

# A Framework for Modifying Quantum Mechanics

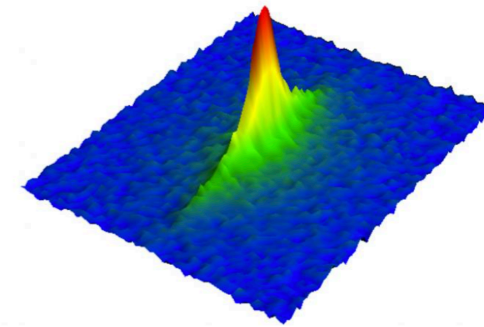
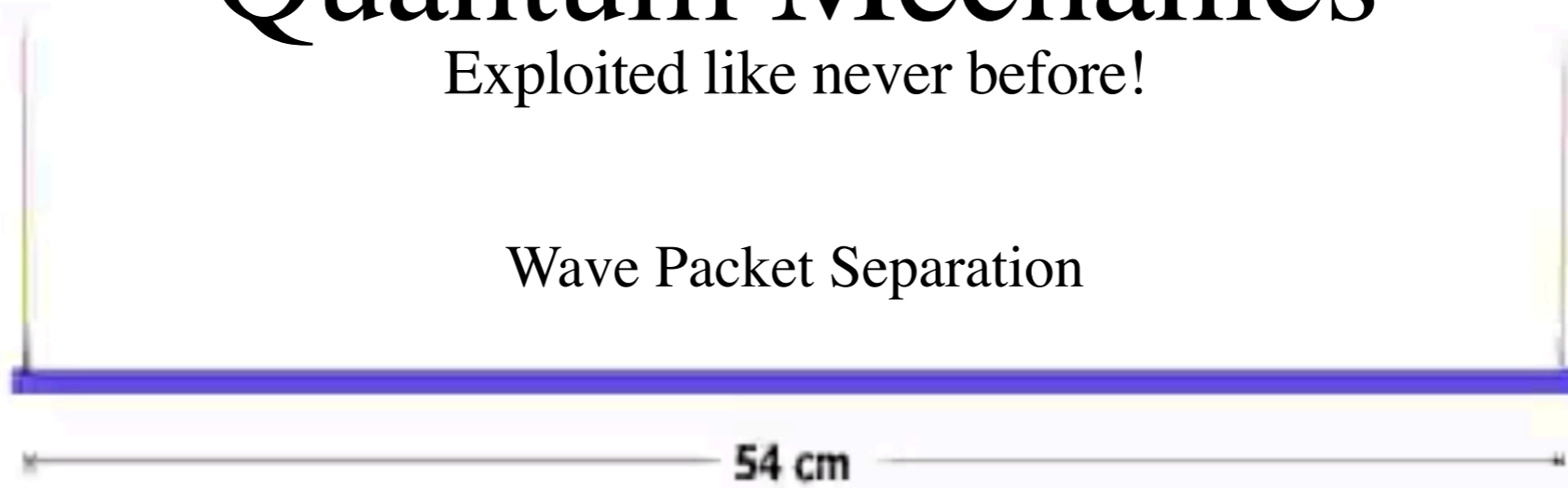
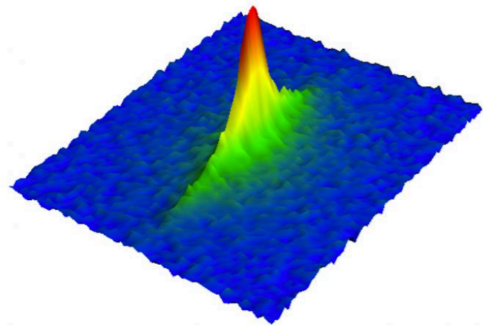
David E. Kaplan

# Quantum Mechanics

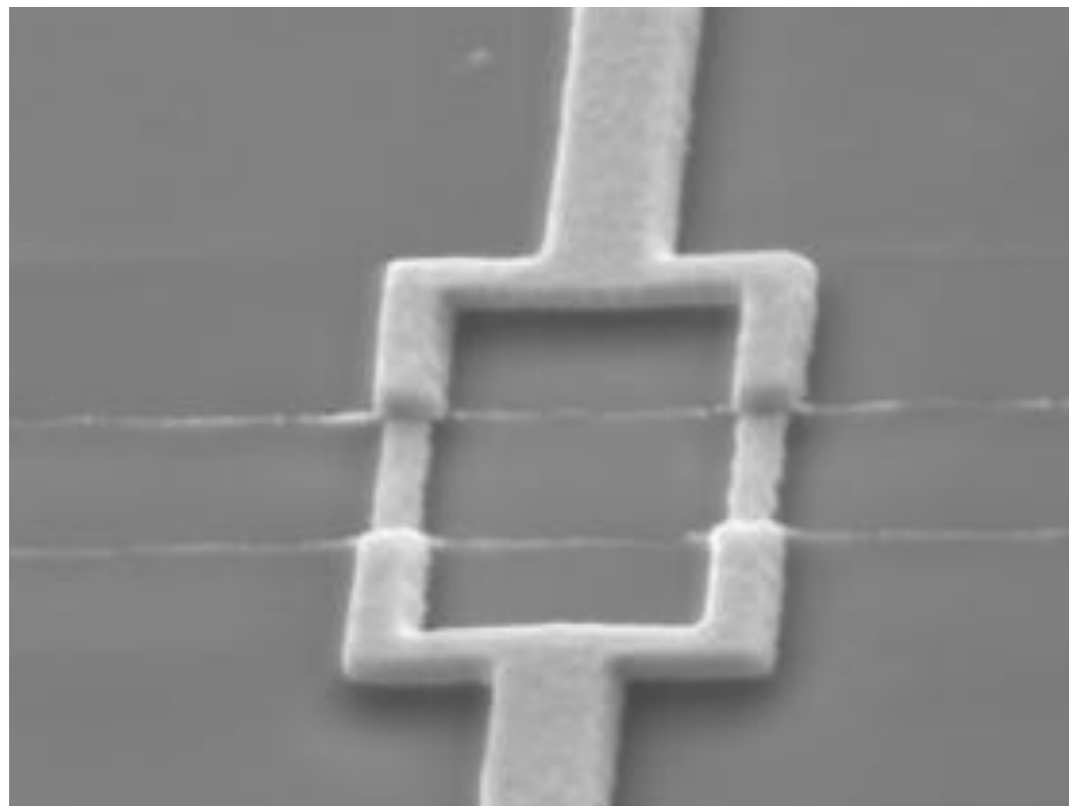
Exploited like never before!

Wave Packet Separation

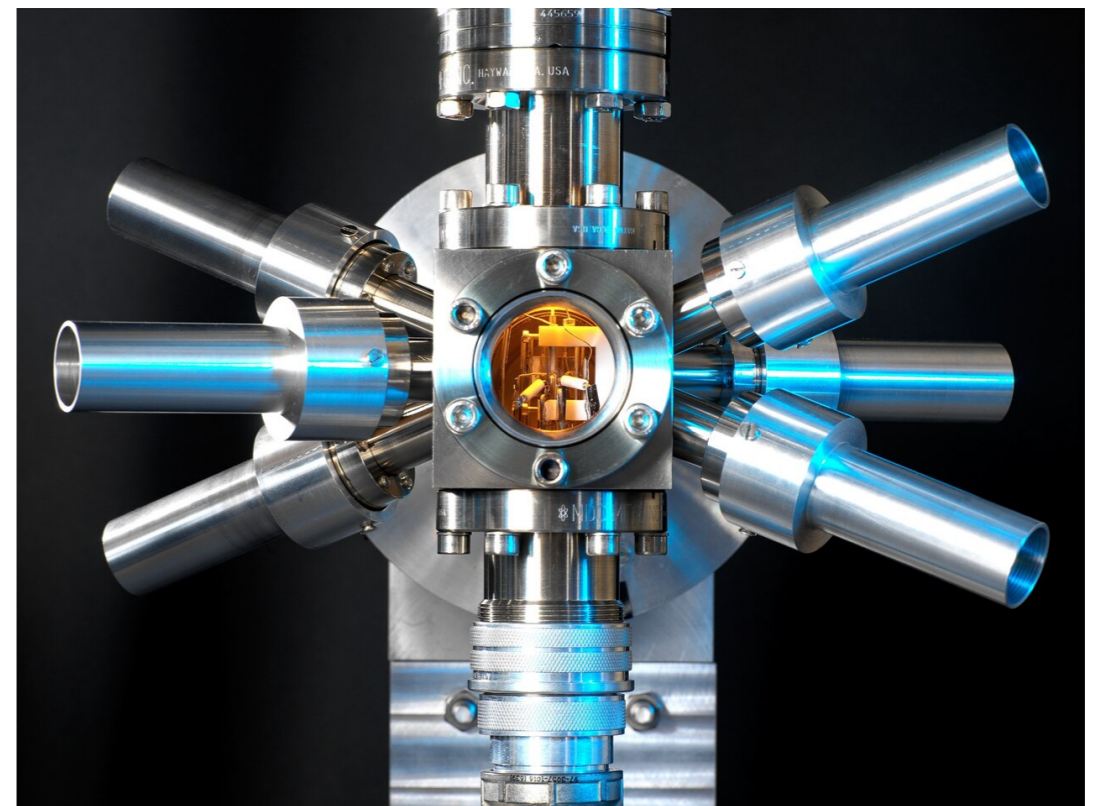
54 cm



SQUID



Optical Lattice

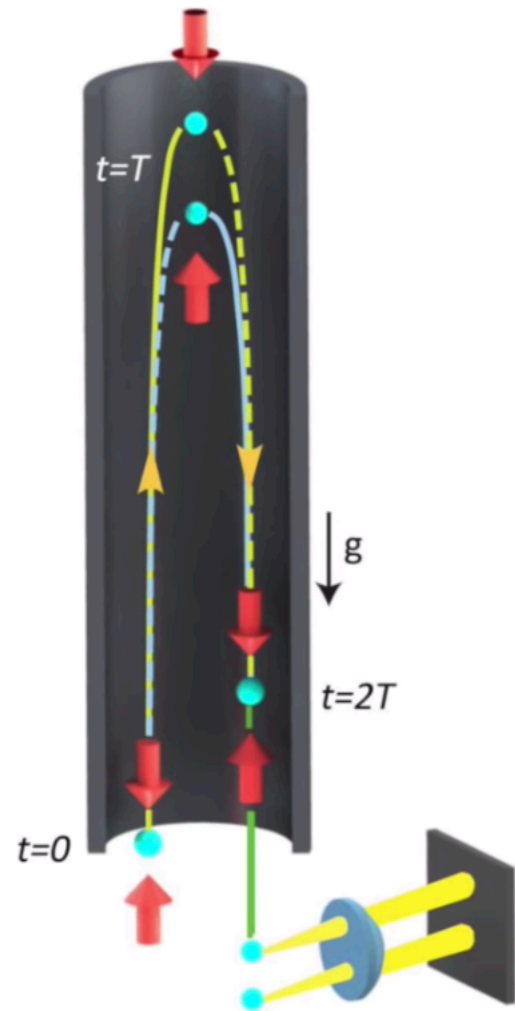


Tweezed Atoms

# Quantum Mechanics

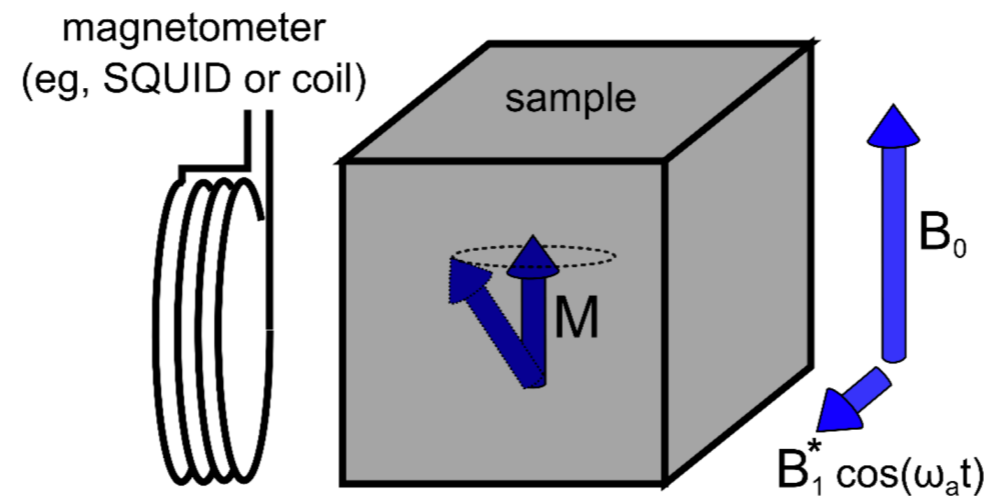
Use to probe fundamental physics!

