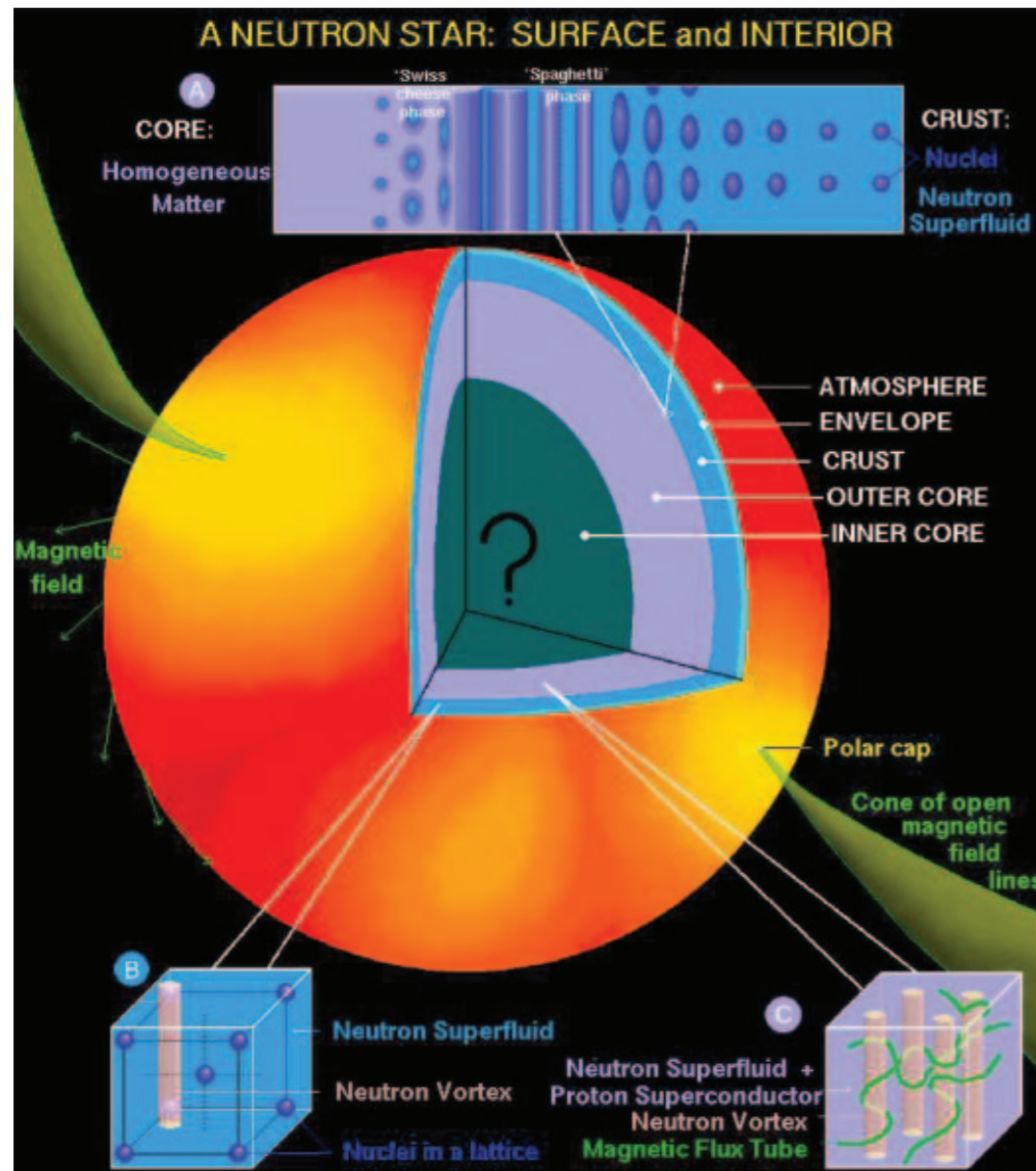


Three generations of matter (fermions)

	I	II	III	
mass	2.4 MeV/c ²	1.27 GeV/c ²	171.2 GeV/c ²	0
charge	2/3	2/3	2/3	0
spin	1/2	1/2	1/2	1
name	u up	c charm	t top	γ photon
	1.8 MeV/c ²	104 MeV/c ²	4.2 GeV/c ²	0
	1/3	-1/3	-1/3	0
	1/2	1/2	1/2	1
Quarks	d down	s strange	b bottom	g gluon
	<2.2 eV/c ²	<0.17 MeV/c ²	<15.5 MeV/c ²	91.2 GeV/c ²
	0	0	0	0
	1/2	1/2	1/2	1
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	Z⁰ Z boson
	511 MeV/c ²	105.7 MeV/c ²	1.777 GeV/c ²	80.4 GeV/c ²
	1	-1	-1	±1
	1/2	1/2	1/2	1
Leptons	e electron	μ muon	τ tau	W[±] W boson
				Gauge bosons

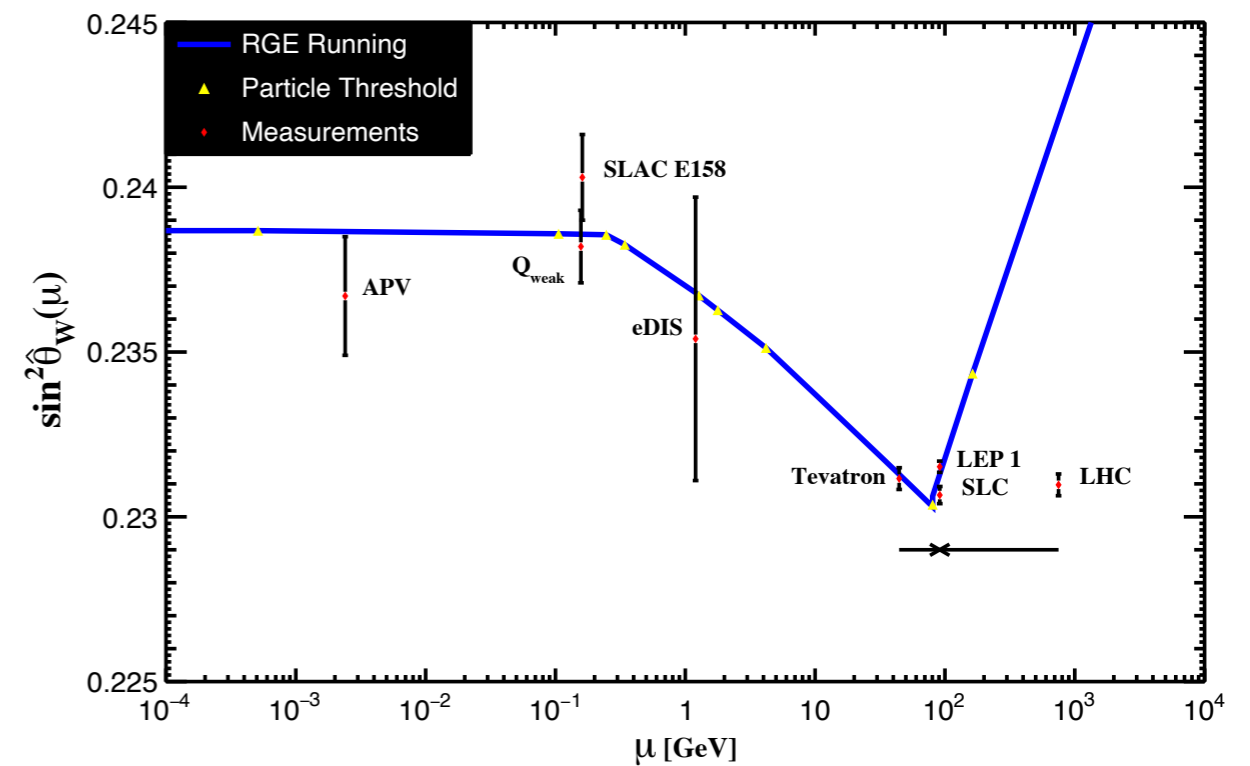
Today's Physics Topics

(1) Neutron Stars



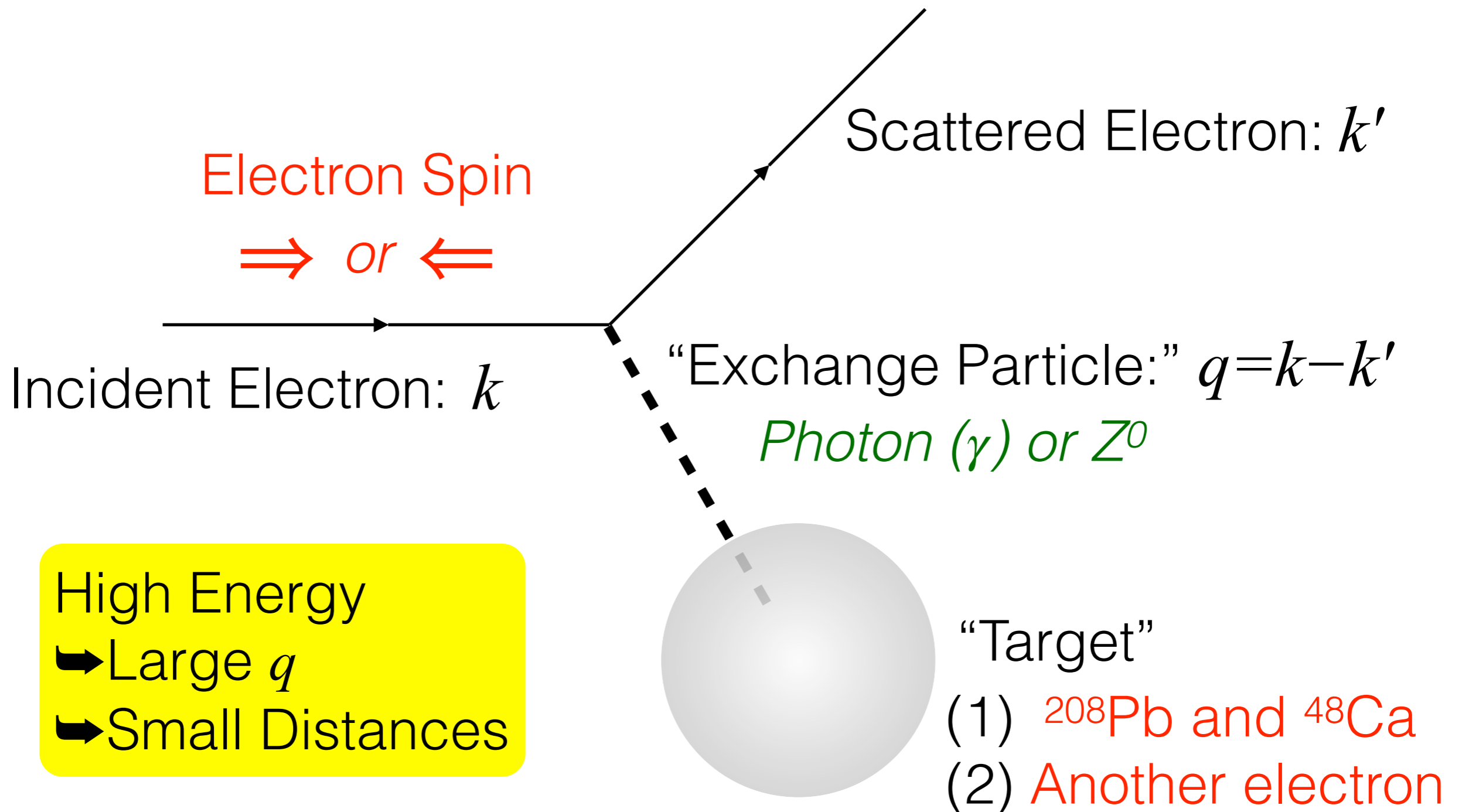
Science 304(2004)536

(2) Beyond the Standard Model



PDG 2020

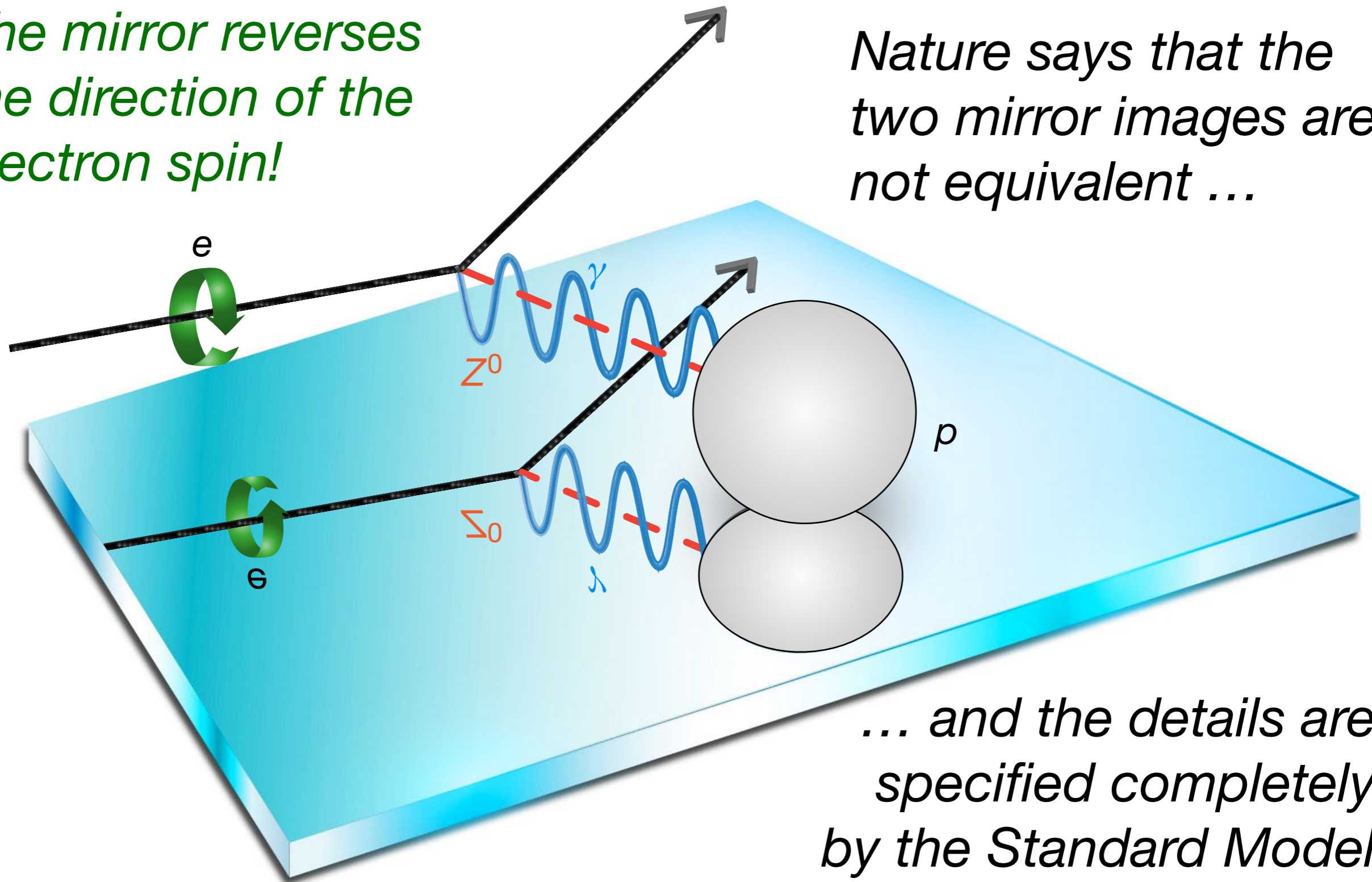
The Tool: Electron Scattering



Parity Violating Electron Scattering

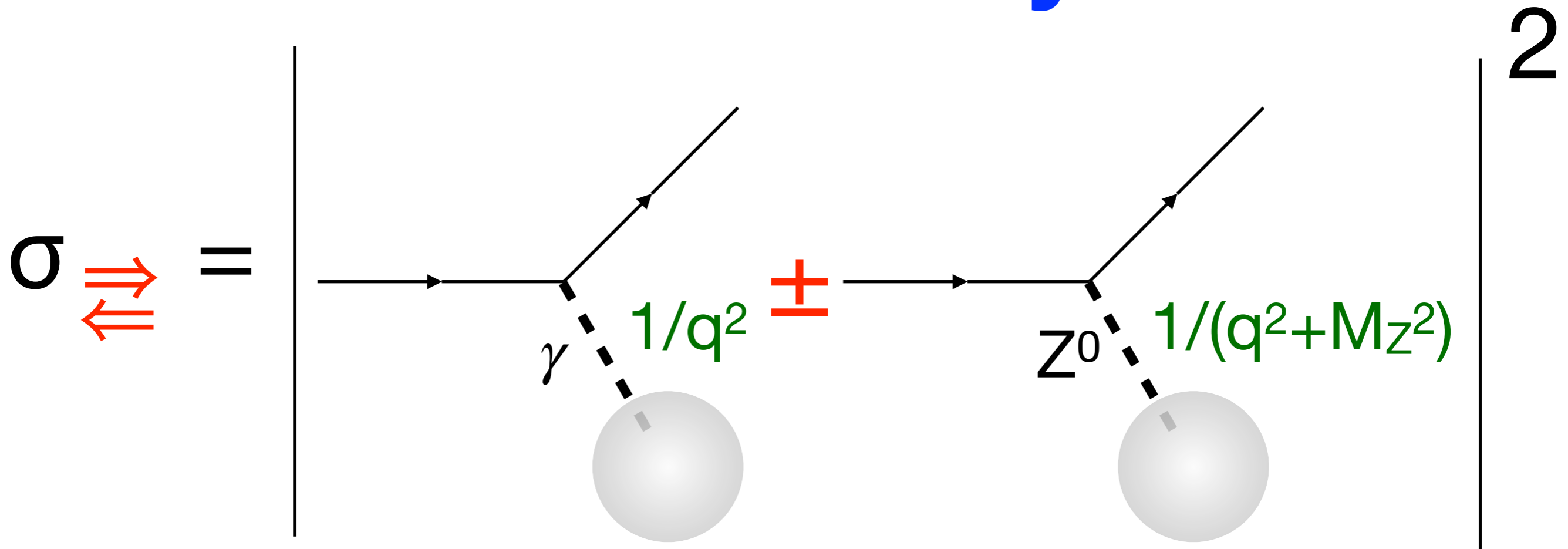
The mirror reverses the direction of the electron spin!

Nature says that the two mirror images are not equivalent ...



... and the details are specified completely by the Standard Model

The Effect is Very Small!



$$A_{PV} = \frac{\sigma_{\Rightarrow} - \sigma_{\Leftarrow}}{\sigma_{\Rightarrow} + \sigma_{\Leftarrow}} \approx \frac{q^2}{M_Z^2} \approx \text{“ppb”}$$

... where details depend on the interaction with the target!

Axial Vector Couplings of Z^0

Getting at the “Physics”

Axial Vector/Vector Interference is the key!

➡ What are the different Z^0 axial couplings?

Electron	$+1 - 4 \sin^2\theta_W$
Up quark (u)	$-1 + (8/3) \sin^2\theta_W$
Down quark (d)	$+1 - (4/3) \sin^2\theta_W$
Proton = 2u+d	$-1 + 4 \sin^2\theta_W$
Neutron = u+2d	$+1$

Note: $\sin^2\theta_W$ is close to 1/4

➡ Electron and proton couplings are small !

Some Previous ePV Experiments

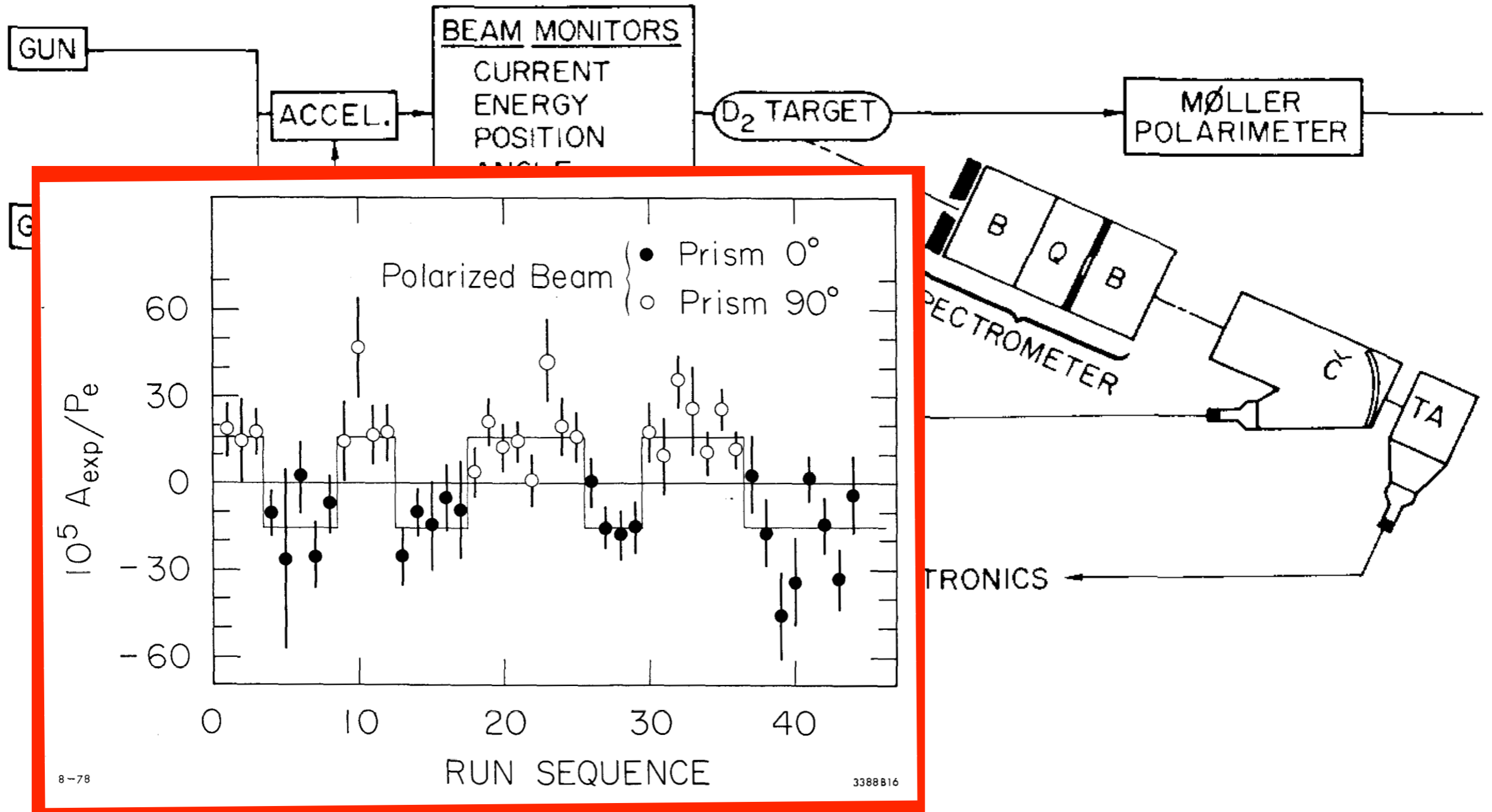
All asymmetries given in parts per billion (ppb)

Year	Lab	Target	A_{PV}	δA_{PV}	Comments
1978	SLAC	H ₂ (DIS)	49,000	8000	Confirmation of the Standard Model
1990	Bates	¹² C (Elastic)	600	140	Isolation of one term in the current
1997	Bates	p (Elastic)	6340	160	SAMPLE: Strangeness in the nucleon
2005	SLAC	e ⁻ (Møller)	117	16	E158: Test of running value of $\sin^2\theta_w$
2006	JLab	⁴ He (Elastic)	5981	810	HAPPEX- ⁴ He: Strange nucleon form factor
2007	JLab	p (Elastic)	1500	110	HAPPEX-II: Strange form factor of proton
2012	JLab	²⁰⁸ Pb (Elastic)	584	53	PREX-I: Neutron skin thickness
2014	JLab	H ₂ (DIS)	80,000	3000	PVDIS: Precision & wide kinematic range
2018	JLab	p (Elastic)	227	9	QWeak: Precision at very low q^2

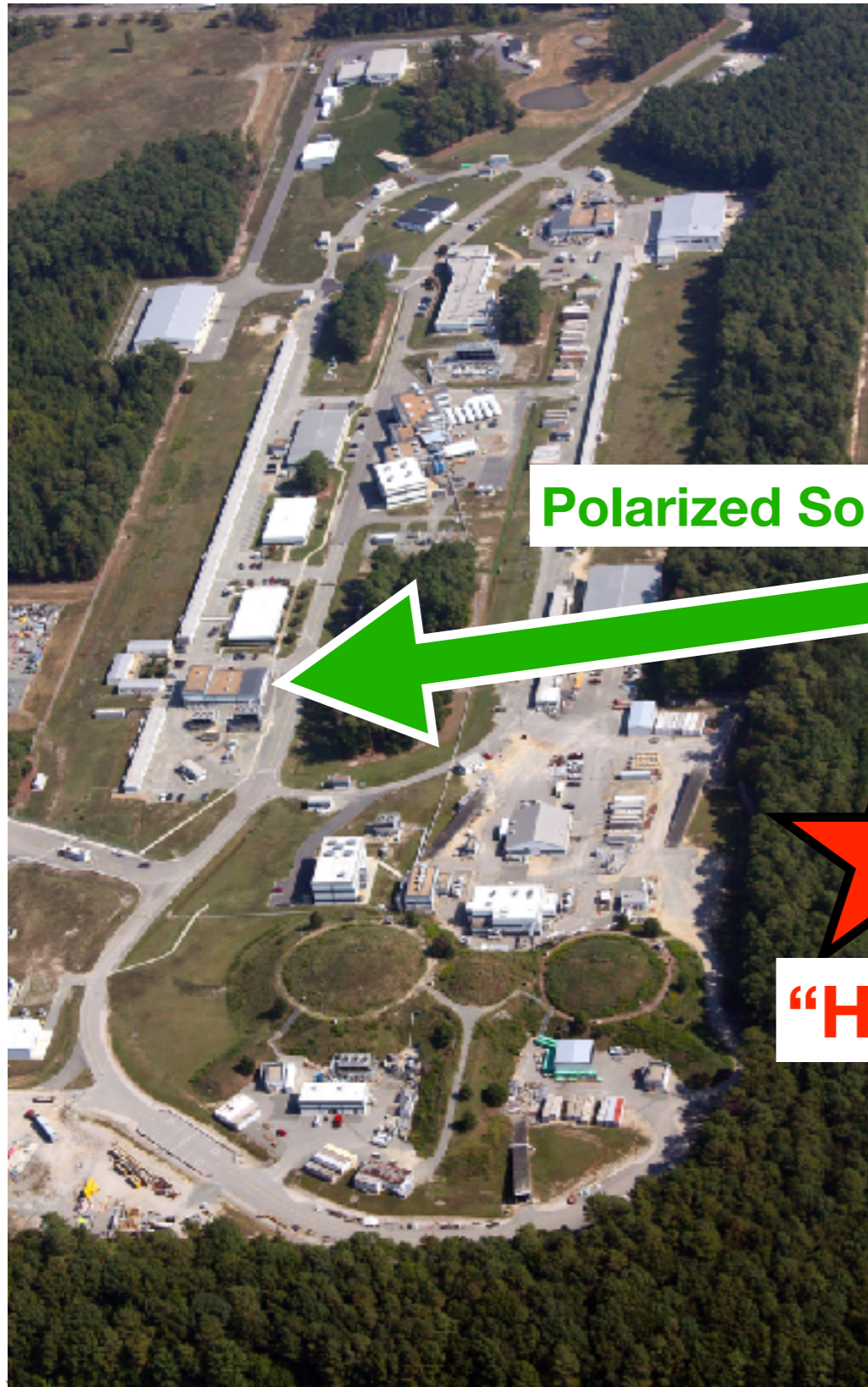


Example: SLAC (1978)

C. Prescott, et al, Phys. Lett. 77B (1978) 347



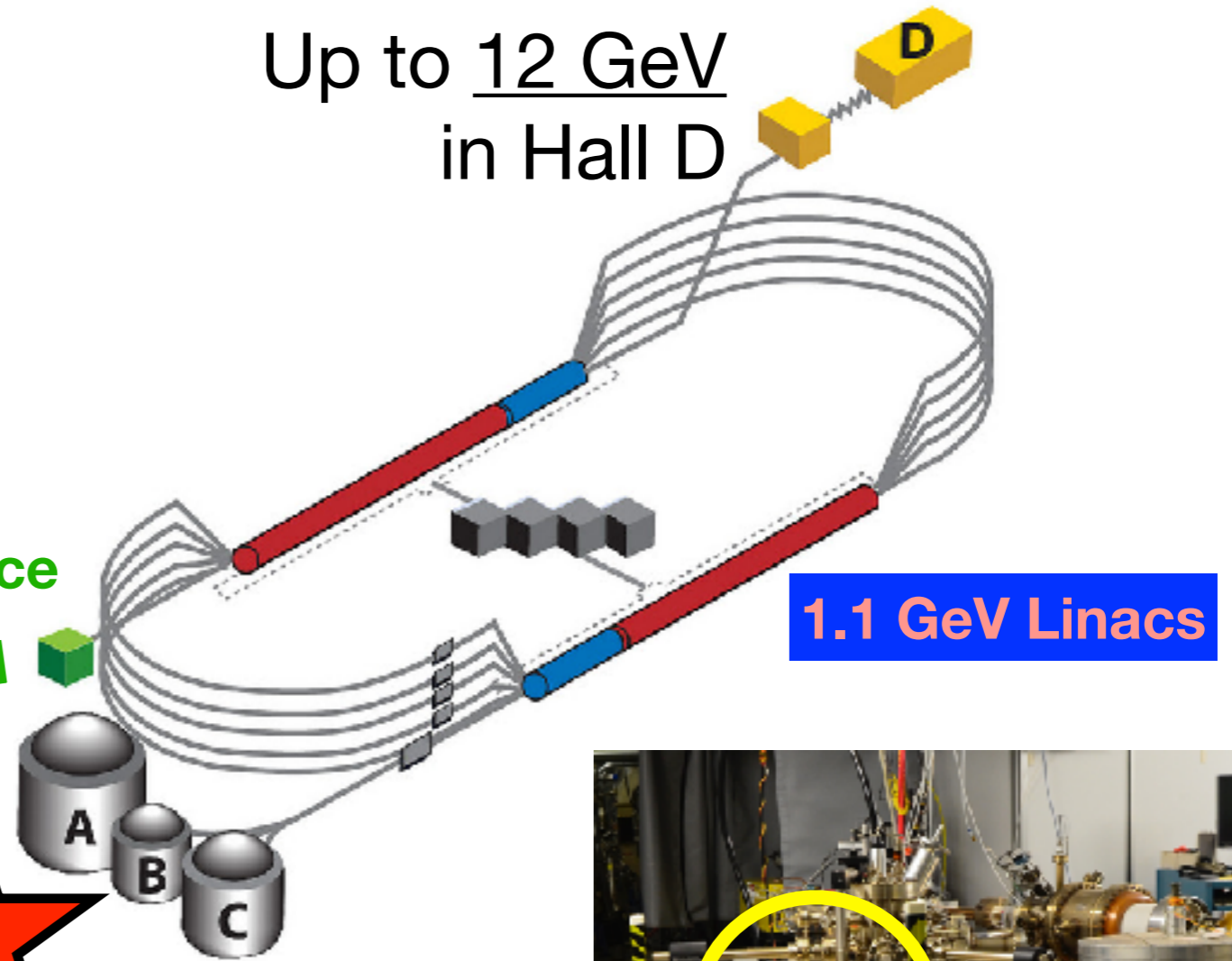
CEBAF @ Jefferson Lab



Polarized Source

“Hall A”