Atom Interferometers



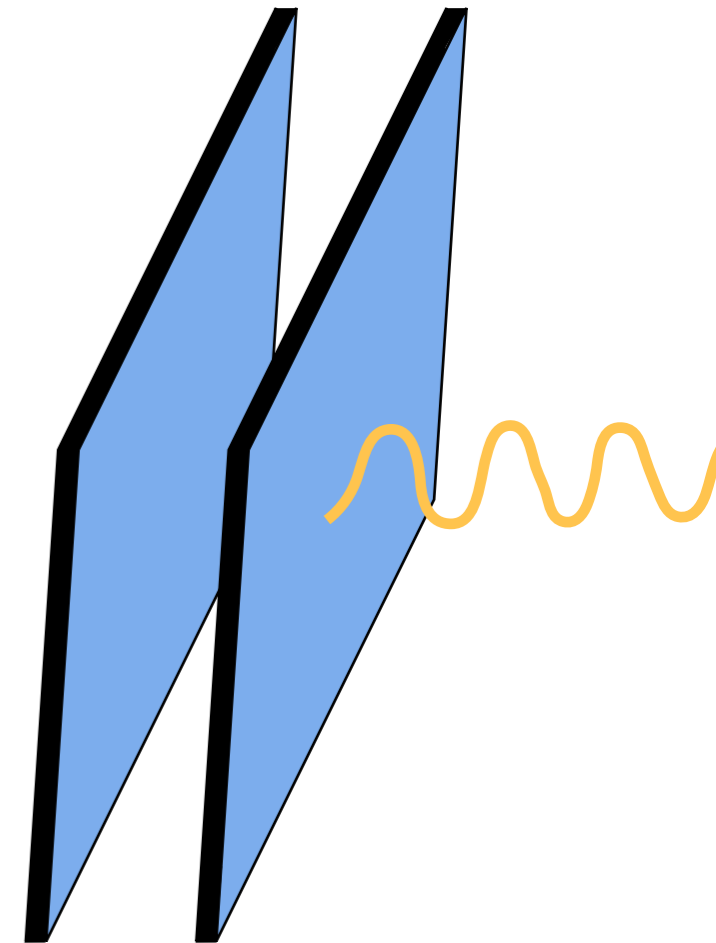
Gravitational Waves!

SQUIDS



Dark Matter!

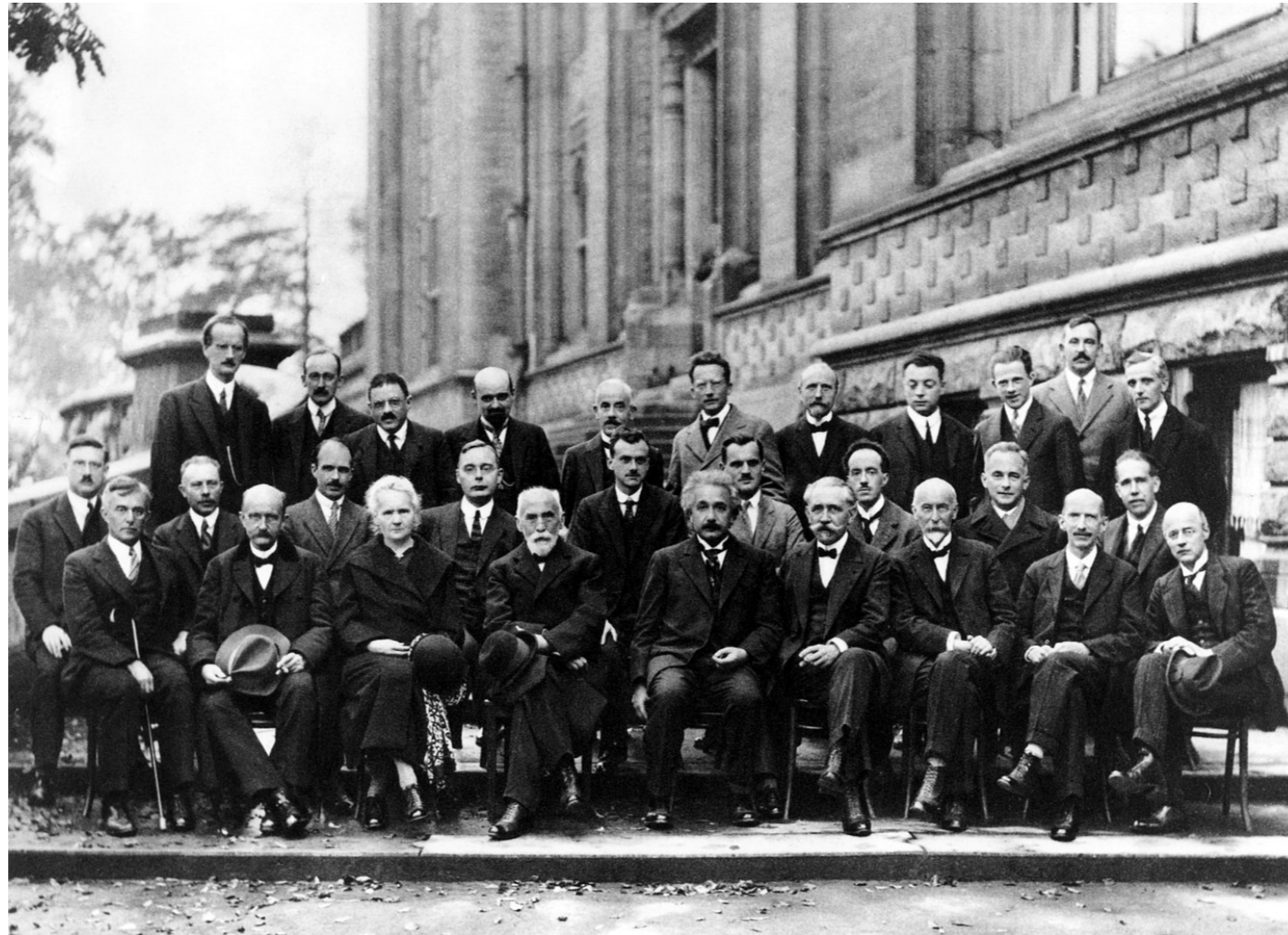
Mössbauer



New Forces!

**We can also test quantum mechanics itself!**

# Quantum Mechanics



1927

Can quantum mechanics be modified or generalized?

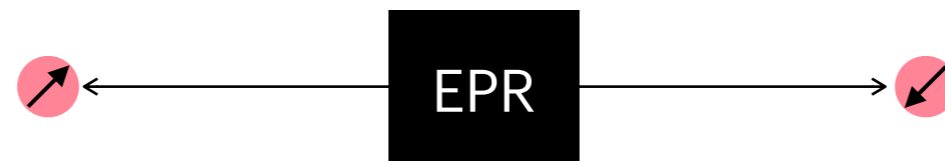
Are there sensible theories?

Are there hints of where it might deviate?

# What to Modify?

Can we get rid of unsightly ‘probability’?

Bell’s inequality, etc., rule out local hidden variables.



Ground state of Hydrogen:

If the electron has a definite position, there is an infinite degeneracy of states



# What to Modify?

Can we get rid of **linearity**?

$$i\hbar \frac{\partial}{\partial t} |\psi\rangle = \hat{H} |\psi\rangle$$

de Broglie (1960) suggested QM is a linear approximation

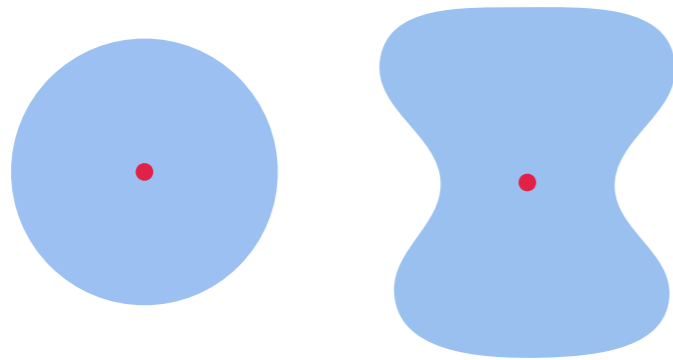
Weinberg (1989) suggested a form

$$\hat{H} \rightarrow \hat{H} [ |\psi\rangle, \langle\psi| ]$$

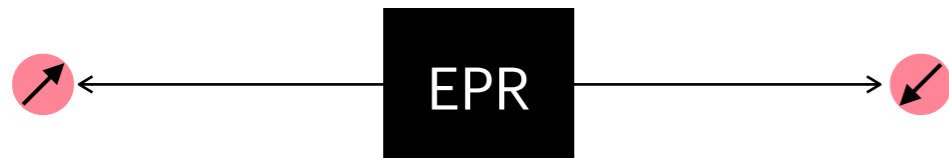
# What to Modify?

Can we get rid of **linearity**?

$$\hat{H} \rightarrow \hat{H} [ |\psi\rangle, \langle\psi| ]$$



Wave-function self-interaction will move energy levels: check sensitive measurements in QM.



Must check causality and understand measurement.

Can this be embedded in quantum field theory?

# Framework



# Non-Linear Time Evolution

The Schrödinger Equation: position basis

$$i\hbar \frac{\partial}{\partial t} \psi(x) = H(\mathbf{x}) \psi(x)$$

Weinberg's attempt (1989)

$$i\hbar \frac{\partial}{\partial t} \psi(x) = h(\psi^*, \psi) \psi(x)$$

Polchinski showed action at a distance with EPR pairs (1990)

# Non-Linear Quantum Mechanics

(warning...  $c = 1$  and  $\hbar = 1$ )

$$i \frac{\partial}{\partial t} \psi(x) = \hat{H}(\mathbf{x}) \psi(x) + \epsilon \int d^4 x' |\psi(x')|^2 G_R(x' - x) \psi(x)$$

Causality guaranteed by the retarded Green's Function (e.g., massless):

$$\square G_R(y; x) = \delta^4(x - y)$$

$$G_R(x - x') = \frac{\delta(t' - (t + |\mathbf{x}' - \mathbf{x}|))}{|\mathbf{x}' - \mathbf{x}|}$$

# Non-Linear Quantum Mechanics

One-particle, non-relativistic limit - the non-linear Schrödinger Equation

$$i \frac{\partial}{\partial t} \psi(x) = \hat{H}(\mathbf{x}) \psi(x) + \epsilon \underbrace{\int d^4 x' |\psi(x')|^2 G_R(x' - x)}_{\delta H(x)} \psi(x)$$

This term is Hermitian, thus  
the norm conserved

New stationary states can be found perturbatively for a fixed potential

# Entangled Systems and Causality

Wave function for two entangled particles:

$$\psi(\mathbf{x}, \mathbf{y}; t) = \sum_{i,j} c_{ij}(t) \psi_i^I(\mathbf{x}) \psi_j^{II}(\mathbf{y})$$

$$\delta \hat{H} \psi = \int d^3x' d^3y' dt' |\psi(\mathbf{x}', \mathbf{y}'; t')|^2 \left( \boxed{G_R(\mathbf{x}, t; \mathbf{x}', t')} + \underbrace{G_R(\mathbf{y}, t; \mathbf{x}', t') + G_R(\mathbf{x}, t; \mathbf{y}', t')}_{\text{causal}} + \boxed{G_R(\mathbf{y}, t; \mathbf{y}', t')} \right) \psi(\mathbf{x}, \mathbf{y}, t)$$

$$\delta \hat{H} \subset \int \underbrace{d^3x' d^3y' dt' |\psi(\mathbf{x}', \mathbf{y}'; t')|^2}_{\text{causal}} G_R(\mathbf{y}, t; \mathbf{y}', t')$$

after measurement at  $\mathbf{x}$ , this integral unchanged

# Quantum Field Theory

$$i \partial_t |\chi\rangle = \hat{H} |\chi\rangle$$

In the Schrödinger picture, the time evolution operator is still:

$$\hat{U} = e^{-i\hat{H}t}$$

with  $\hat{H} = \int d^3x \hat{\mathcal{H}}(\mathbf{x})$  made up of field operators

Add state-dependent terms.