Up to 12 GeV
in Hall D

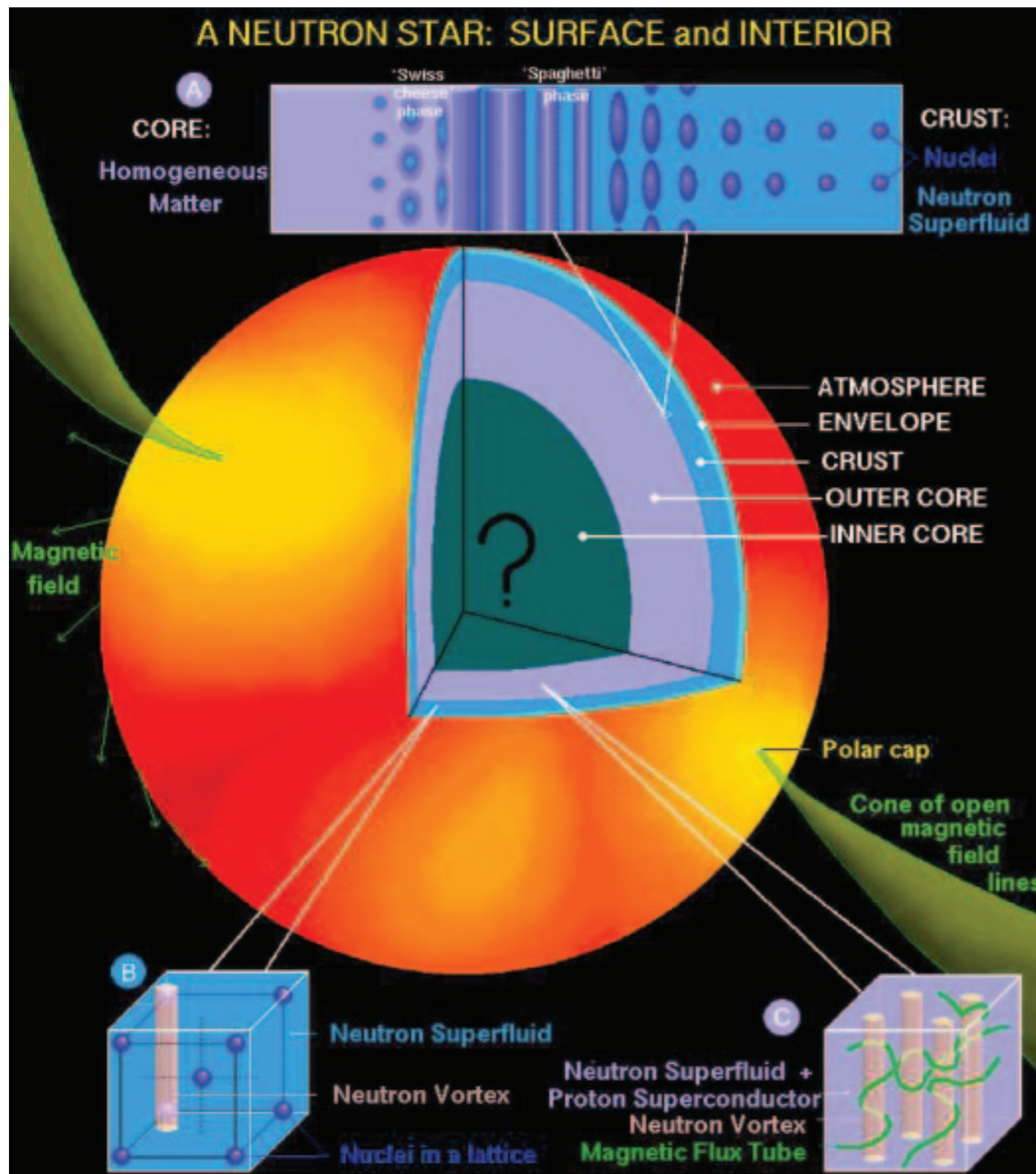


1.1 GeV Linacs



“TJ”

(1) Neutron Stars



Using neutron-rich nuclei to learn about neutron matter.

Can this help us to understand some of the mysteries about neutron stars?

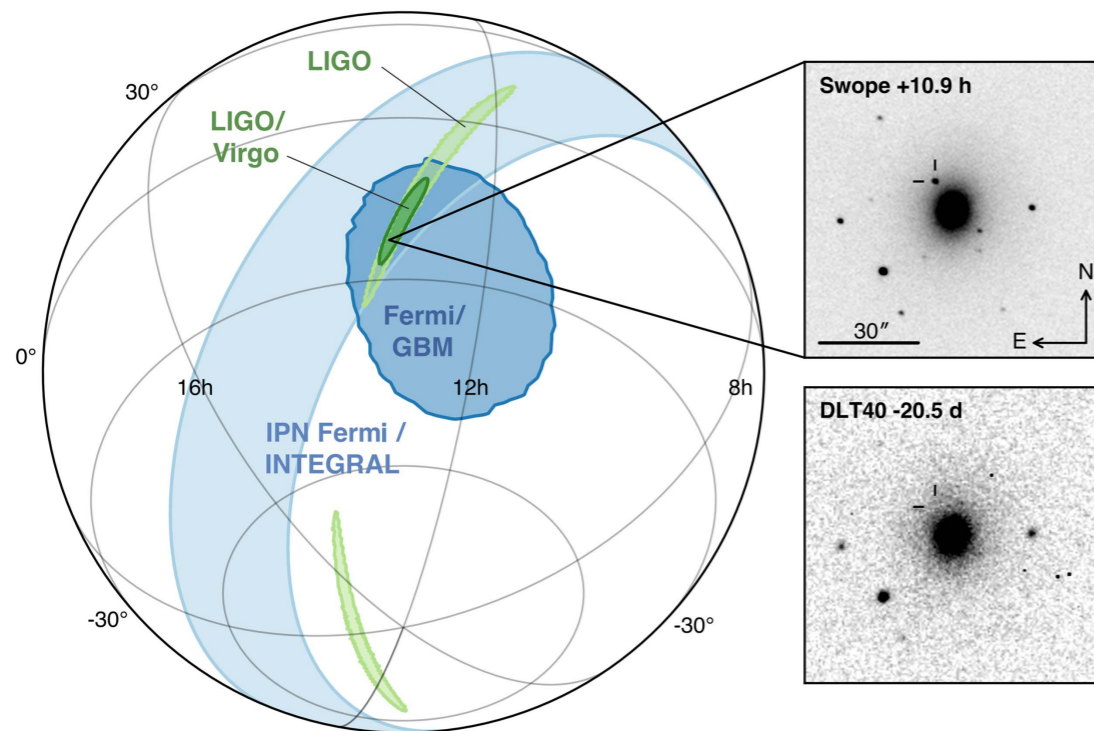
Neutron Stars in the News

Neutron Star Binary Mergers in the Multi-Messenger Age

GW170817:

PRX 9 (2019)011001

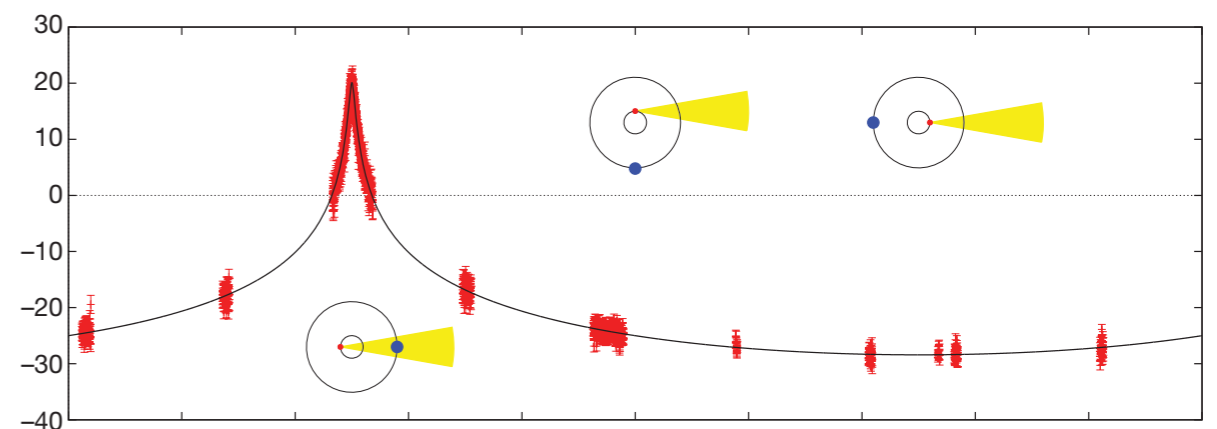
ApJ 848 2017 L12



Discovery of Neutron Stars with Mass >2 Solar Masses

PSR J1614-2230:

Nature 467(2010)1081



Also

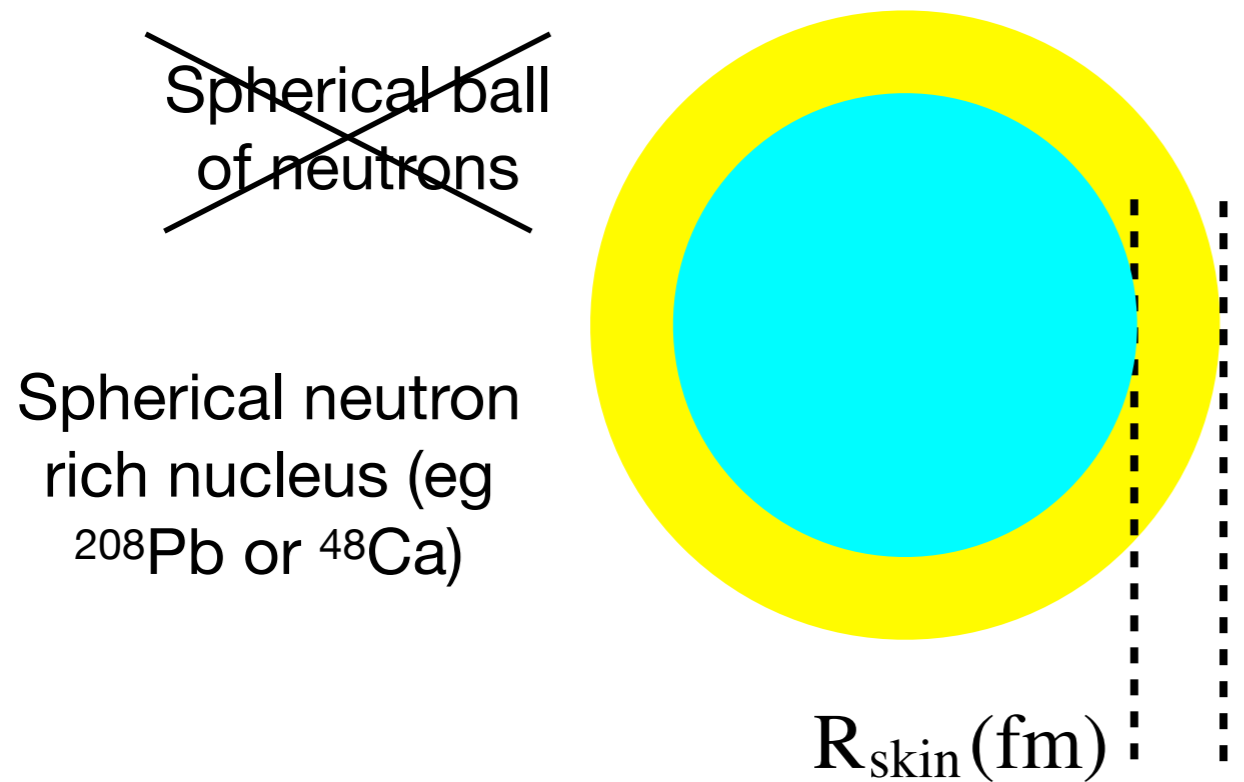
PSR J0348+0432:

Science 340(2013)448

➡ It is imperative to learn about neutron matter

Neutron Stars & Neutron Skins

See Reed, et al, Phys.Rev.Lett 126(2021)172503

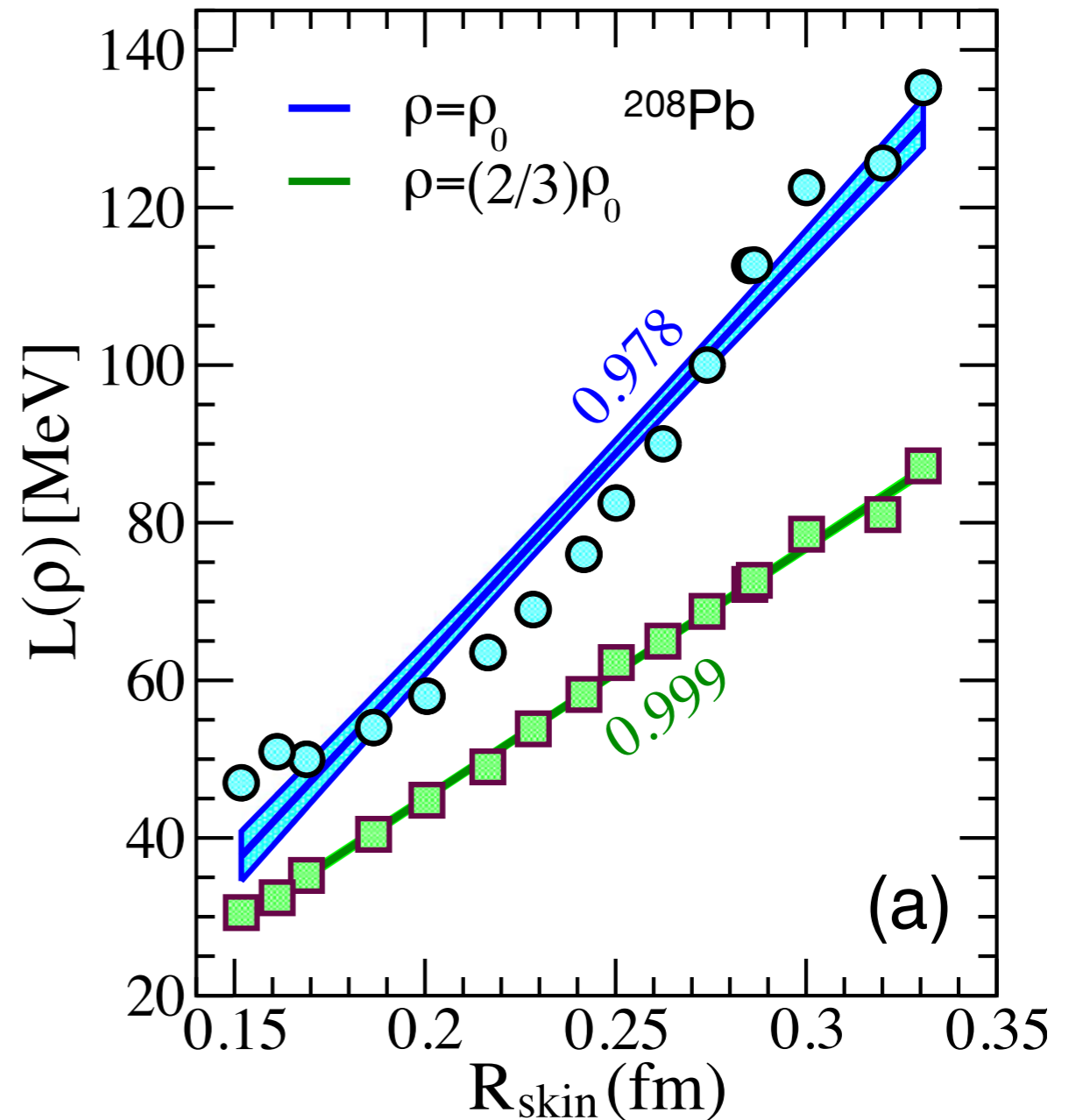


“Liquid Drop”

$$\frac{E}{A} - M \approx \mathcal{E}_{\text{SNM}} + \left(\frac{\rho_n - \rho_p}{\rho} \right)^2 S(\rho)$$

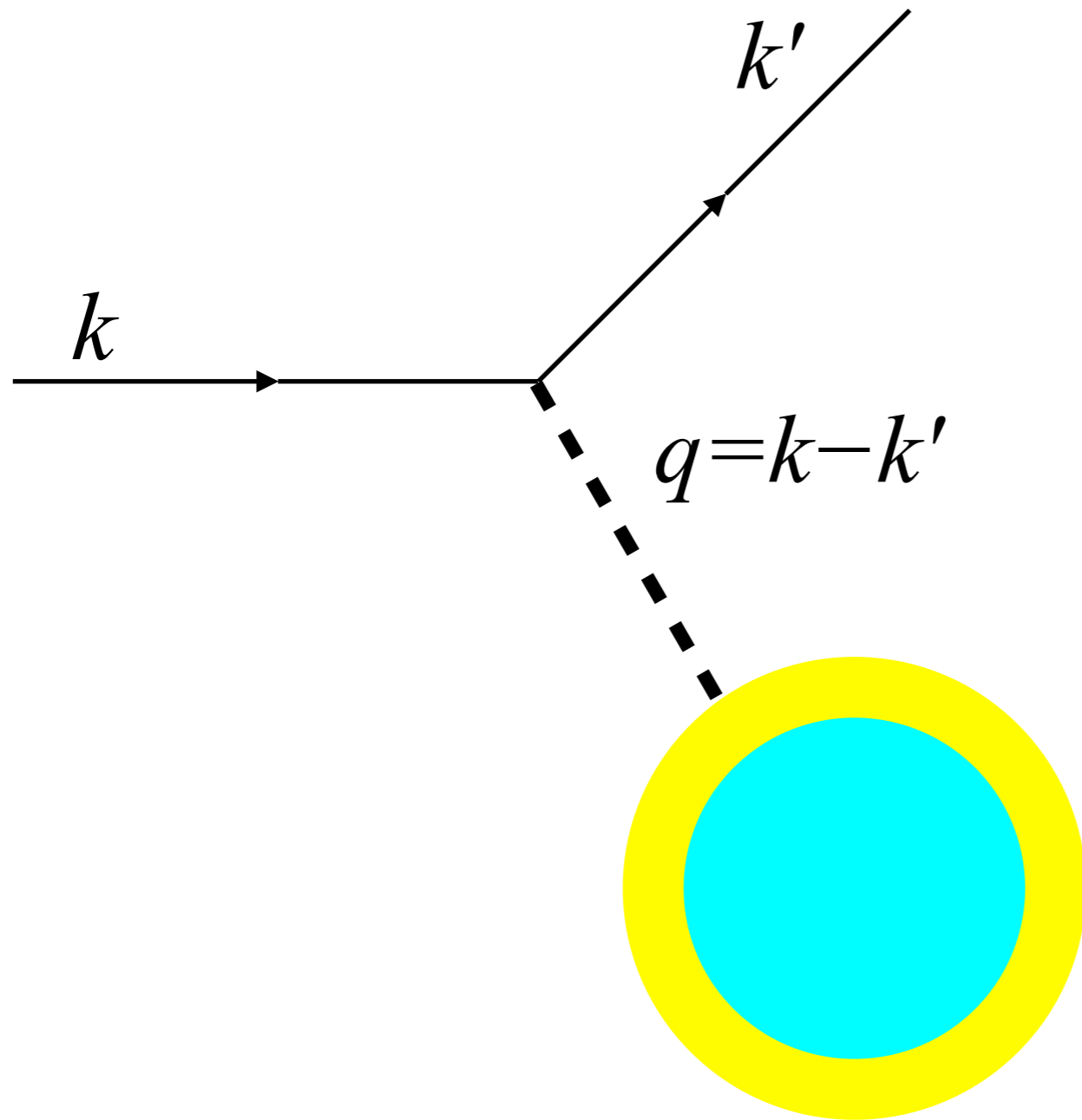
“Symmetry Energy”

$$S(\rho) = J + L(\rho) \frac{\rho - \rho_0}{3\rho_0}$$



The “Physics” is in $L(\rho)$

Elastic Scattering Form Factors

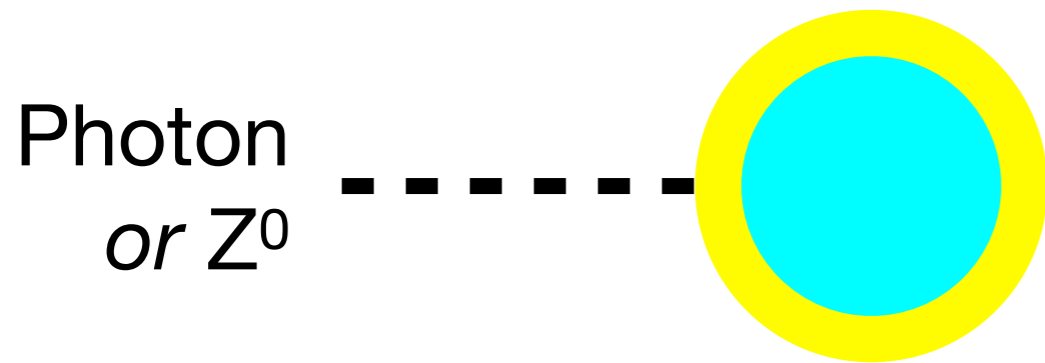


Scattering amplitude is proportional to

$$\begin{aligned} F(q) &= \int \rho(r) e^{i\mathbf{q}\cdot\mathbf{r}} dV \\ &= 1 - \frac{q^2}{6} \int r^2 \rho(r) dV + \dots \\ &= 1 - \frac{q^2}{6} \langle r^2 \rangle + \dots \end{aligned} \text{Note!}$$

The mean square radius $\langle r^2 \rangle$ gives the “radius” of the distribution $\rho(r)$.

Measuring the Neutron Skin



<u>Recall:</u>	<u>Photon</u>	<u>Z⁰</u>
Proton	1	Small
Neutron	≈0	1

$$A_{PV} = \frac{d\sigma(\Rightarrow) - d\sigma(\Leftarrow)}{d\sigma(\Rightarrow) + d\sigma(\Leftarrow)}$$

$$= \frac{G_F q^2 F_Z(q)}{4\pi\alpha F_\gamma(q)}$$

$$= \frac{G_F q^2}{4\pi\alpha} \left\{ 1 - \frac{q^2}{6} \left[\langle r_n^2 \rangle - \langle r_p^2 \rangle \right] + \dots \right\}$$

... but this term is small.

*Parity Violating
elastic scattering
directly measures
the neutron skin...*

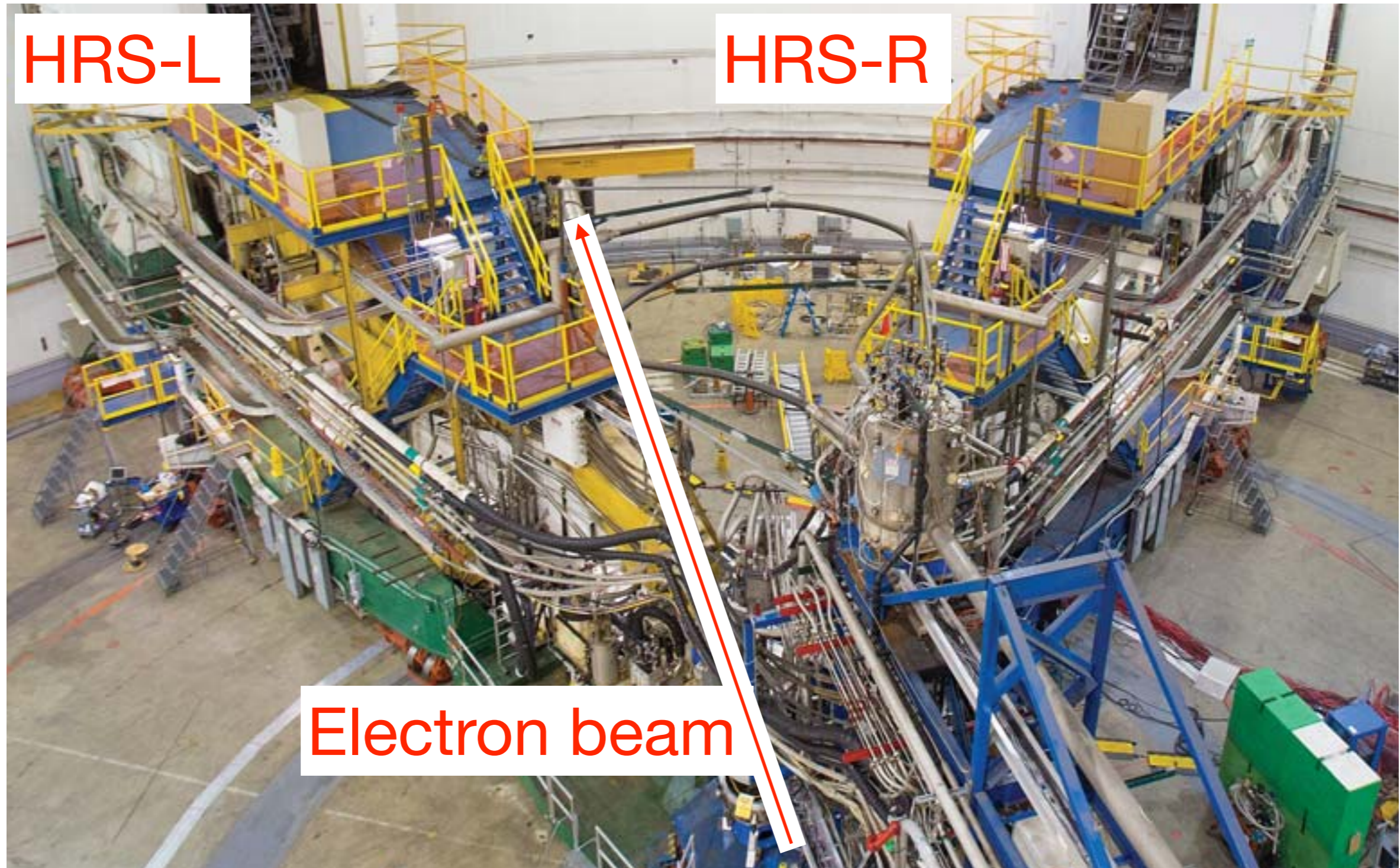
PREX and CREX

Energy	^{208}Pb	1.0 GeV	^{48}Ca	2.2 GeV
Angle		5 degrees		4 degrees
A_{PV}		0.6 ppm		2 ppm
1 st Ex. State		2.60 MeV		3.84 MeV
beam current		70 μA		150 μA
rate		1 GHz		100 MHz
run time		35 days		45 days
A_{PV} precision	9% (PREX-I)	3% (PREX-II)		2.4%
Error in R_N		0.06 fm (PREX-II)		0.02 fm

^{208}Pb : Large nucleus \Rightarrow nuclear matter

^{48}Ca : Microscopic calculation tractable

Hall A at Jefferson Lab



Put in Numbers for ^{208}Pb

Precision is the key!

$$A_{\text{PV}} = \frac{G_F q^2}{4\pi\alpha} \left\{ 1 - \frac{q^2}{6} \left[\langle r_n^2 \rangle - \langle r_p^2 \rangle \right] + \dots \right\}$$

$$\begin{aligned} \langle r_n^2 \rangle - \langle r_p^2 \rangle &= \left(\langle r_n^2 \rangle^{1/2} + \langle r_p^2 \rangle^{1/2} \right) \left(\langle r_n^2 \rangle^{1/2} - \langle r_p^2 \rangle^{1/2} \right) \\ &\approx \quad \quad \quad 2R_{\text{Pb}} \quad \quad \quad \Delta_{\text{Skin}} \end{aligned}$$