# Quantum Field Theory

‘Non-linear’  $\rightarrow$  state-dependent

$$i \partial_t |\chi\rangle = \left[ \int d^3x \hat{\mathcal{H}}(\mathbf{x}) + \epsilon \left( \langle \chi | \hat{\mathcal{O}}_1(\mathbf{x}) | \chi \rangle \hat{\mathcal{O}}_2(\mathbf{x}) + \hat{\mathcal{O}}_1(\mathbf{x}) \langle \chi | \hat{\mathcal{O}}_2(\mathbf{x}) | \chi \rangle \right) \right] |\chi\rangle$$

Time evolution includes terms that depends on the state itself

If  $\mathcal{O}_1$  and  $\mathcal{O}_2$  are Hermitian, the norm is constant

$$\partial_t \langle \chi | \chi \rangle = 0$$

Probabilistic interpretation of observables can be maintained

# QFT Examples

YUKAWA THEORY

$$\mathcal{H} \supset y \phi(\mathbf{x}) \bar{\Psi}(\mathbf{x}) \Psi(\mathbf{x})$$

Add non-linearity

$$\langle \phi \rangle \equiv \langle \chi | \phi | \chi \rangle$$

$$\mathcal{H} \supset y (\phi(\mathbf{x}) + \epsilon \langle \phi(\mathbf{x}) \rangle) \bar{\Psi}(\mathbf{x}) \Psi(\mathbf{x})$$

Assuming  $\langle \chi | \chi \rangle = 1$

Perturbatively, compute background source in the  $\epsilon = 0$  theory

# QFT Examples

QED

$$\mathcal{L} \supset e A_\mu J^\mu$$

Add non-linearity

$$\mathcal{L} \supset e \left( \frac{A_\mu + \epsilon_\gamma \langle A_\mu \rangle}{1 + \epsilon_\gamma} \right) J^\mu$$

$A_\mu$  and  $\langle A_\mu \rangle$  have the same gauge transformations —  
gauge fix and generate the Hamiltonian

GRAVITY

Replace  $g_{\mu\nu} \rightarrow \frac{g_{\mu\nu} + \epsilon_G \langle g_{\mu\nu} \rangle}{1 + \epsilon_G}$  in interaction terms. Remains a tensor.



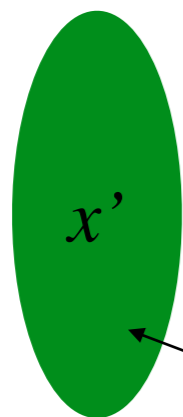
# Non-Relativistic Limit — One Particle

$$\mathcal{L} \supset y(\phi + \epsilon \langle \chi | \phi | \chi \rangle) \bar{\Psi} \Psi$$

To get NR theory for fermions  $\Psi$ , compute  $\langle \phi \rangle$ .

Will depend on initial conditions and sources. At zeroth order,  $\Psi$  sources  $\phi$ :

$\psi$  wave function for single fermion  $\Psi$



$$\langle \phi \rangle(x) \supset \int d^4x' |\psi(x')|^2 G_R(x - x')$$

Charge density of  $\Psi$

Causal Green's Function

# **The Classical Limit**

# Classical Physics from QM

$$i\hbar \frac{\partial}{\partial t} \langle \hat{O} \rangle = \left\langle \left[ \hat{H}, \hat{O} \right] \right\rangle$$

This leads to  $\frac{\partial \langle p \rangle}{\partial t} = - \left\langle \frac{\partial V}{\partial x} \right\rangle$  and  $\frac{\partial \langle x \rangle}{\partial t} = \frac{\langle p \rangle}{m}$

Or,  $F = ma$  on average

Coherent states (or classical-like states) are

ones in which, *e.g.*,  $\left\langle \frac{\partial V(x)}{\partial x} \right\rangle \simeq \frac{\partial V(\langle x \rangle)}{\partial \langle x \rangle}$

# Classical Physics from NLQM

$$\text{Say } \mathcal{H} \supset \hat{A}_\mu \hat{J}^\mu + \epsilon_\gamma \left( \langle \hat{A}_\mu \rangle \hat{J}^\mu + \hat{A}_\mu \langle \hat{J}^\mu \rangle \right)$$

$$\begin{aligned} i\hbar \frac{\partial}{\partial t} \langle \hat{\mathcal{O}} \rangle &= \int \langle [\hat{\mathcal{H}}, \hat{\mathcal{O}}] \rangle \supset \int \langle [\hat{A} \cdot \hat{J}, \hat{\mathcal{O}}] \rangle + \epsilon_\gamma \langle [\langle \hat{A} \rangle \cdot \hat{J}, \hat{\mathcal{O}}] \rangle + \epsilon_\gamma \langle [\hat{A} \cdot \langle \hat{J} \rangle, \hat{\mathcal{O}}] \rangle \\ &= \int \left( \langle \hat{A} \cdot [\hat{J}, \hat{\mathcal{O}}] \rangle + \epsilon_\gamma \langle \hat{A} \rangle \cdot \langle [\hat{J}, \hat{\mathcal{O}}] \rangle \right) + \left( \langle [\hat{A}, \hat{\mathcal{O}}] \cdot \hat{J} \rangle + \epsilon_\gamma \langle [\hat{A}, \hat{\mathcal{O}}] \rangle \cdot \langle \hat{J} \rangle \right) \end{aligned}$$

The point is,  $\langle \hat{A}_\mu \hat{J}^\mu \rangle \simeq \langle \hat{A}_\mu \rangle \langle \hat{J}^\mu \rangle$  for classical states

In fact, the non-linear terms make physics *more classical!*

# Constraints from Quantum Systems

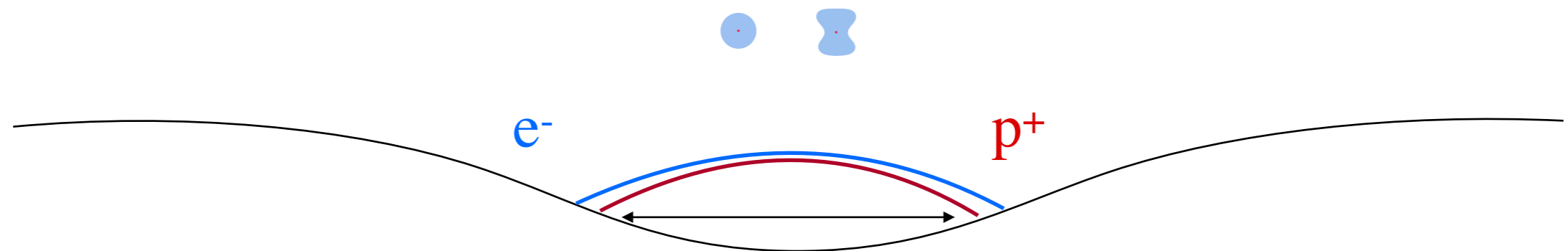
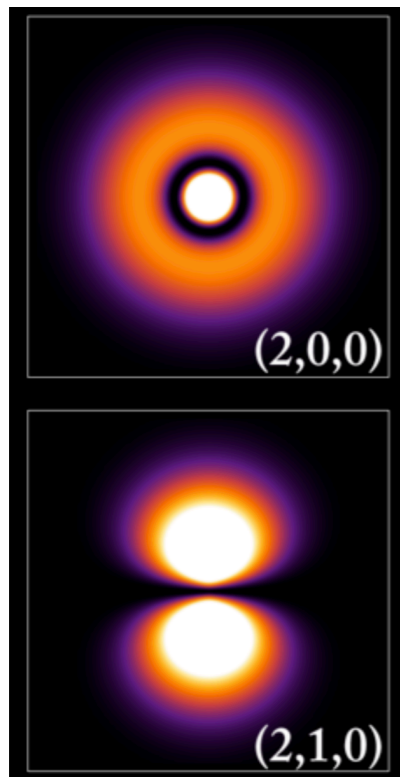
Atomic levels — lamb shift, (g-2) of the electron, ...

# Atomic Levels

What does this do to the Lamb Shift?

Say charged particles see their own w.f.:

$$i \frac{\partial}{\partial t} \psi(x) = H(\mathbf{x}) \psi(x) + \epsilon_\gamma \alpha \int d^4 x' |\psi(x')|^2 G_R(x' - x) \psi(x)$$



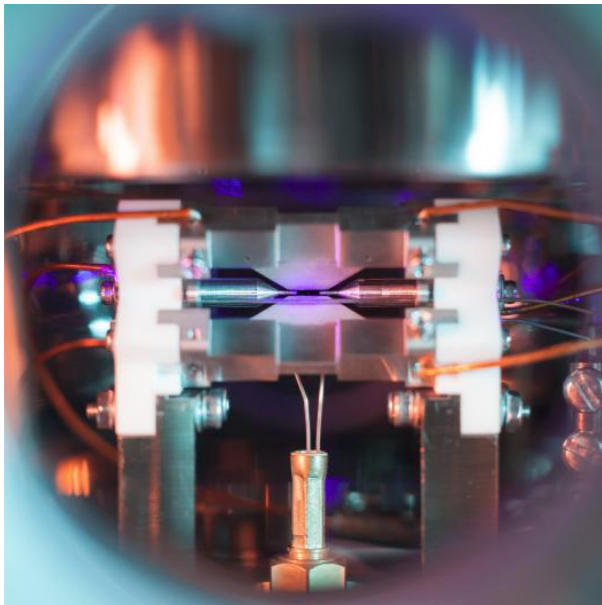
Electron spread over the trap (micron) dilutes the electric field and thus the level splitting  
Proton's wave function also produces a field that nearly cancels the electron wave function.

Key — center of mass coordinate cannot be separated from relative coordinate due to locality.

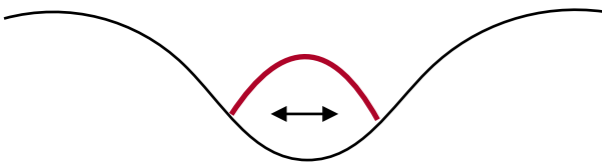
$$\epsilon_\gamma < 10^{-2}$$

# Constraints

## Leading Constraint



Ion Traps



For  $\varepsilon_\gamma > 0$  (repulsive interaction)

Too large a repulsion, can't trap ion  $\varepsilon_\gamma < 10^{-5}$

No direct limit on  $\varepsilon_\gamma < 0$  (attractive interaction)

Perhaps from mapping of ion in trap?

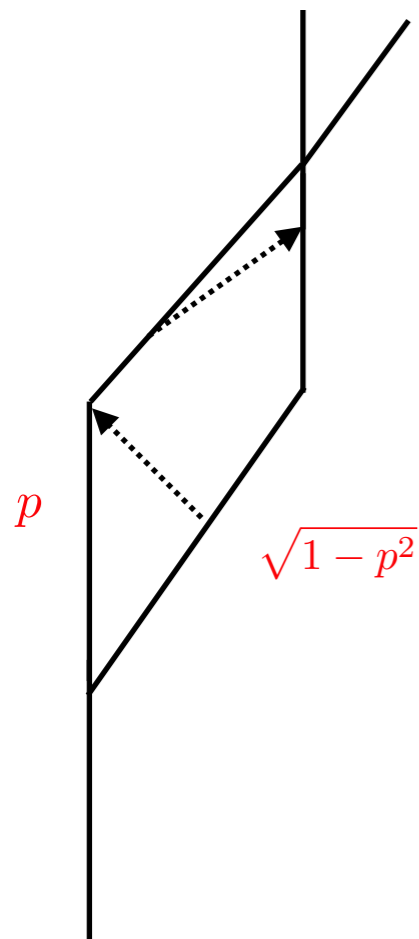
# Experimental Tests

Interferometry - interaction between paths

Take an ion - split its wave-function

Coulomb Field of one path interacts with the other path

Gives rise to phase shift that depends on the intensity  $p^2$  of the split



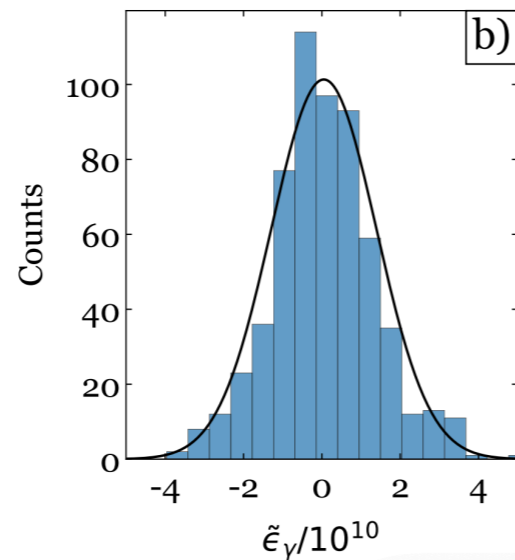
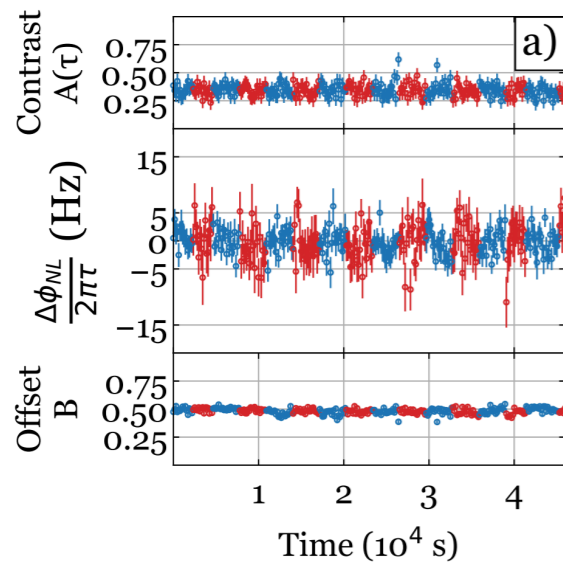
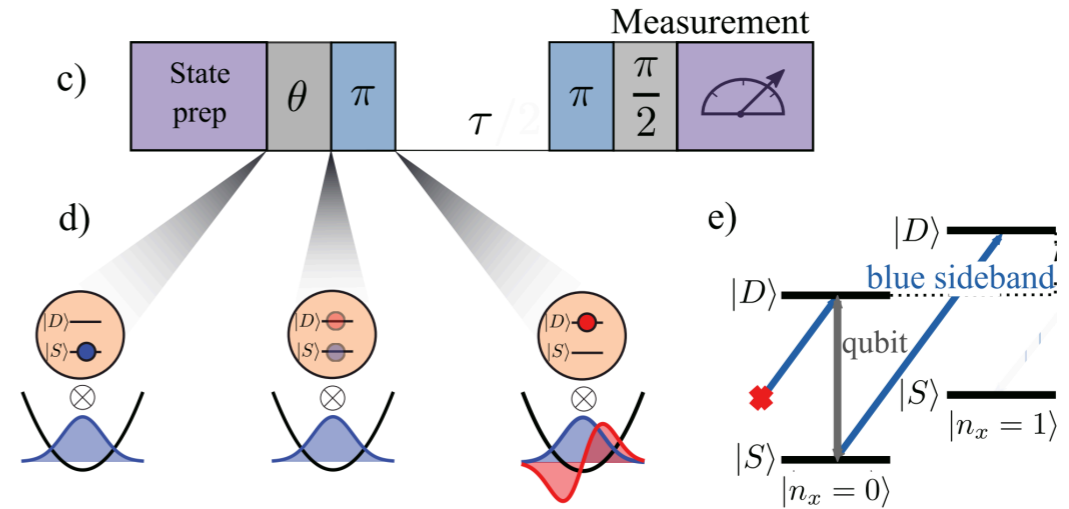
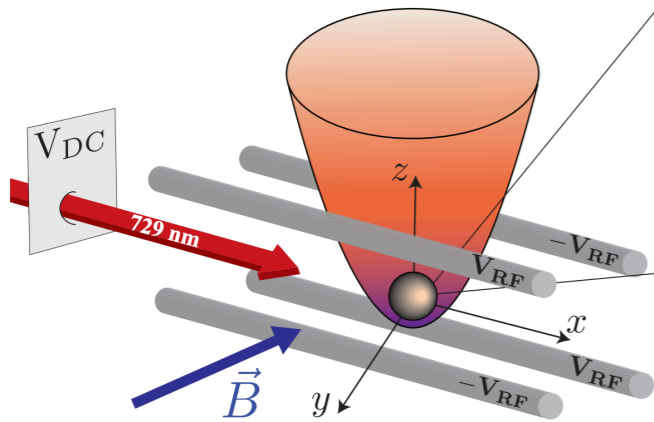
Use intensity dependence to combat systematics



# Test ... Using a Vibrational Mode of a Trapped Ion

J. Broz, et al. (2022)

State:  $|\psi\rangle = a|0\rangle + b|1\rangle$

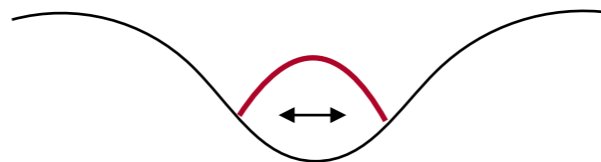


$$\text{Non-linear phase: } \phi_{NL} = \epsilon_{\gamma} \alpha \frac{10a^2 + b^2}{30\sqrt{2\pi\hbar}x_0} \tau$$

$$\epsilon_{\gamma} = 5 \pm 5.4 \times 10^{-12}$$

# Experimental Tests

## Atomic Aging



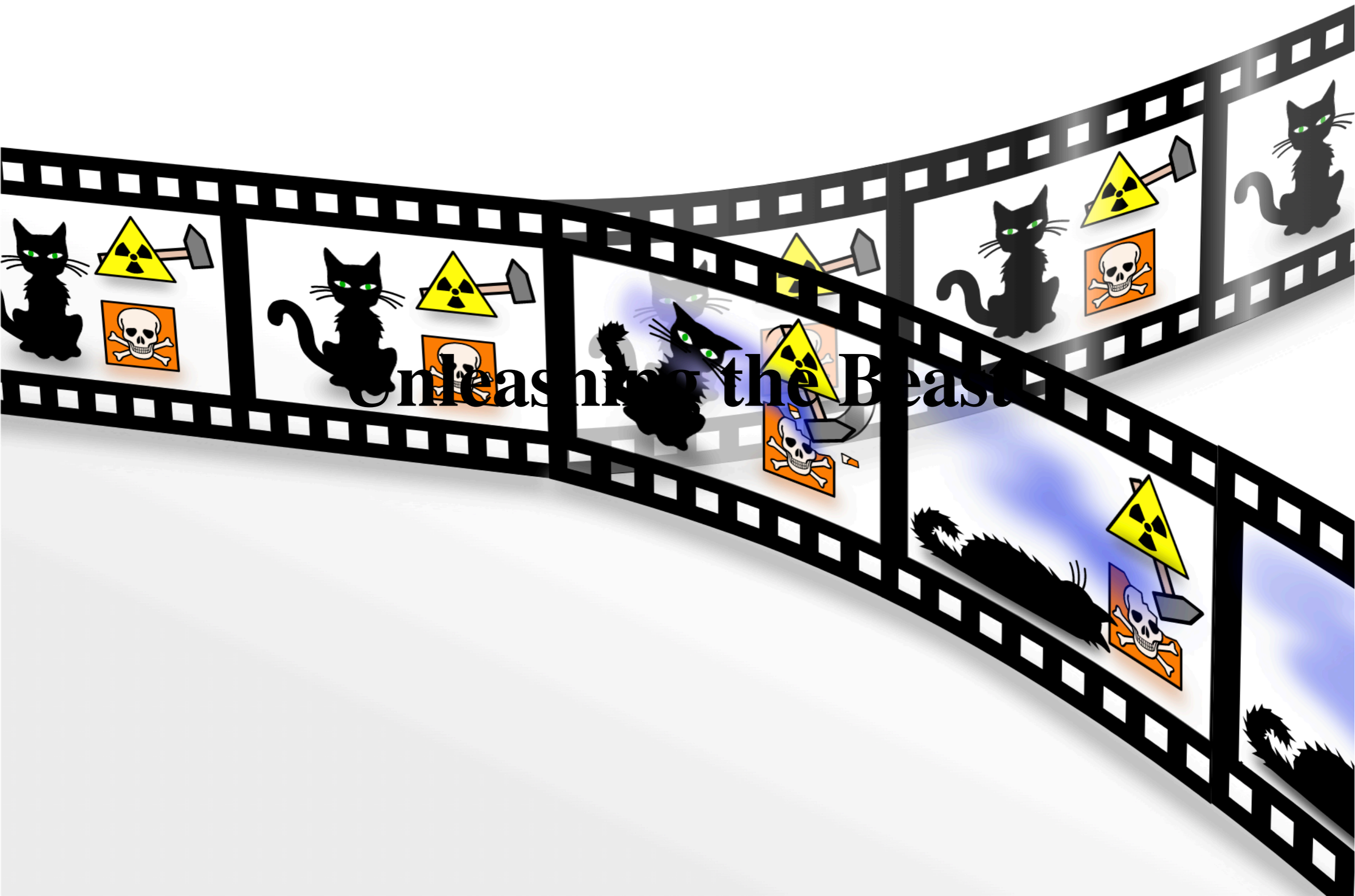
Ion Trap with decaying nucleus — wave function evolves over time:

$$|\psi\rangle = e^{-\Gamma t/2} |\text{Trapped}\rangle + \sqrt{1 - e^{-\Gamma t}} |\text{Decayed}\rangle$$



Background wave function produces E-field in trap and second order Stark shift in the atomic states.  $\delta\phi \sim E_{\text{NL}}^2 \Delta\alpha_p T$

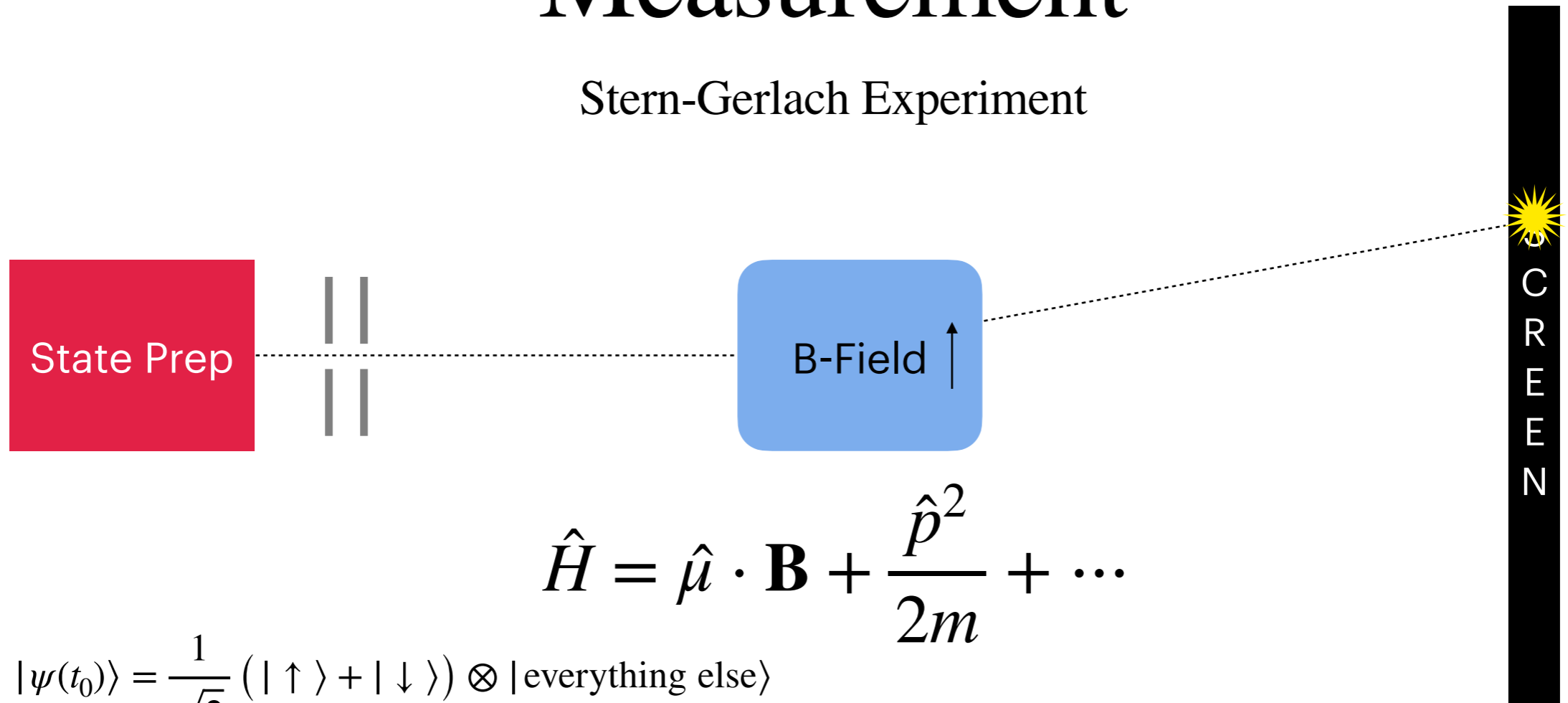
(not ideal as background suppression suppresses this effect)