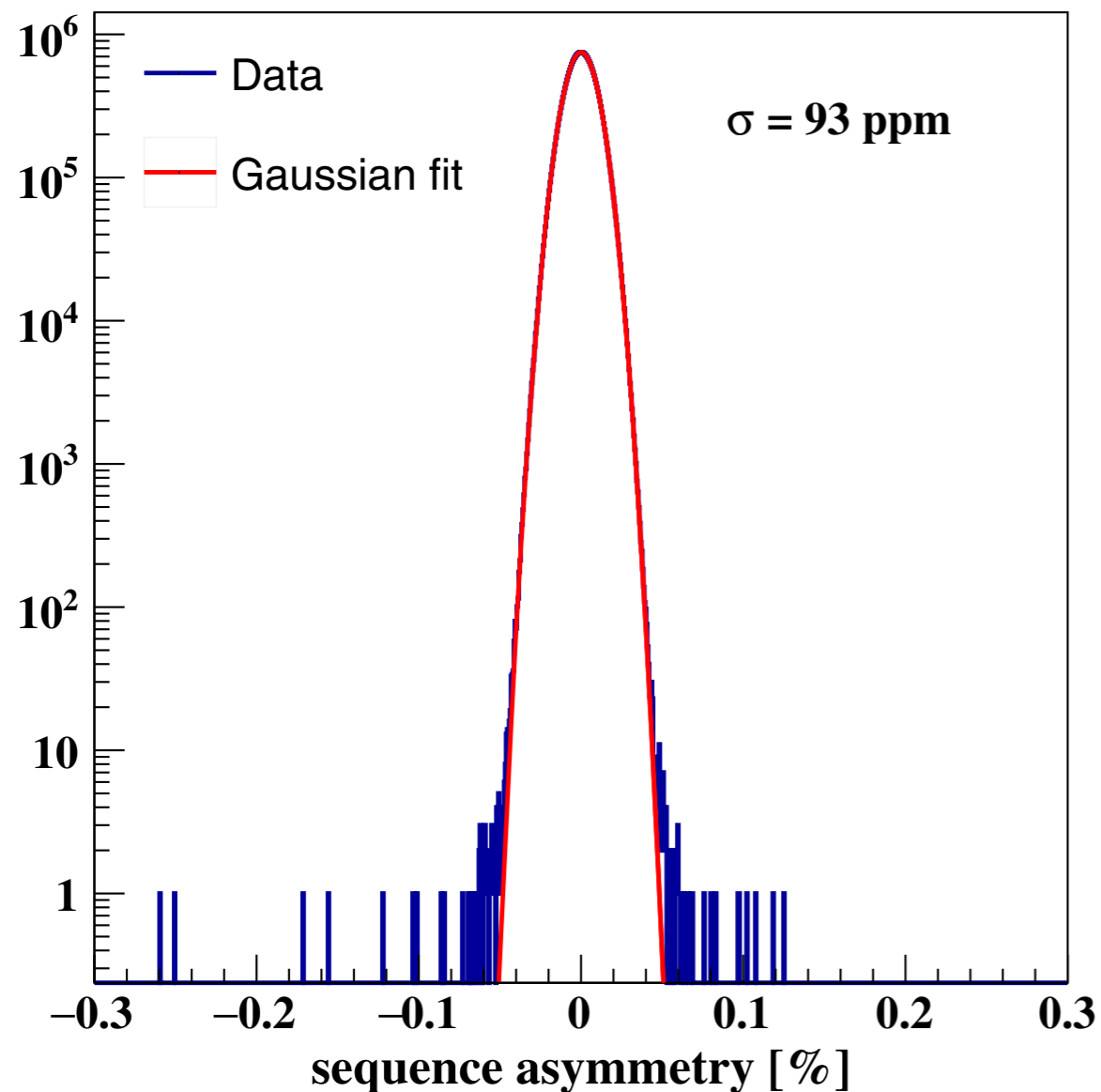
q^2	0.076 GeV ²
R_{Pb}	6 fm
Δ_{Skin}	~ 0.2 fm
$q^2 [\langle r_n^2 \rangle - \langle r_p^2 \rangle] / 6$	0.076
δA_{PV}	3%
$\delta \Delta_{\text{Skin}}$	0.07 fm

If we can measure the (small!) parity violating asymmetry itself to about 3%, then we should be able to determine the skin thickness with an uncertainty near 0.07 fm.

The PREX Result

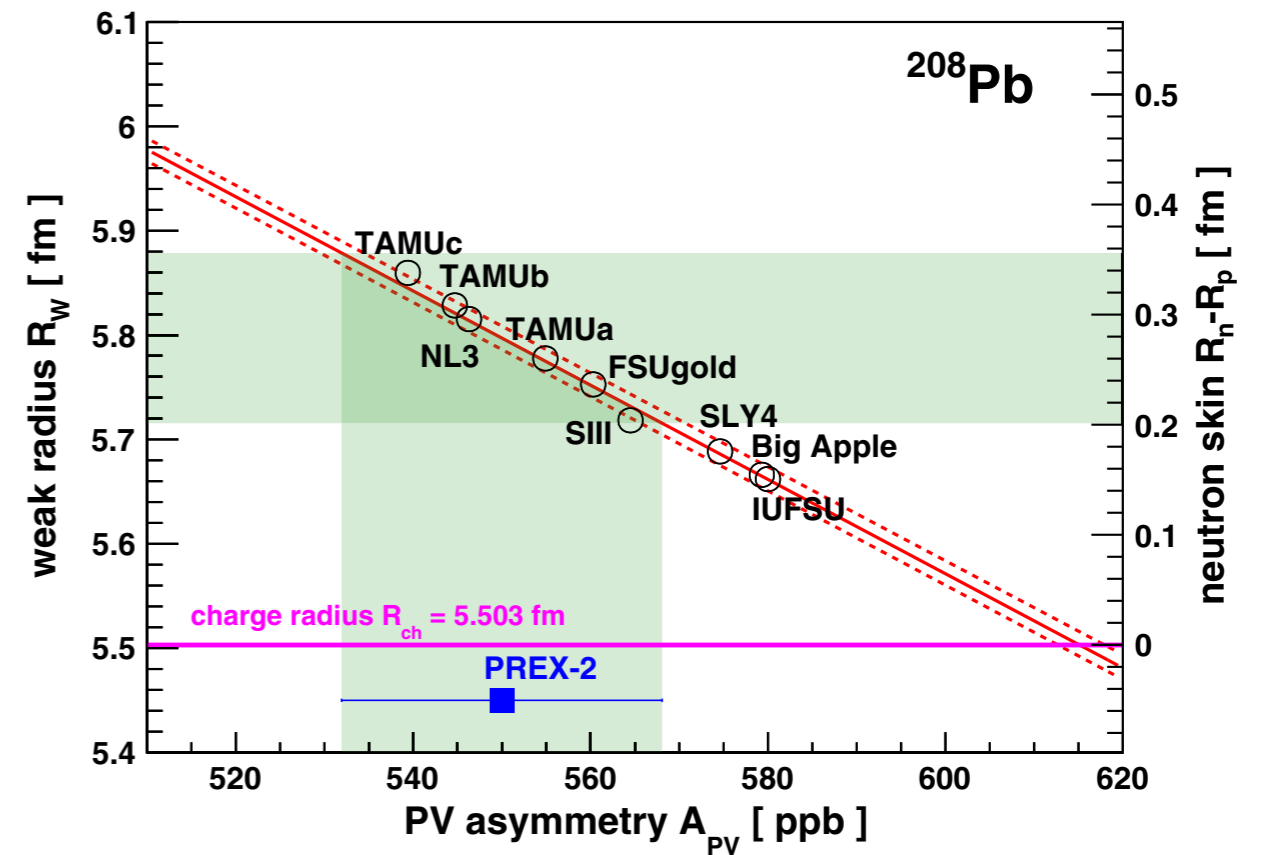
Adhikari, et al, Phys.Rev.Lett 126(2021)172502

3×10^6 1/30 sec @ 240 Hz



$$A_{PV}^{\text{meas}} = 550 \pm 16 \text{ (stat)} \pm 8 \text{ (syst)} \text{ ppb}$$

$$F_W(\langle Q^2 \rangle) = 0.368 \pm 0.013 \text{ (exp)} \pm 0.001 \text{ (theo)}$$



^{208}Pb Parameter	Value
Weak radius (R_W)	5.800 ± 0.075 fm
Interior weak density (ρ_W^0)	-0.0796 ± 0.0038 fm $^{-3}$
Interior baryon density (ρ_b^0)	0.1480 ± 0.0038 fm $^{-3}$
Neutron skin ($R_n - R_p$)	0.283 ± 0.071 fm

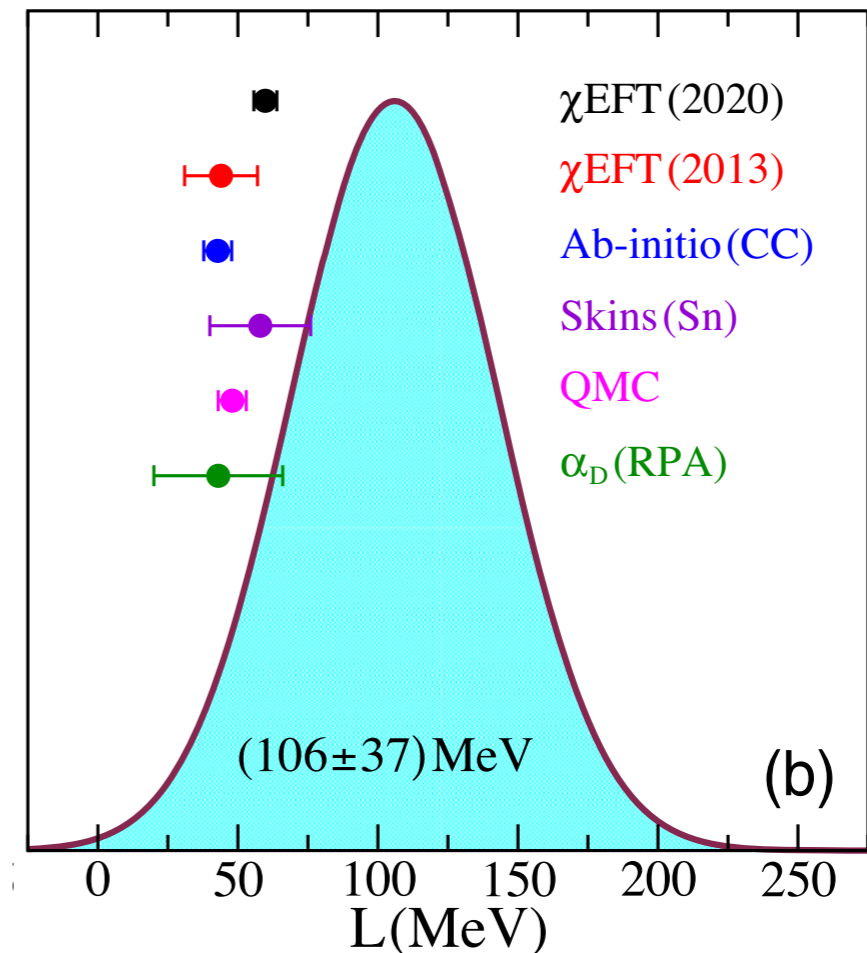
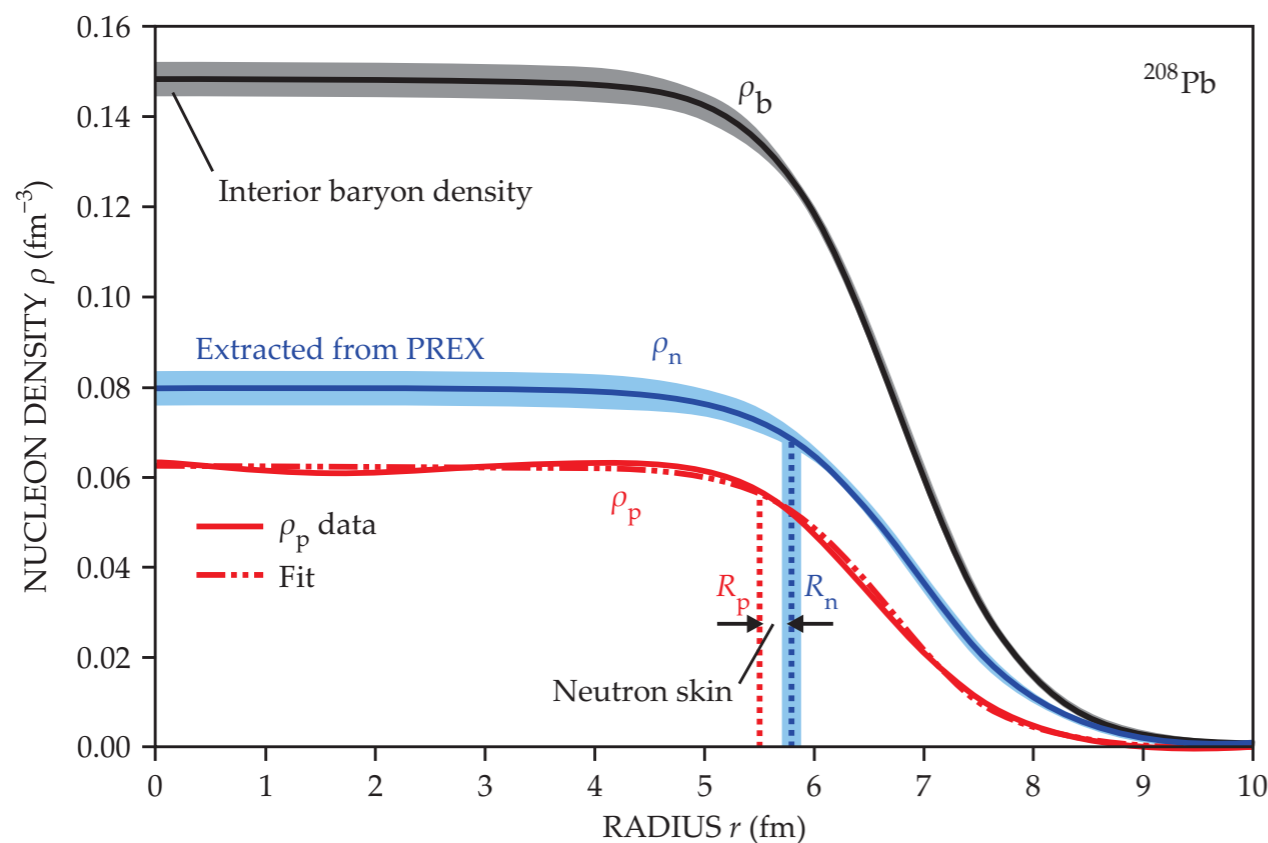
PREX Media Attention