Releasing the Beast

# Measurement

## Stern-Gerlach Experiment



$$\hat{H} = \hat{\mu} \cdot \mathbf{B} + \frac{\hat{p}^2}{2m} + \dots$$

$$|\psi(t_0)\rangle = \frac{1}{\sqrt{2}} (|\uparrow\rangle + |\downarrow\rangle) \otimes |\text{everything else}\rangle$$

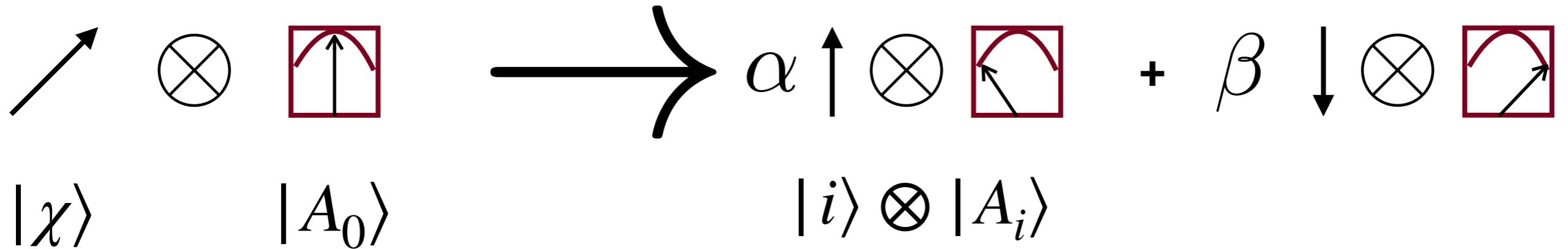
$$|\psi(t_1)\rangle = \frac{1}{\sqrt{2}} (|\uparrow\rangle|+p\rangle + |\downarrow\rangle|-p\rangle) \otimes |\text{everything else}\rangle$$

$$|\psi(t_2)\rangle = \frac{1}{\sqrt{2}} (|\uparrow\rangle|+p\rangle|\text{upper pixel}\rangle_{\text{screen}} + |\downarrow\rangle|-p\rangle|\text{lower pixel}\rangle_{\text{screen}}) \otimes |\text{everything else}\rangle$$

$$|\psi(t_3)\rangle = \frac{1}{\sqrt{2}} (|\uparrow\rangle|+p\rangle|\text{upper pixel}\rangle_{\text{screen}} | \text{me} \rangle + |\downarrow\rangle|-p\rangle|\text{lower pixel}\rangle_{\text{screen}} | \text{me} \rangle) \otimes |\text{everything else}\rangle$$

# Measurement in Quantum Mechanics

Time evolution with interaction between the system and measuring device



$$|\chi\rangle \otimes |A_0\rangle \rightarrow \sum_i c_i |i\rangle \otimes |A_i\rangle$$

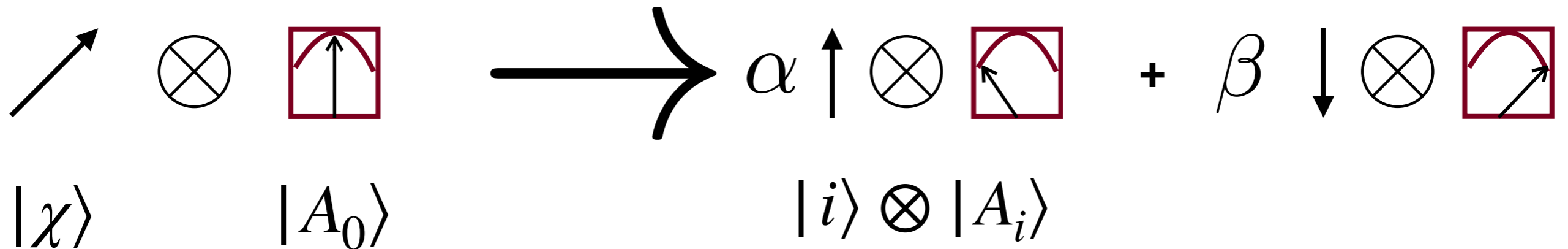
Prediction of Quantum Mechanics (“Many Worlds”)

Pick:  $\langle A_j | A_i \rangle = \delta_{ij} \implies \rho_{|\Psi\rangle} = \sum_i c_i c_i^* |i\rangle \langle i|$

“Interpret” as direct sum of “worlds”

# Measurement in Quantum Mechanics

Time evolution with interaction between the system and measuring device



In linear QM, can pick orthogonal basis vectors just by knowing the interaction Hamiltonian

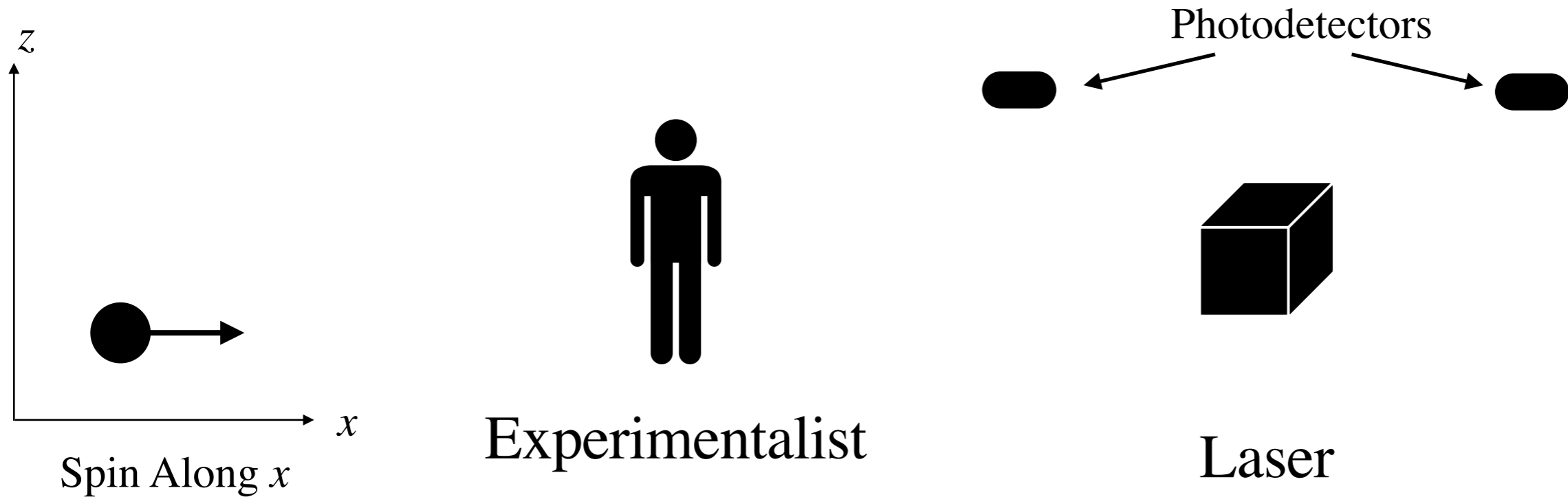
$$\langle A_j | A_i \rangle = \delta_{ij}$$

In non-linear QM, stationary states are generally **not** orthogonal — the effective Hamiltonian depends on the initial state of the system

$$\text{No Guarantee: } \langle A_j | A_i \rangle \neq 0$$

$$|\Psi\rangle \otimes |A_0\rangle \rightarrow \sum_i c_i |i\rangle \otimes |A_i\rangle + \epsilon \sum_{i,j} d_{i,j} |i\rangle \otimes |A_j\rangle \quad \text{Measurement noise}$$

# Linear Quantum Mechanics



Initial State :  $|\chi(0)\rangle$

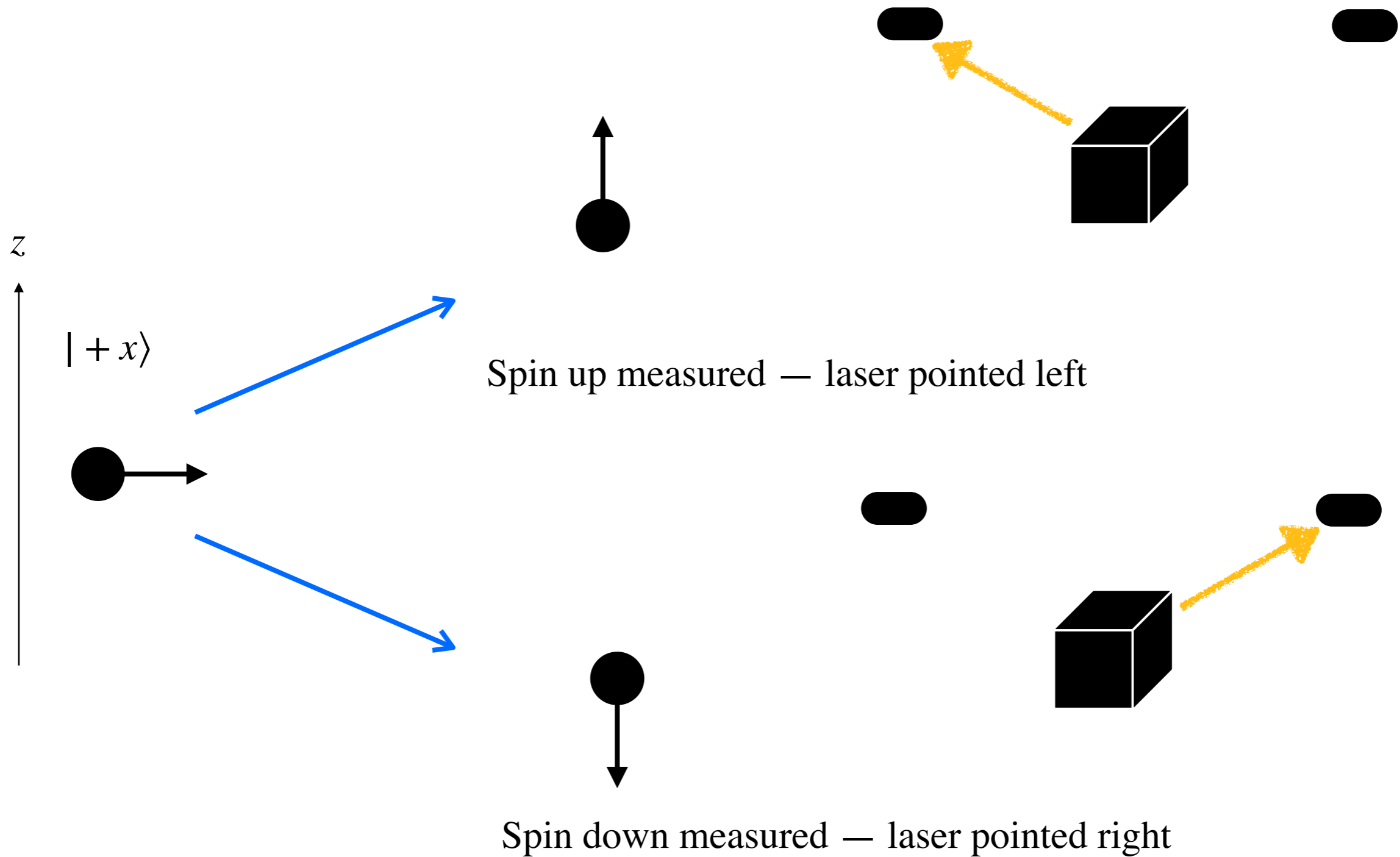
Represents Full Quantum State (spin, experimentalist...)

Goal: Create Macroscopic Superposition

Method: Measure spin along  $z$ .