Physics Today 74(2021)12 & K Scholberg <https://physics.aps.org/articles/v14/58>

SEARCH & DISCOVERY

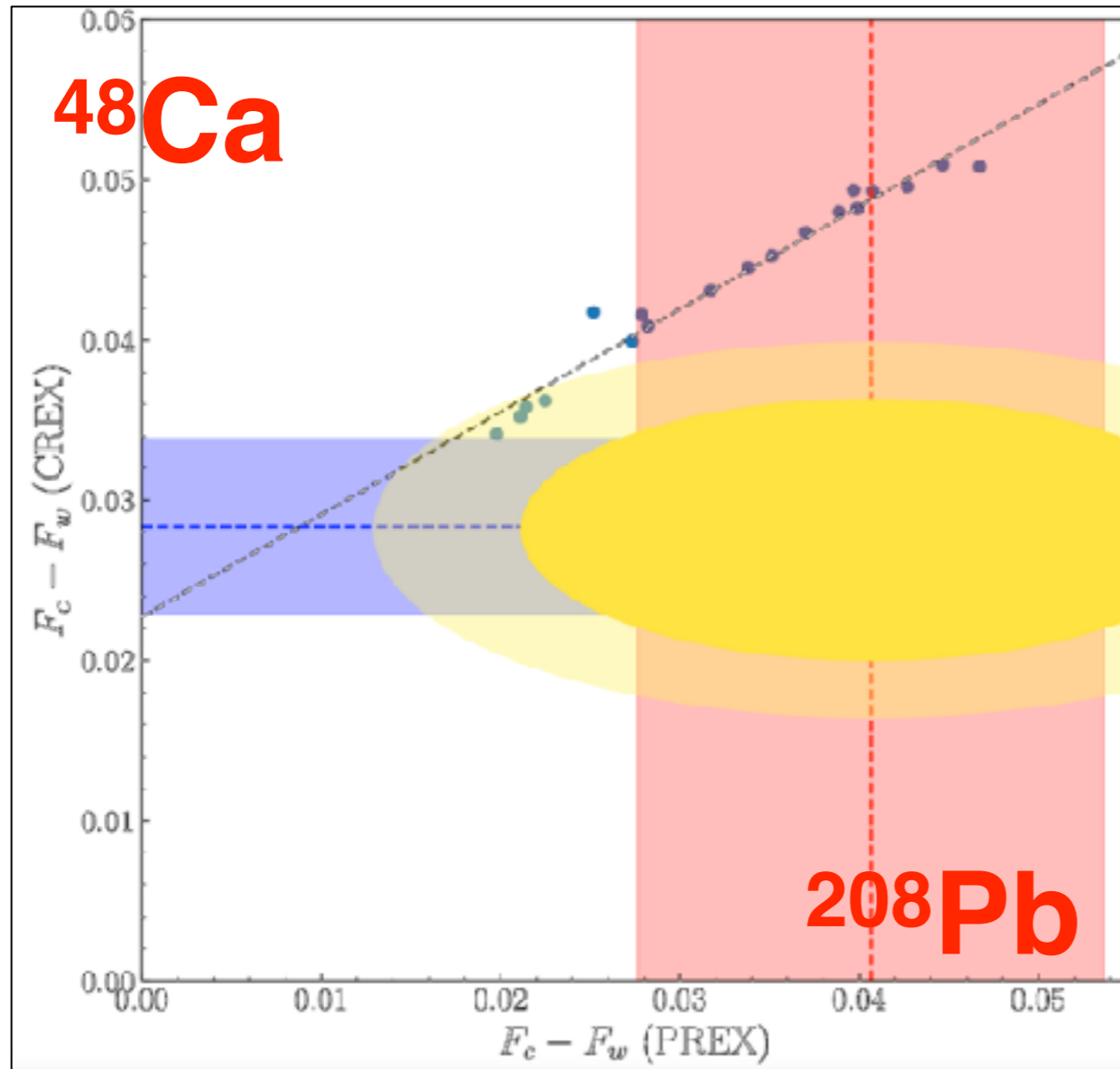
Lead-208 nuclei have thick skins

Nuclear matter calculations generally predicted a smaller value for the “pressure” L .



CREX Just Released

$$A_{pV} = 2658.6 \pm 113.2 \text{ ppb (4.3\%)}$$



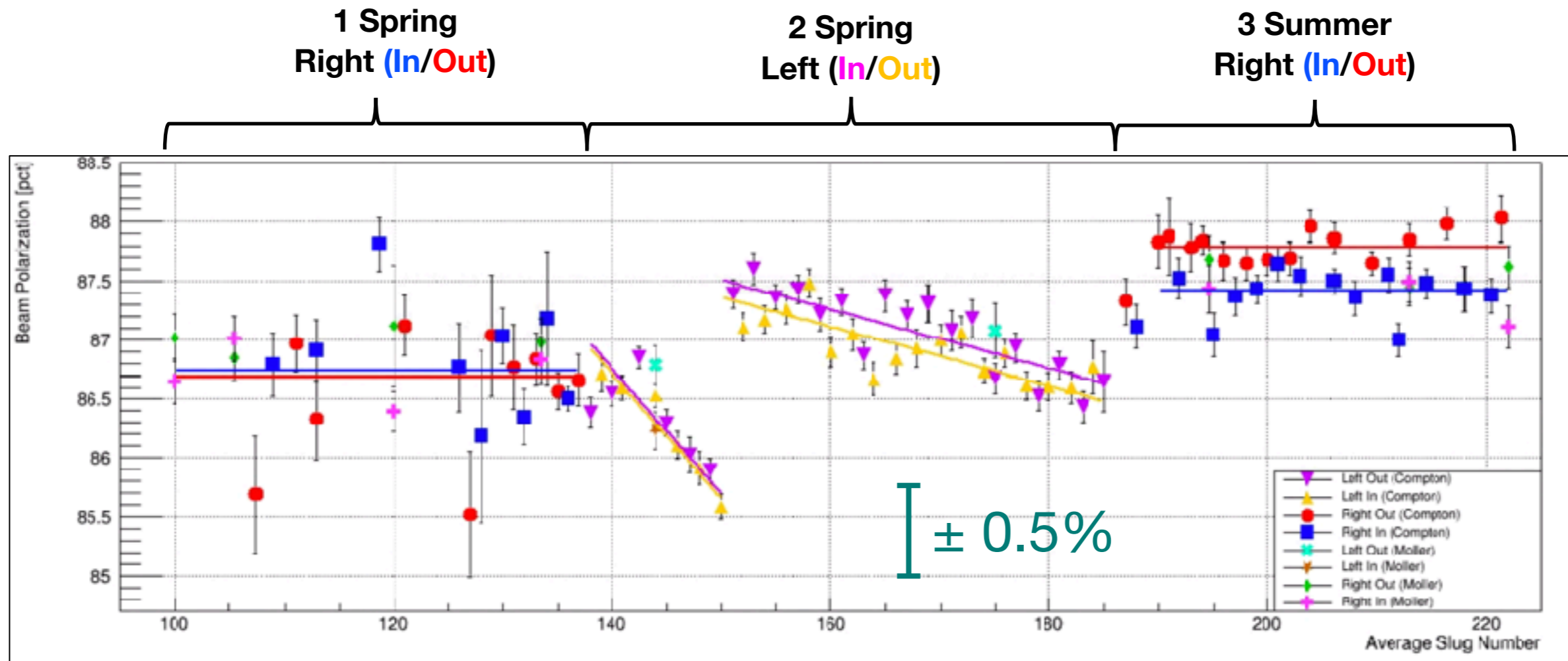
Points are a collection of Relativistic and Non-Relativistic DFT calculations

Compare to predictions of microscopic calculations?
eg G. Hagen et al, Nature Physics 12 (2016) 186

First need to sort out some issues, including a strong “weak spin orbit” effect that mainly effects ^{48}Ca
eg Horowitz and Piekarewicz, Phys Rev C86(2012)045503

Electron Beam Polarization

High Precision Necessary



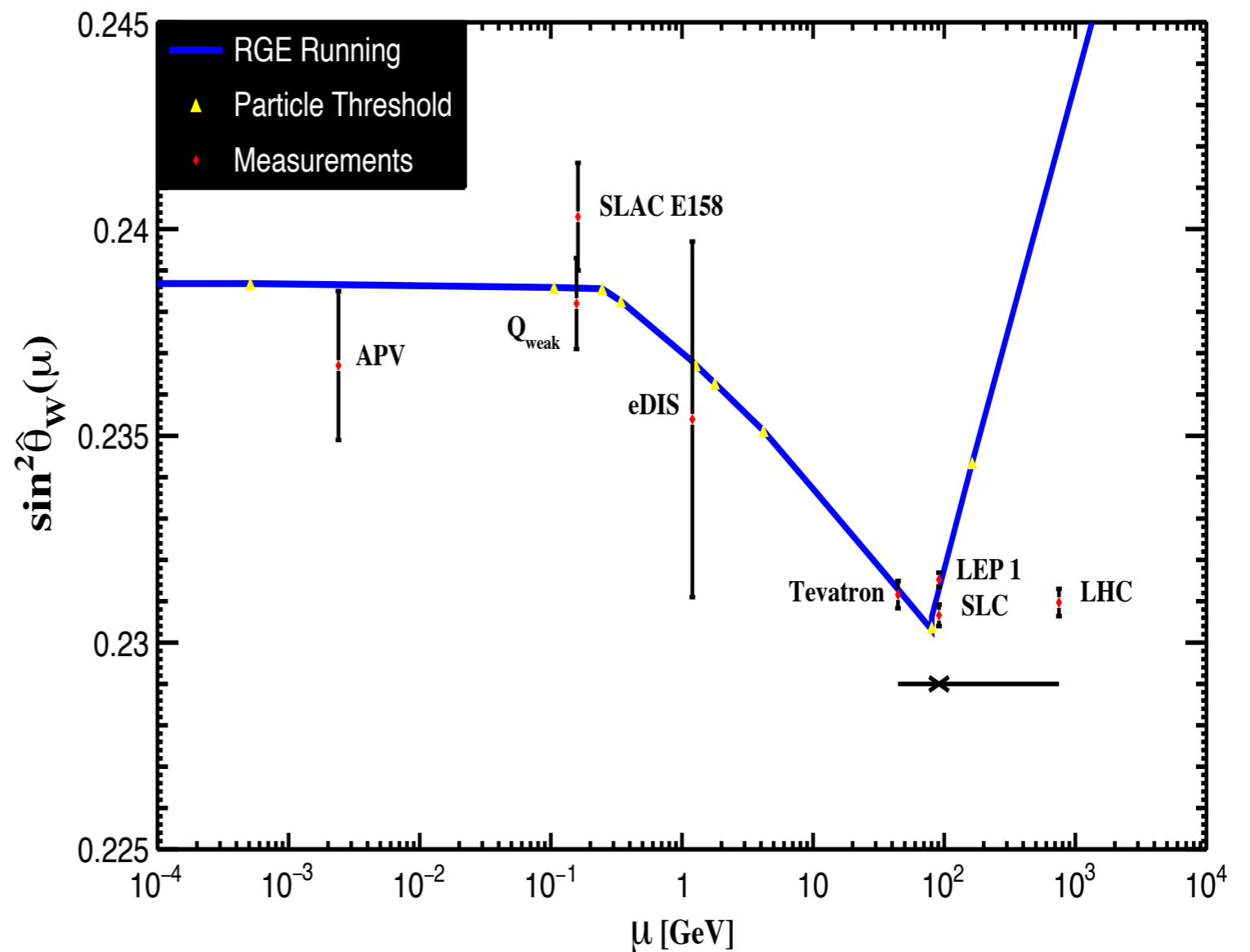
Two Approaches: Møller and Compton

$$P_e = 87.09 \pm 0.44\%$$

(2) Beyond the Standard Model

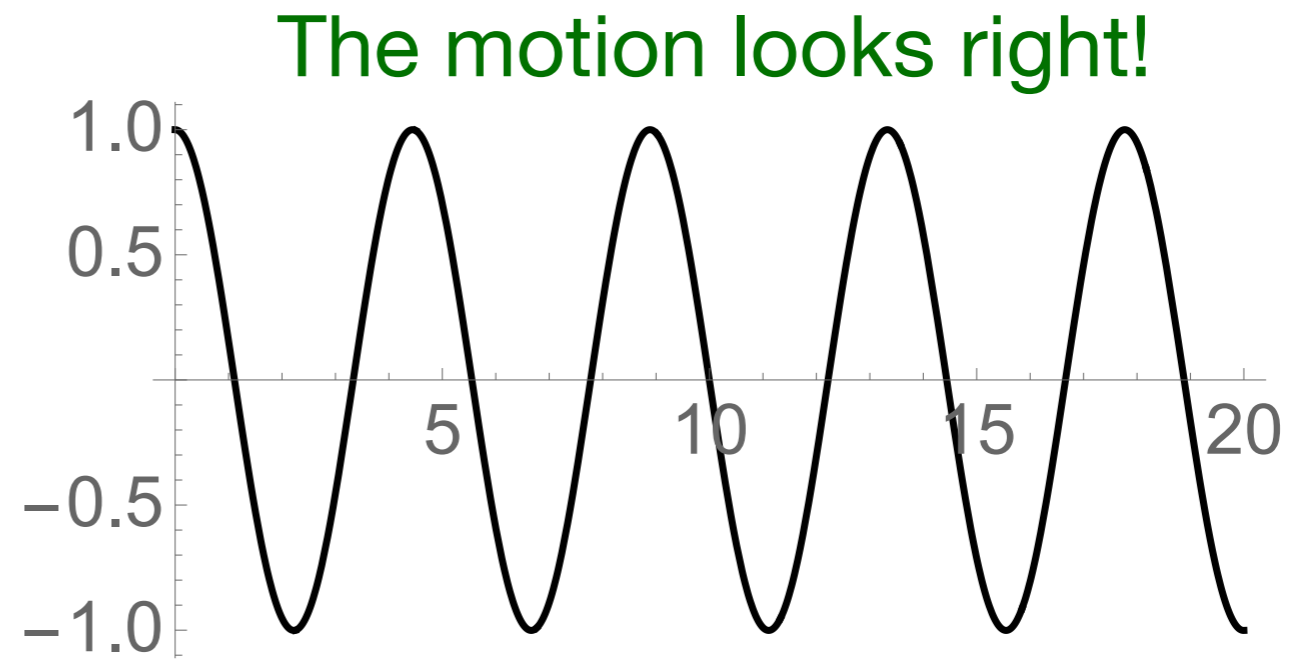
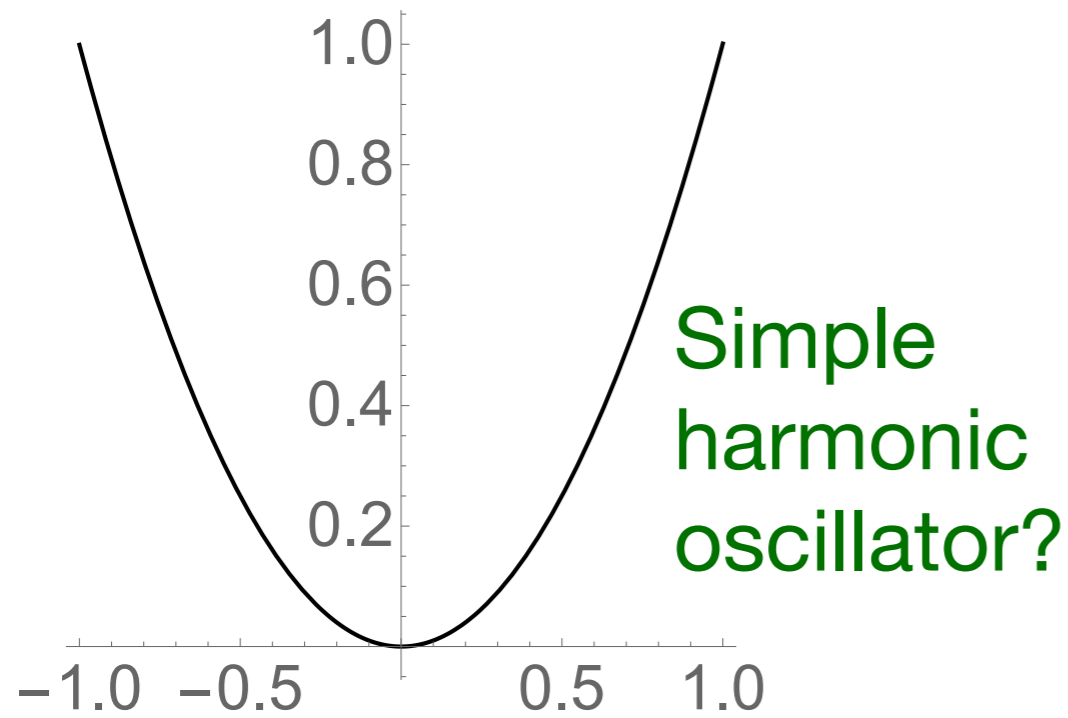
Standard Model gives the weak neutral couplings to high precision.