Depending upon outcome, send laser along different directions

# Linear Quantum Mechanics



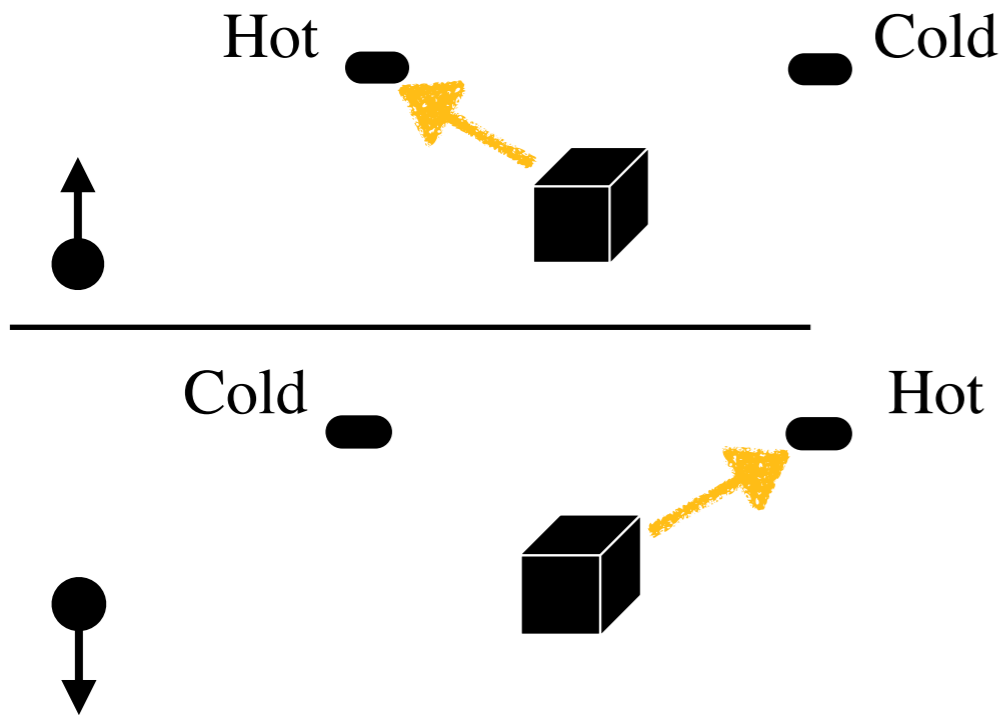
$$|\chi(t)\rangle = \frac{1}{\sqrt{2}} (|+z\rangle |L\rangle |Env_L\rangle + |-z\rangle |R\rangle |Env_R\rangle)$$



# Linear Quantum Mechanics

Which photodetectors light up?

$$\mathcal{H} \supset eA_\mu J^\mu$$



Transition Matrix Elements

$$\langle +z | \langle L | \langle Env_L | A_\mu(x_L) J^\mu(x_L) | +z \rangle | L \rangle | Env_L \rangle \neq 0$$

$$\langle +z | \langle L | \langle Env_L | A_\mu(x_R) J^\mu(x_R) | +z \rangle | L \rangle | Env_L \rangle = 0$$



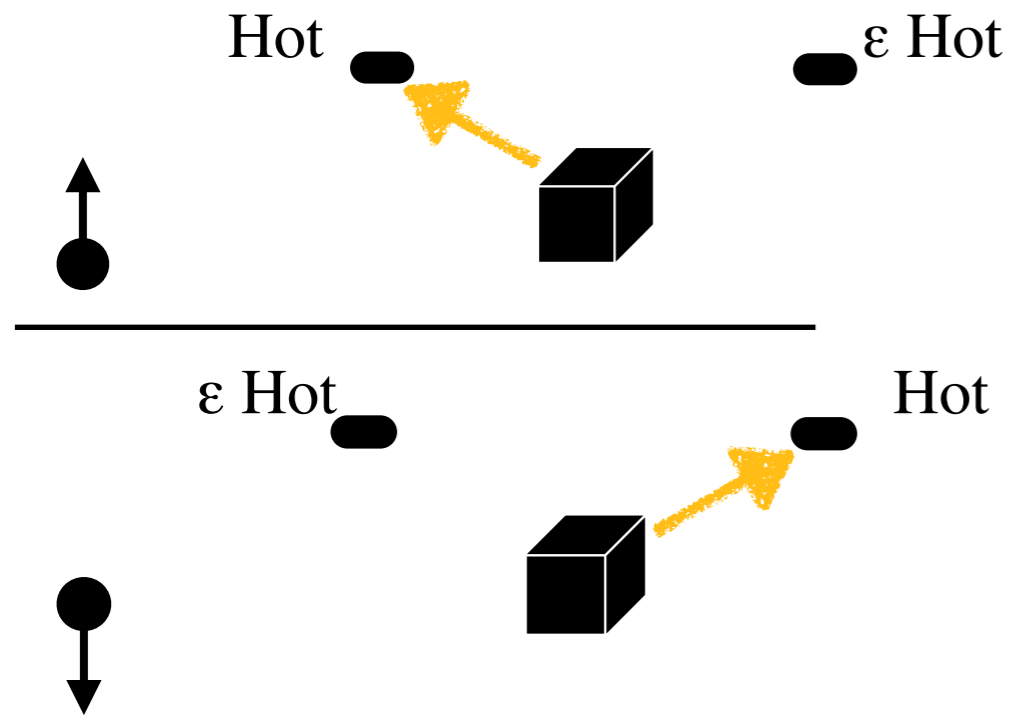
$$\langle L | A_\mu(x_R) | L \rangle = 0$$

$$|\chi(t)\rangle = \frac{1}{\sqrt{2}} (|+z\rangle |L\rangle |Env_L\rangle + |-z\rangle |R\rangle |Env_R\rangle)$$

# Non-Linear Quantum Mechanics

Which photodetectors light up?

$$\mathcal{H} \supset eA_\mu J^\mu + e\varepsilon \langle A_\mu \rangle J^\mu$$



Transition Matrix Elements

$$\langle +z | \langle L | \langle Env_L | \langle A_\mu(x_L) \rangle J^\mu(x_L) | +z \rangle | L \rangle | Env_L \rangle \neq 0$$

$$\langle +z | \langle L | \langle Env_L | \langle A_\mu(x_R) \rangle J^\mu(x_R) | +z \rangle | L \rangle | Env_L \rangle \neq 0$$

$$\langle \chi | A_\mu(x_L) | \chi \rangle \neq 0, \langle \chi | A_\mu(x_R) | \chi \rangle \neq 0$$

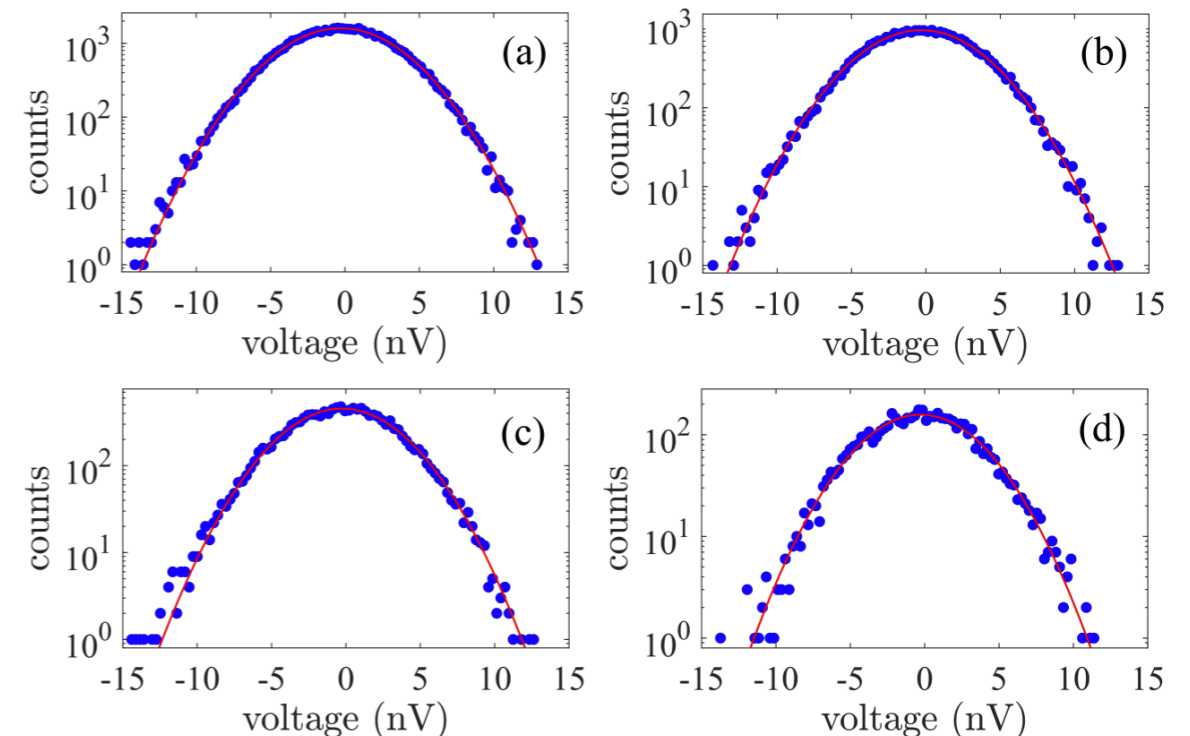
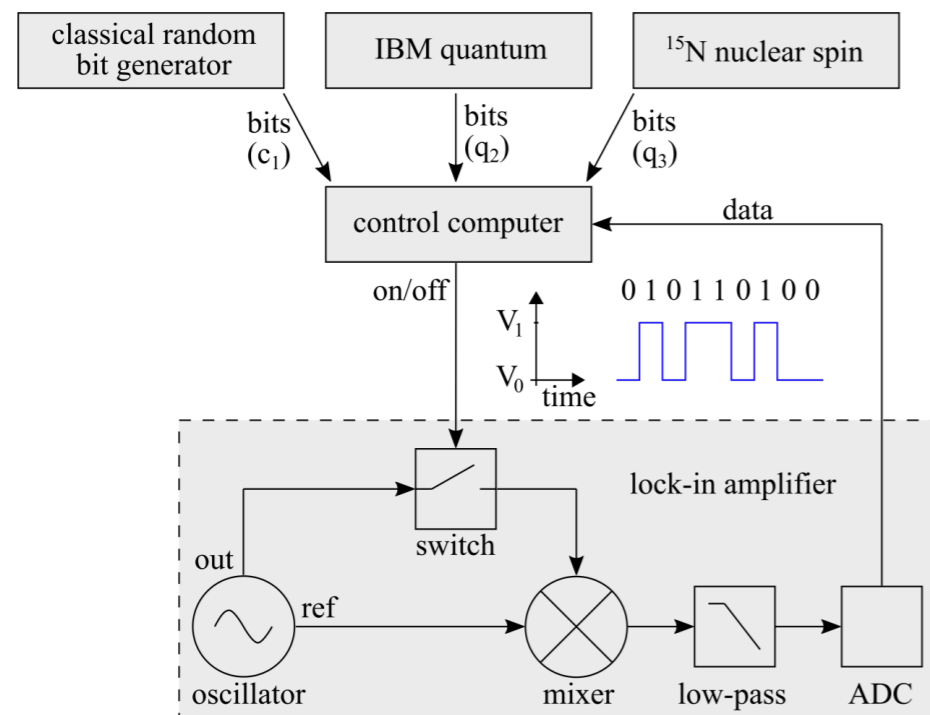
Communication between “worlds”

Non-linearity visible despite Environmental De-coherence!

Polchinski: “Everett Phone”

# Experimental limit on non-linear QM using a voltmeter and quantum bits

M. Polkovnikov, et al (2022)

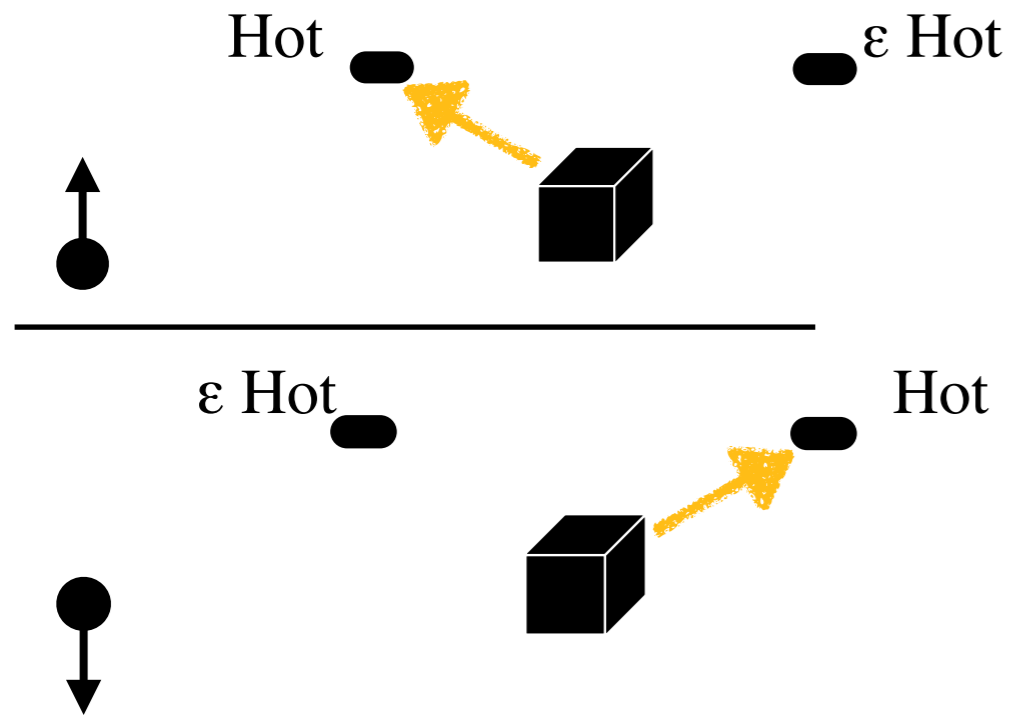


$$|\epsilon_\gamma| \leq 4.7 \times 10^{-11}$$

# **Quantum Pollution**

# Delicate Non-Linearity

$O$  performs the laser experiment on October 24th  
 - discovers non-linear quantum mechanics!