But is there a hint of “new physics” in the LEP1/SLC discrepancy?

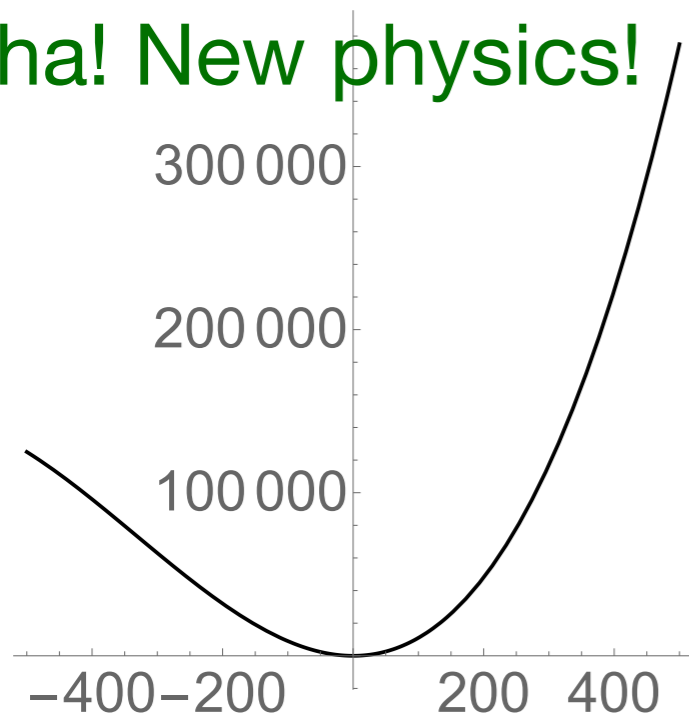


➡ Want to look for new bosons at high masses!

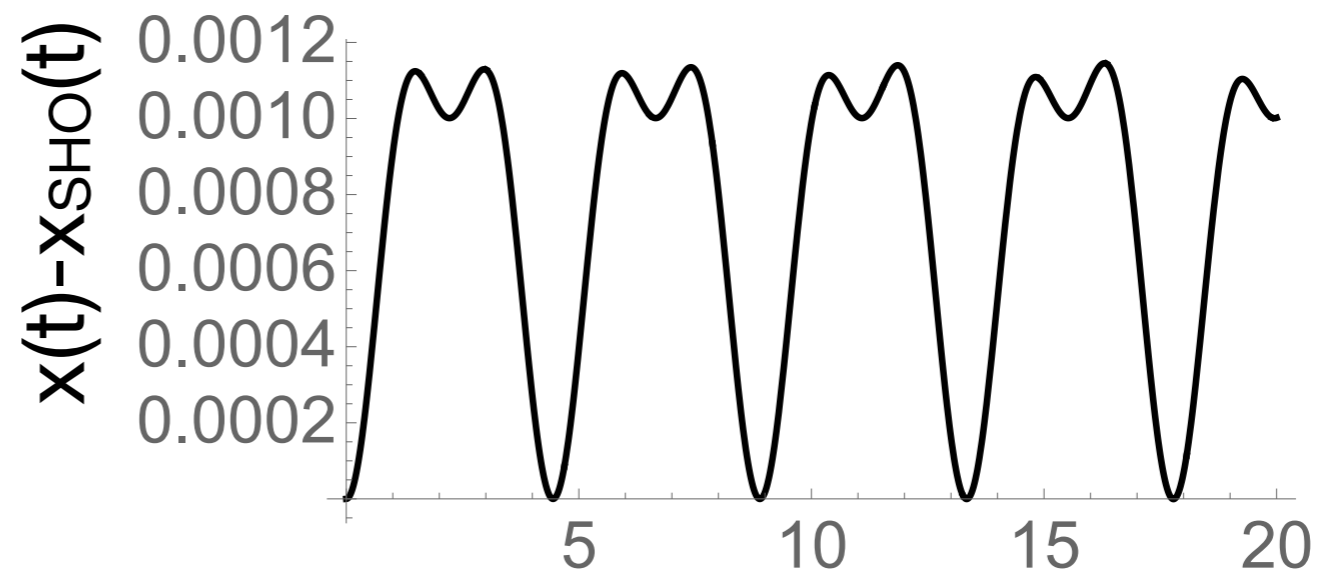
“High Energy Physics” at Low Energies



Aha! New physics!



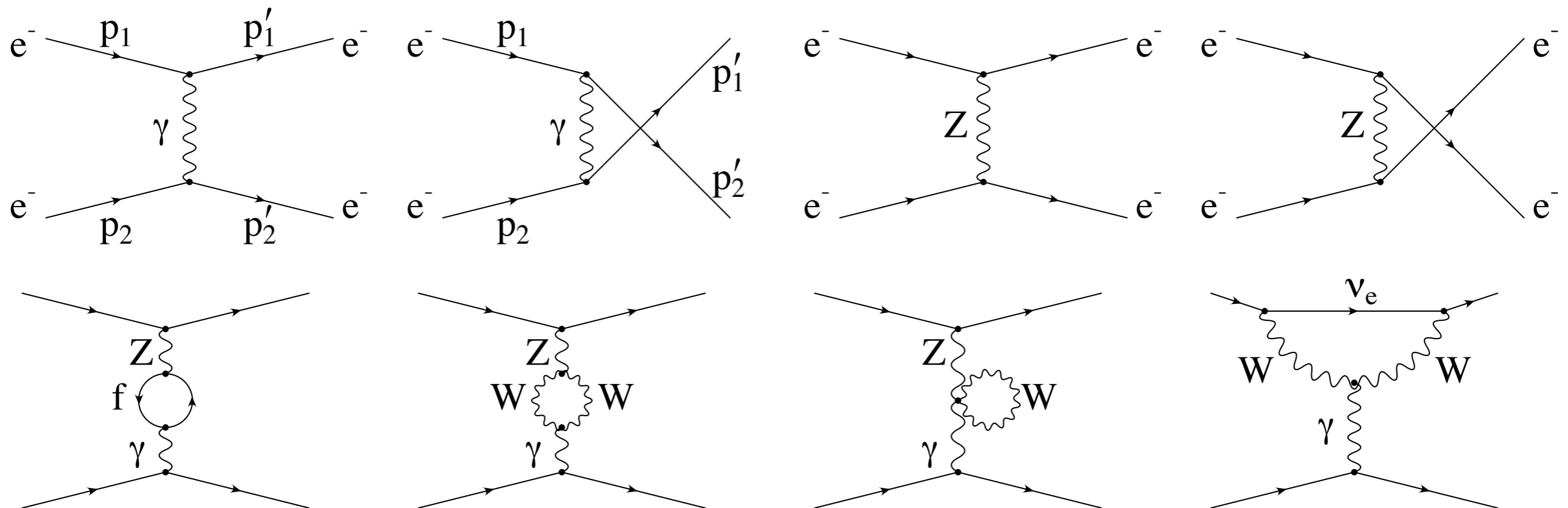
Unless you take a closer look!



MOLLER @ JLab

Measurement Of Lepton-Lepton Elastic Reaction

Parity Violation in Electron-Electron Elastic Scattering



➔ *Look for deviations from the calculated result!*

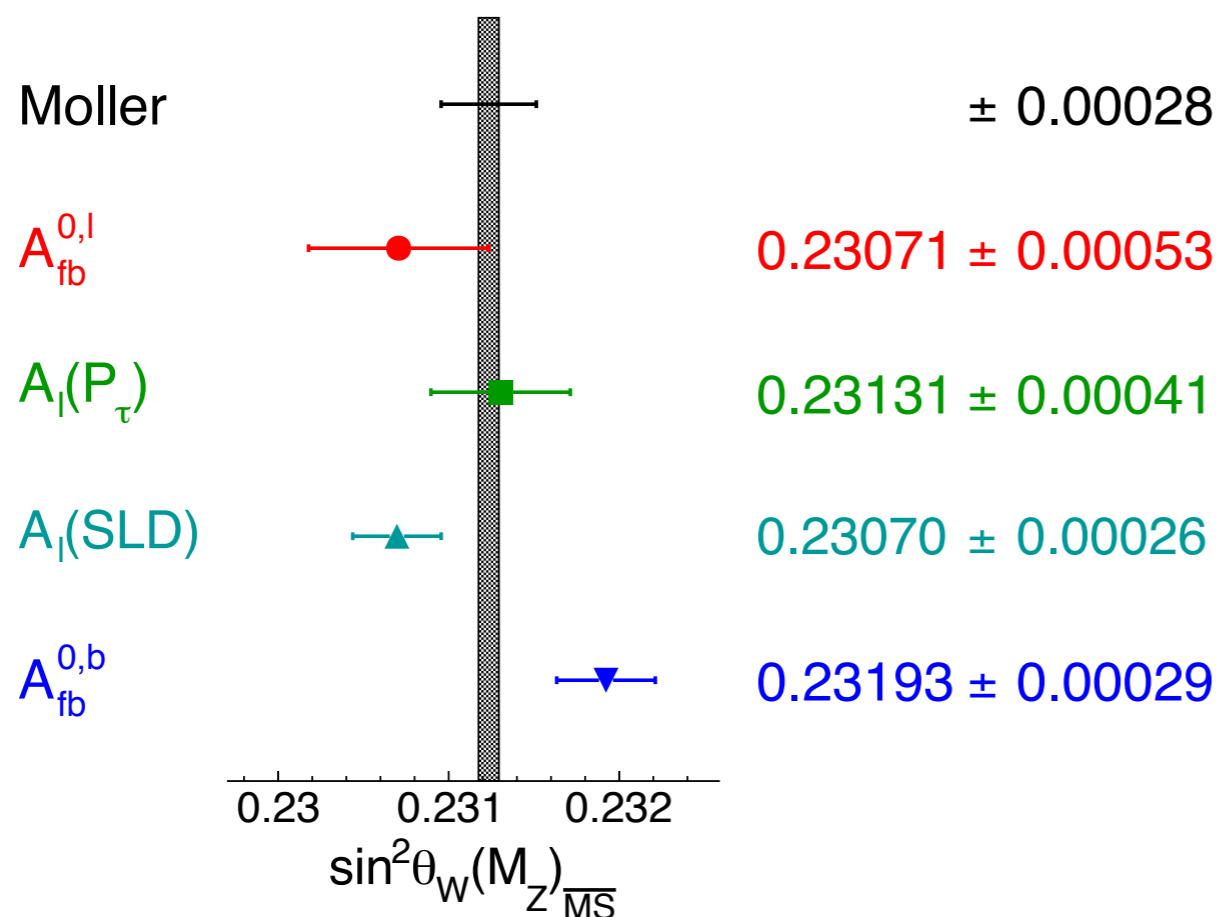
Asymmetry is 33 ppb. Precision goal is 2.4% (0.7 ppb)

Sensitivity to New Physics

Standard Model → *New Physics Added*

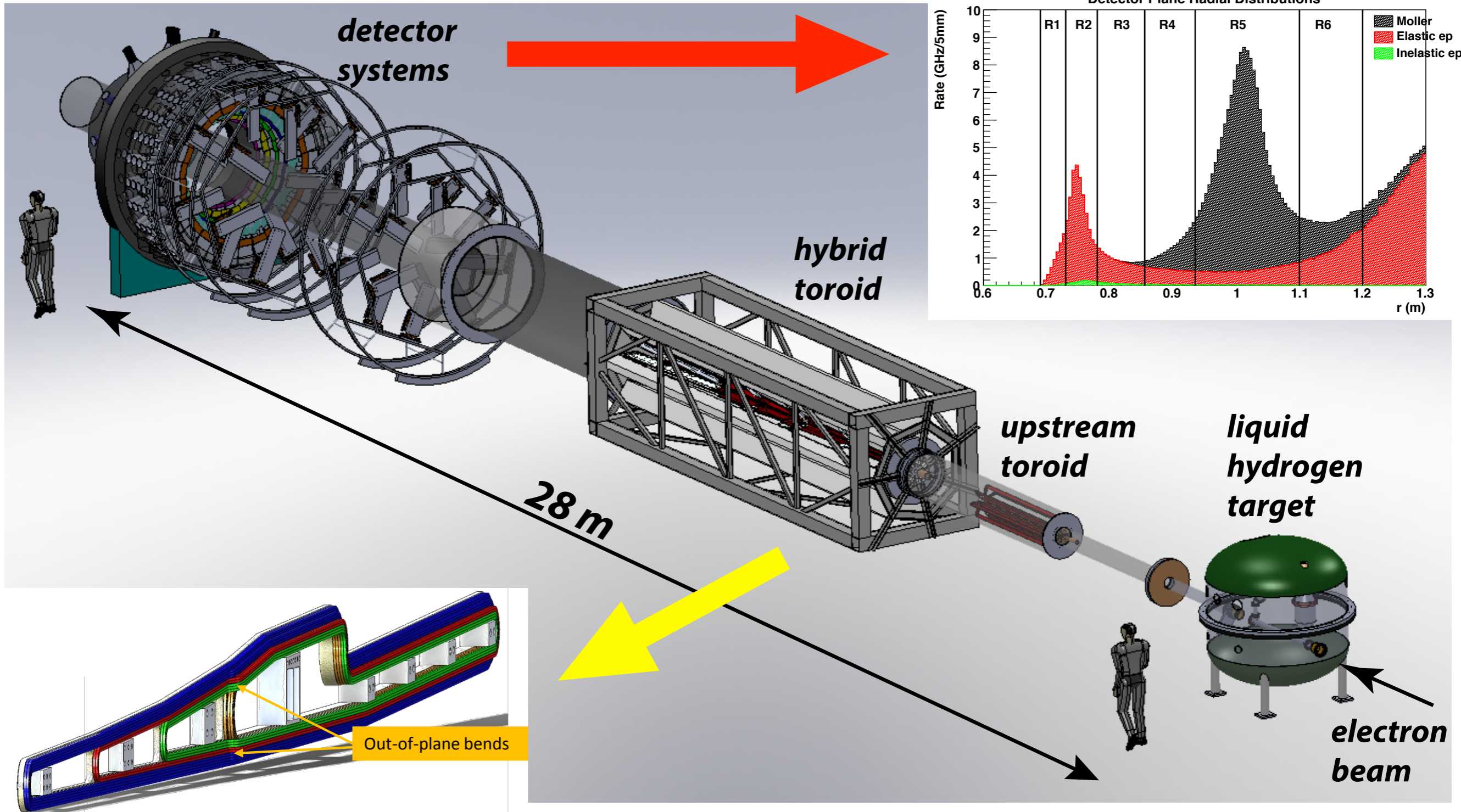
$$\frac{g_{AV}^{eq}}{2v^2} \bar{e} \gamma^\mu \gamma^5 e \bar{q} \gamma_\mu q \rightarrow \left[\frac{g_{AV}^{eq}}{2v^2} + \frac{4\pi}{(\Lambda_{AV}^{eq})^2} \right] \bar{e} \gamma^\mu \gamma^5 e \bar{q} \gamma_\mu q$$

Higgs
A New Scale!

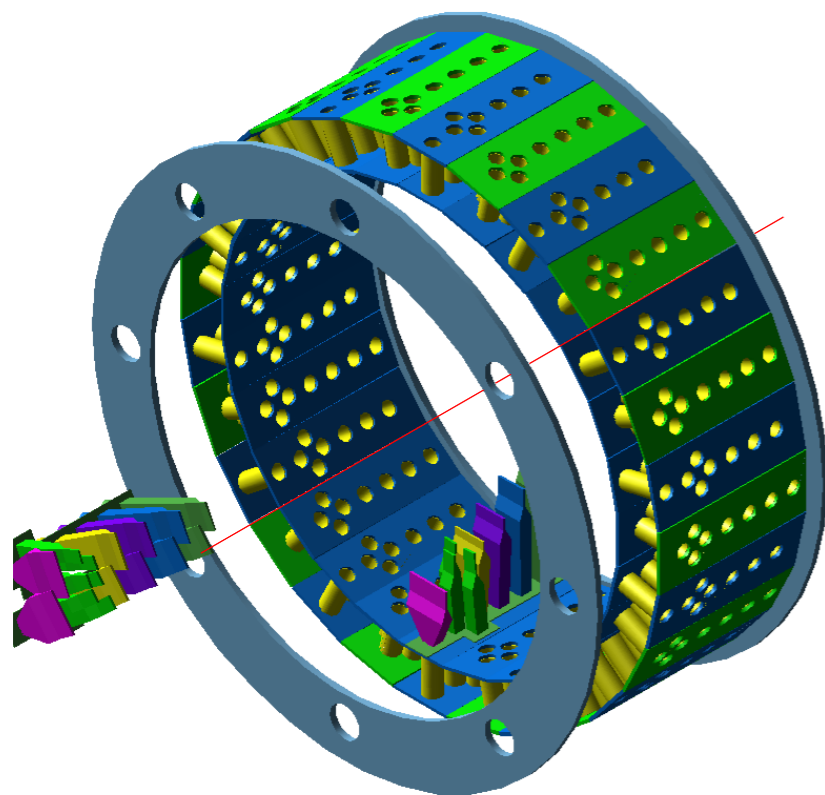
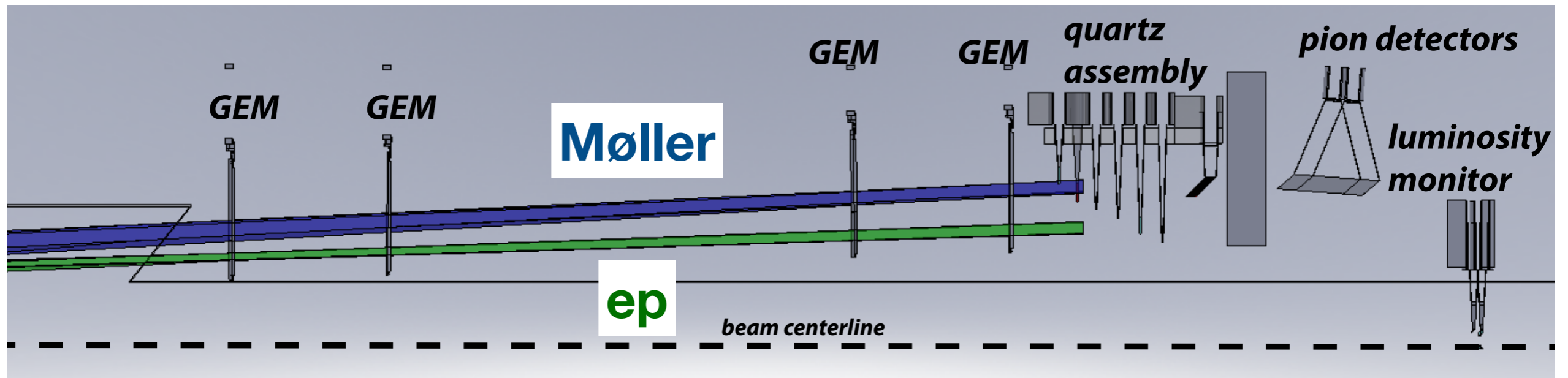


Depending on the new couplings, the MOLLER result will test for Higgs-like particles with mass up to ≈ 10 TeV (roughly 100×Higgs)

Experimental Setup

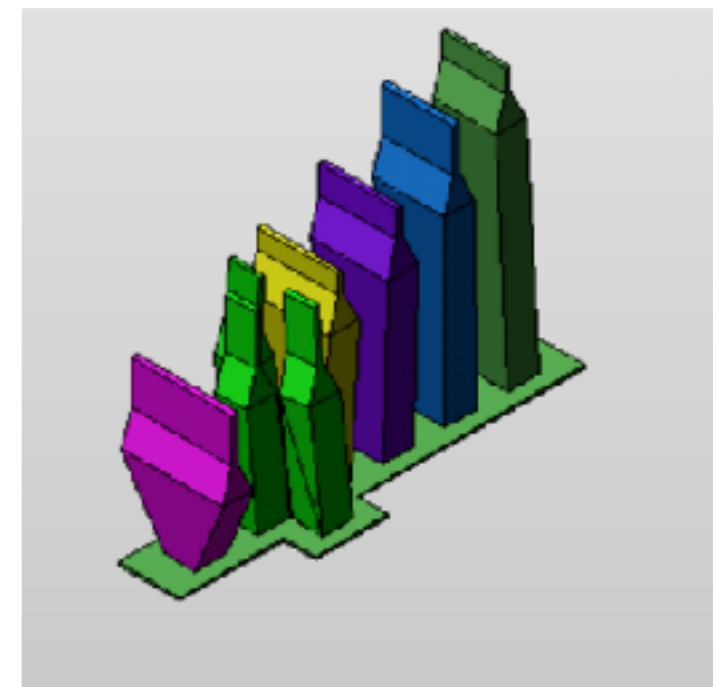


Detector Arrangement



Quartz Cherenkov detectors to gather light from scattered electrons.

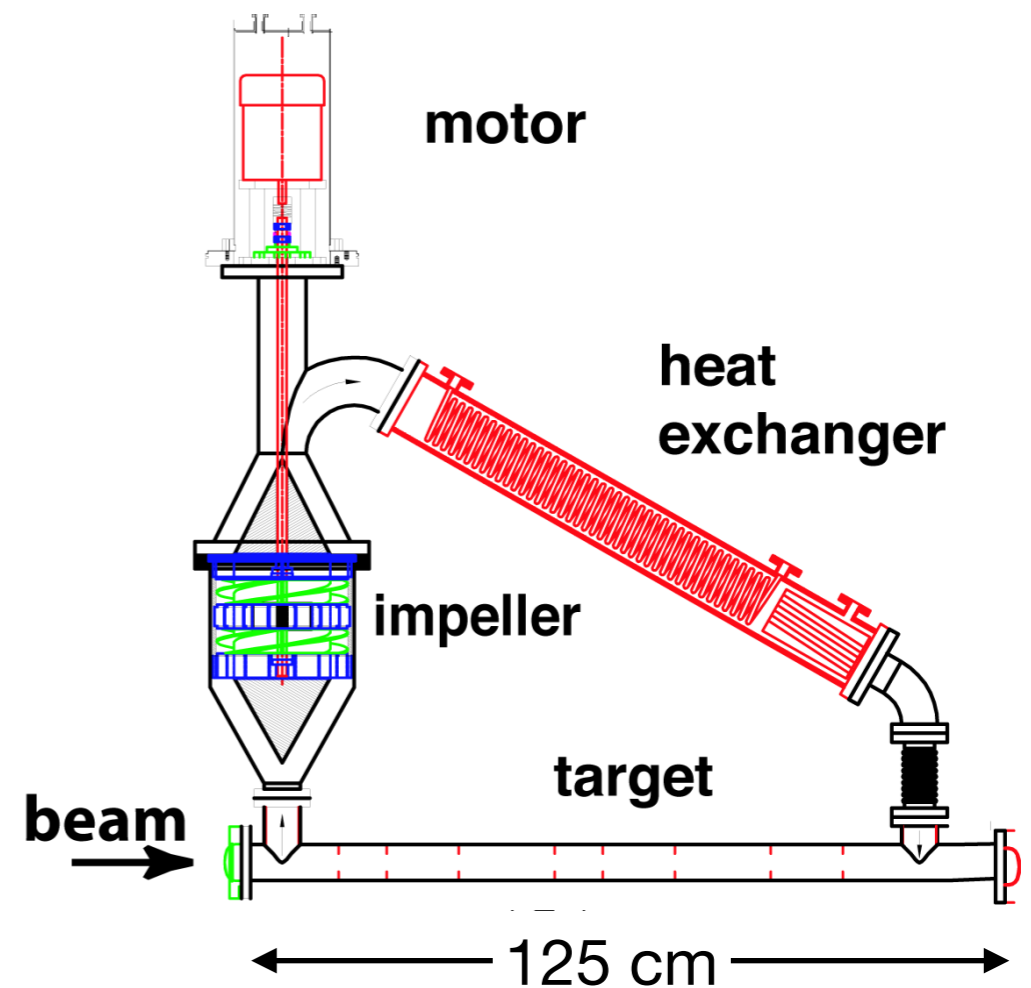
PMT signals are integrated for each polarization state.



Some Parameters

Beam E	11 GeV
Møller E'	1.7 - 8.5 GeV
θ_{CM}	$46^\circ - 127^\circ$
θ_{Lab}	$0.23^\circ - 1.1^\circ$
$\langle Q^2 \rangle$	0.0058 GeV^2
Current	$70 \mu\text{A}$
Møller rate	123 GHz
Flip rate	1920 Hz

Liquid Hydrogen Target



Deposited power = 4kW

Uncertainty Budget

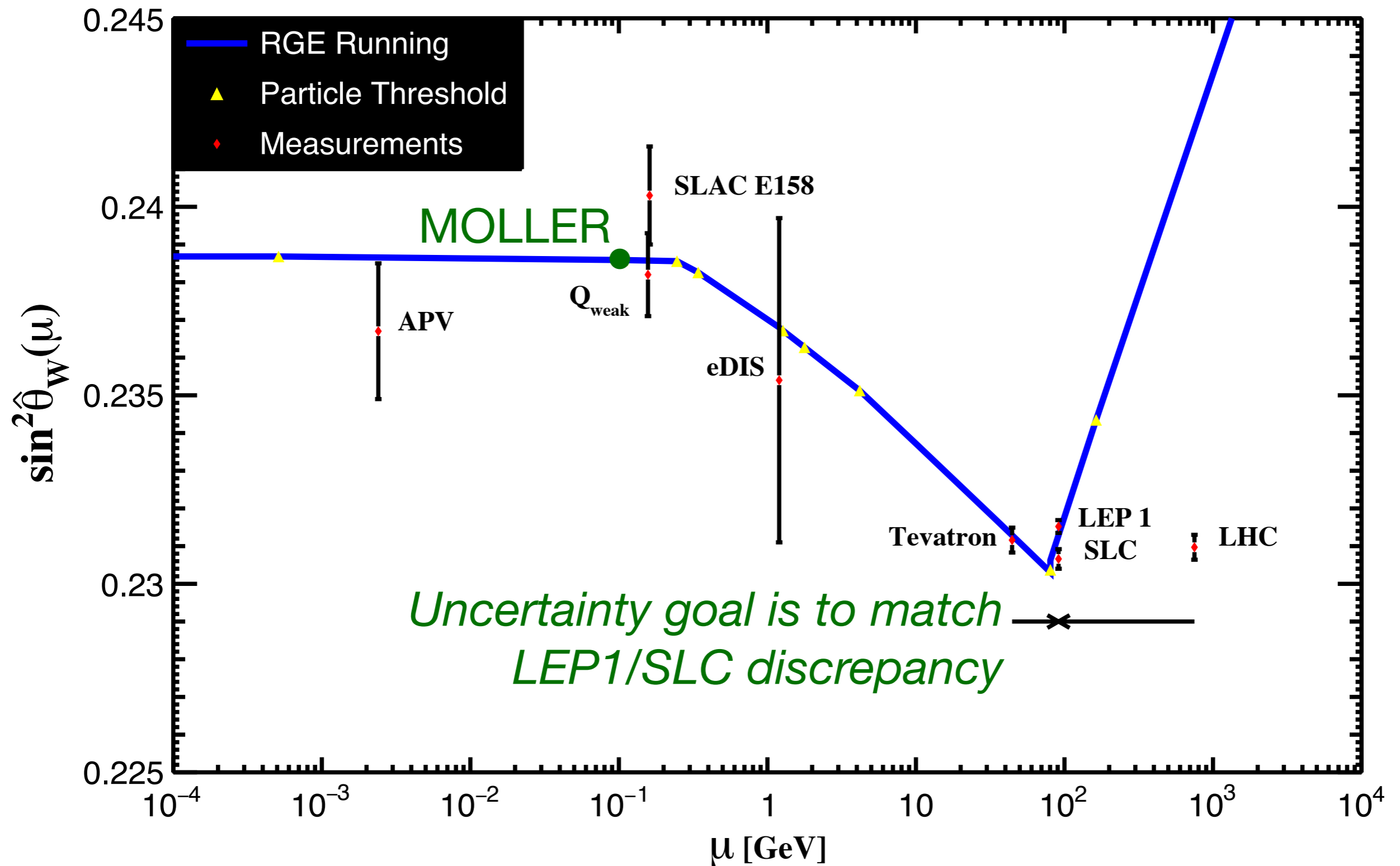
Challenging Experiment!

But the scientific case is solid, CD1 is granted, R&D to finalize design is underway, and we are on a path for installation some time in 2025.

Error Source	Fractional Error (%)
Statistical	2.1
Absolute value of Q^2	0.5
beam (second order)	0.4
beam polarization	0.4
$e + p(+\gamma) \rightarrow e + X(+\gamma)$	0.4
beam (position, angle, energy)	0.4
beam (intensity)	0.3
$e + p(+\gamma) \rightarrow e + p(+\gamma)$	0.3
$\gamma^{(*)} + p \rightarrow \pi + X$	0.3
Transverse polarization	0.2
neutrals (soft photons, neutrons)	0.1
Total systematic	1.1

Beam Property	Assumed Sensitivity	Accuracy of Correction	Required 1 kHz random fluctuations	Required cumulative helicity-correlation	Systematic contribution
Intensity	1 ppb / ppb	$\sim 1\%$	< 1000 ppm	< 10 ppb	~ 0.1 ppb
Energy	-1.4 ppb / ppb	$\sim 10\%$	< 286 ppm	< 0.7 ppb	~ 0.05 ppb
Position	0.85 ppb / nm	$\sim 10\%$	< 47 μm	< 1.2 nm	~ 0.05 ppb
Angle	8.5 ppb / nrad	$\sim 10\%$	< 4.7 μrad	< 0.12 nrad	~ 0.05 ppb

Comparison with Other Results



Conclusions

Parity Violating electron scattering with GeV beams is a useful tool for studying nature.

We now have new information that will help us understand the structure of neutron stars!

Next: MOLLER will search for New Physics at multi-TeV energies through Precision ePV.

Thank You!