$$|\chi\rangle = \frac{1}{\sqrt{2}} (|L\rangle |O_L\rangle + |R\rangle |O_R\rangle)$$

Now  $O$  wants to repeat experiment

Suppose  $|O_U\rangle$  decides to run experiment at 9am on Oct 26  
 But  $|O_D\rangle$  runs experiment on 9am on Nov 3rd

State just after 9am on Oct 26

$$|\chi\rangle = \frac{1}{\sqrt{2}} \left( |L\rangle |O_L\rangle \frac{|L'\rangle |O'_L\rangle + |R'\rangle |O'_D\rangle}{\sqrt{2}} + |R\rangle |O_R\rangle \right)$$

# Delicate Non-Linearity

State after 9am on Oct 26

Compare with State on Oct 24

$$|\chi\rangle = \frac{1}{\sqrt{2}} \left( |L\rangle |O_L\rangle \frac{|L'\rangle |O'_L\rangle + |R'\rangle |O'_D\rangle}{\sqrt{2}} + |R\rangle |O_R\rangle \right) \quad |\chi\rangle = \frac{1}{\sqrt{2}} (|L\rangle |O_L\rangle + |R\rangle |O_R\rangle)$$

$$\langle L | \langle O_L | \langle L' | \langle O'_L | \langle A_\mu(x_R) \rangle J^\mu(x_R) | \chi(t = \text{Oct 26}) \rangle = \frac{1}{2} \langle L | \langle O_L | \langle A_\mu(x_R) \rangle J^\mu(x_R) | \chi(t = \text{Oct 24}) \rangle$$

Effect is 1/2 of prior effect!

But, full effect if  $O_U$  and  $O_D$  perform experiment at same time!

Quantum Pollution: Without adequate care, superpositions may diverge wildly, preventing exploitability. Not automatic - but need careful protocols!

But hasn't there already been dilution?

# What part of the wave function...

$$|\chi\rangle = \alpha |Us\rangle + \beta |Them\rangle$$

$$\mathcal{H} \supset eA_\mu J^\mu + e\varepsilon \langle A_\mu \rangle J^\mu$$

$$|\chi\rangle = \alpha |Us\rangle + \beta |Them\rangle \rightarrow \langle \chi | A_\mu | \chi \rangle = |\alpha|^2 \langle U | A_\mu | U \rangle + |\beta|^2 \langle T | A_\mu | T \rangle$$

$$\langle \chi | A_\mu | \chi \rangle \rightarrow |\alpha|^2 \langle U | A_\mu | U \rangle$$

For  $\alpha \ll \beta$ , the wave function is dominated by something we are not a part of. Can't turn on coherent fields over there.

Local exploitability completely determined by unchangeable initial conditions dramatic difference from linear QM

# Classical World?

$$|U(t)\rangle = | \text{🌍} \rangle + \delta | \blacksquare \rangle \quad \text{Or} \quad |U(t)\rangle = \delta | \text{🌍} \rangle + | \blacksquare \rangle$$

Are there natural quantum amplifiers, for e.g. in chaotic systems?



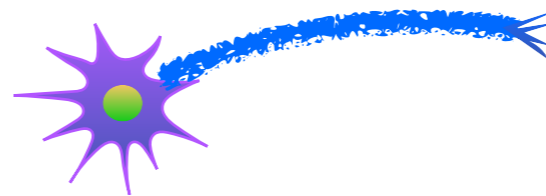
$$\Delta x \sim 100\text{nm}$$

Changing classical evolution of a system requires coherent motion of N atoms

Probability that N atoms coherently move in some way:  $p^N$

With  $p \sim O(1)$  scattering probability

What about the weather?  
What about my brain??



100's + ions to get one neuron to fire

$$|U(t)\rangle = | \text{🌍} \rangle + \delta | \blacksquare \rangle \quad \text{Reasonable}$$

**Quantum Amplifiers are Hard!**

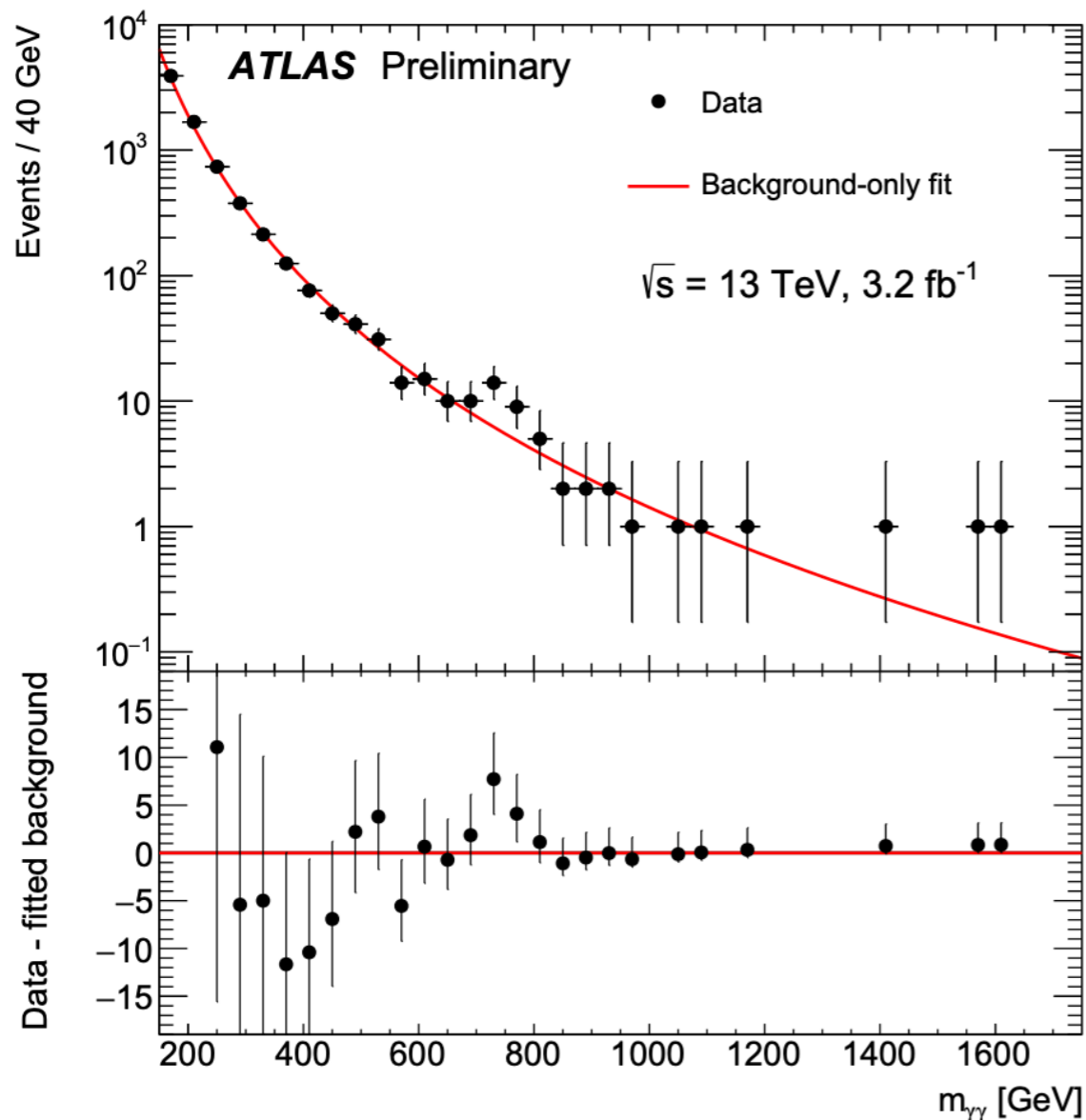


# Natural Quantum Dilution

The '750 GeV' resonance!

Search for resonances decaying to photon pairs in  $3.2 \text{ fb}^{-1}$  of  $pp$  collisions at  $\sqrt{s} = 13 \text{ TeV}$  with the ATLAS detector

Dec 15, 2015



ATLAS-CONF-2015-081

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Have we been diluting our wave function on Earth for the past 100 years?

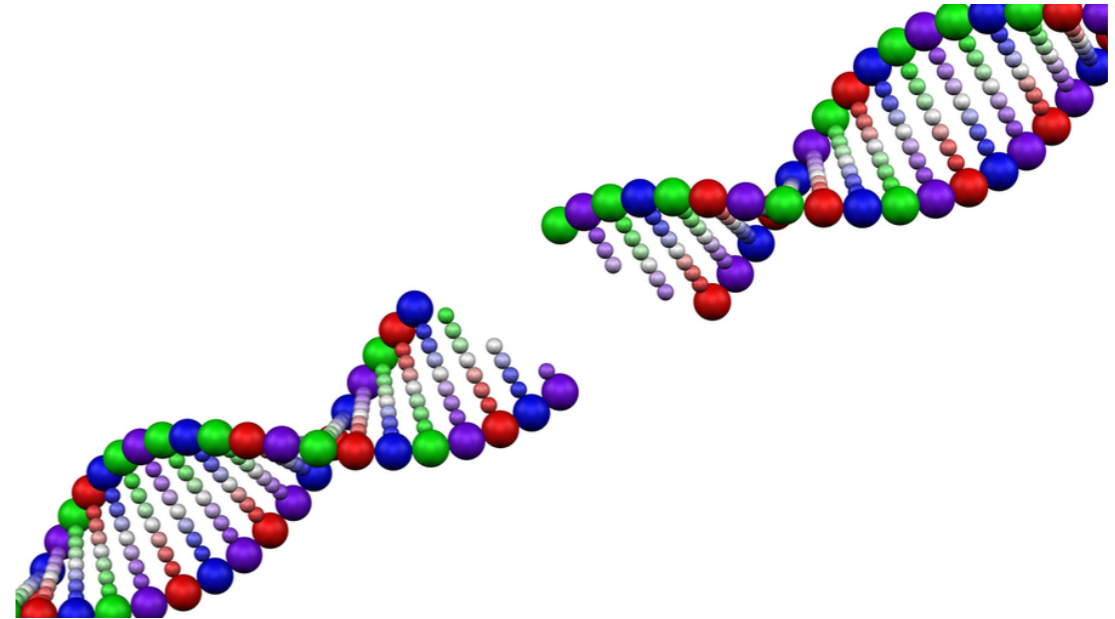
# Evolutionary Dilution?

Is  $N \sim O(\text{few})$  for evolution?

Maybe for RNA/DNA?

RNA formation?

Evolution in an amplifier!



$$|U(t = 0)\rangle = \text{[Image of a rocky, barren landscape]}$$



$$|U(t)\rangle = |\text{Earth}\rangle(|\text{Human}\rangle + |\text{No life}\rangle + \dots)$$

$$|U(t = 0)\rangle = \text{[Image of a rocky, barren landscape]}$$



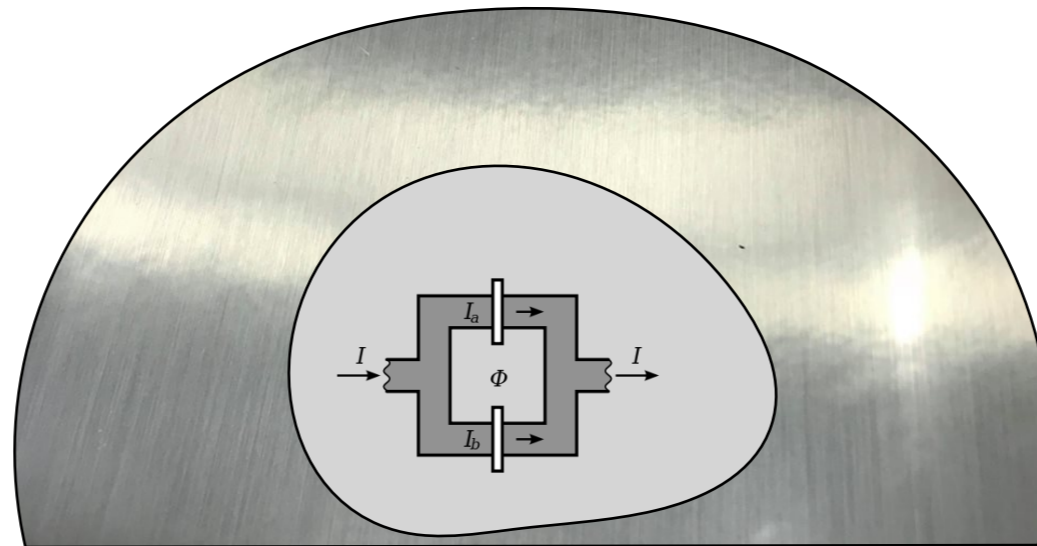
$$|U(t)\rangle = |\text{Earth}\rangle(|\text{Human}\rangle + |\text{Alien}\rangle + \dots)$$

# Tests for a Quantum-Diluted Earth

Look for coherent fields turned on in all parts of the wavefunction:  
The magnetic field of the Earth!

$$eJ^\mu(A_\mu + \epsilon_\gamma \langle A_\mu \rangle)$$

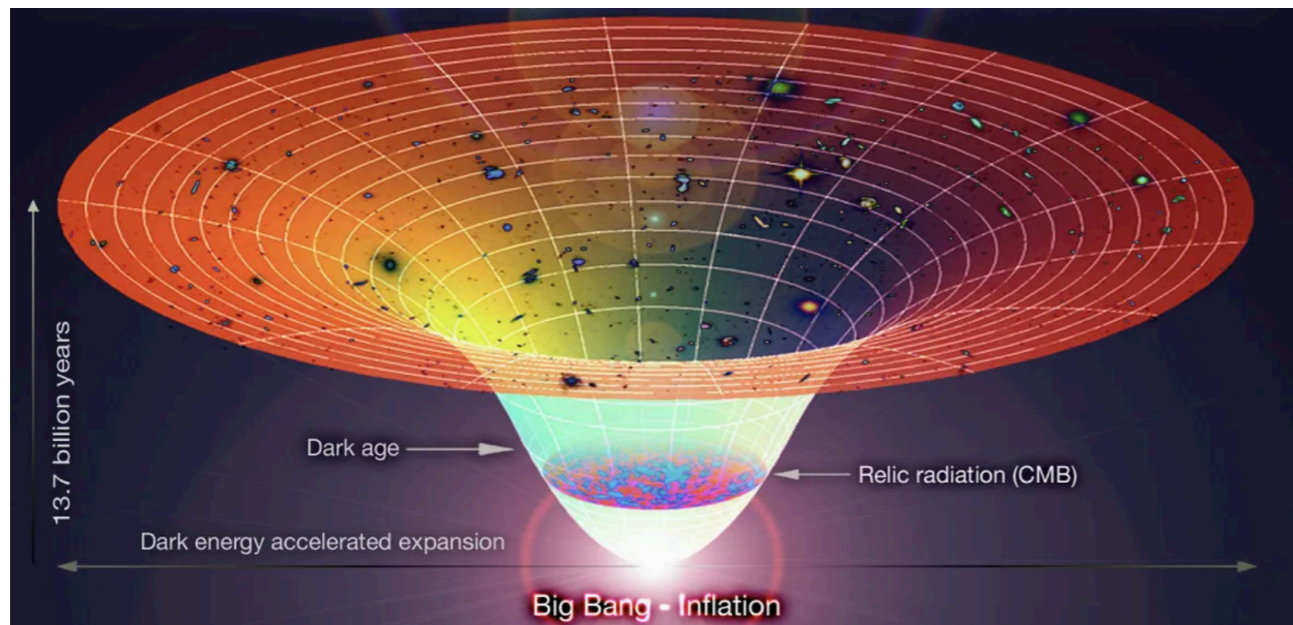
Build a magnetic field structure unique to our part of the wave function and measure the field inside.



$$|U(t)\rangle = |\text{🌍}\rangle(\alpha |\text{shield}\rangle + \beta |\text{No shield}\rangle)$$

# Cosmological Quantum Amplifier: Inflation

Standard cosmic inflation:  
rapidly places quantum state in a homogenous and isotropic state  
(Bunch-Davies Vacuum )



How could homogeneous state  
become inhomogeneous?

Answer: Massive Superposition of  
Statistically Similar Universes!

$$|\chi\rangle = \sum_i c_i |U_i\rangle, \quad c_i \sim e^{-N}$$

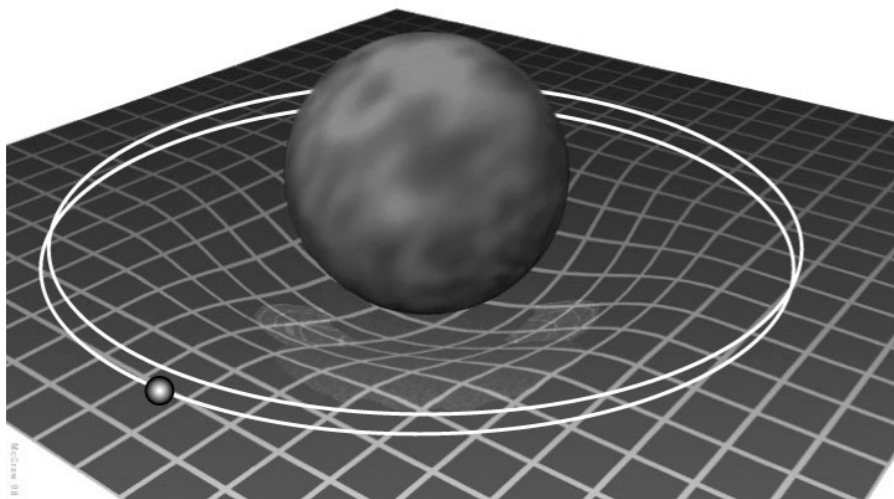
Most of the Universe: The space-time point where the  
Earth is located is in intergalactic space!

# Tests for a Quantum-Diluted Universe(!)

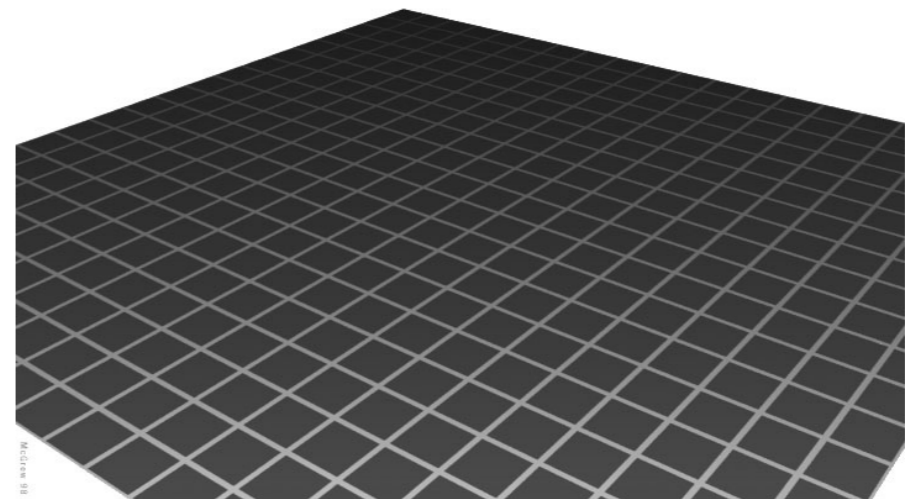
Look for coherent fields turned on in all parts of the wavefunction:  
The magnetic field of the Earth!

$$T^{\mu\nu}(g_{\mu\nu} + \epsilon_G \langle g_{\mu\nu} \rangle) + \dots$$

Objects in our part of the wave function will produce  
different gravitational fields than the average



+



# Tests for a Quantum-Diluted Universe

$$g_{\mu\nu} \rightarrow \frac{g_{\mu\nu} + \epsilon_G \langle g_{\mu\nu} \rangle}{1 + \epsilon}$$

$$g_s = - \left(1 - \frac{r_s}{r}\right) dt^2 + \frac{dr^2}{\left(1 - \frac{r_s}{r}\right)} + r^2 d\Omega^2 \quad \langle g \rangle = -dt^2 + dr^2 + r^2 d\Omega^2$$

Renormalize and Expand

$$g_{\text{eff}} \simeq \left[ - \left(1 - \frac{R_s}{r}\right) dt^2 + \left(1 + \frac{R_s}{r} + \left(\frac{R_s}{r}\right)^2 (1 + \epsilon_G)\right) dr^2 \right] + r^2 d\Omega^2$$

Looks like a long-distance modification of gravity!

Corrects second-order GR term  $\rightarrow$  Strong field tests of GR

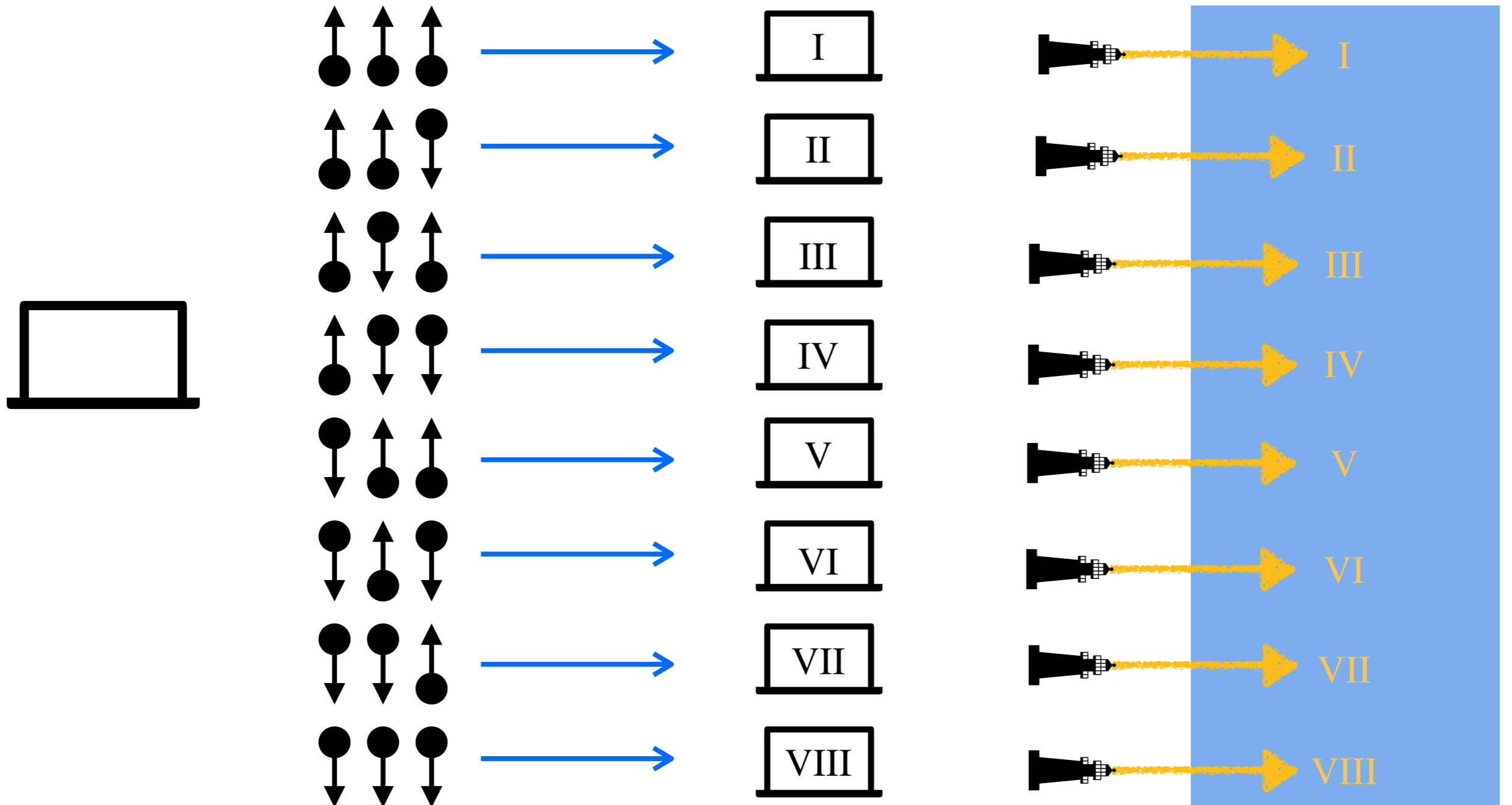
Will potentially make a black hole horizon more singular

# **Implications**

# If we have a Classical Universe

Macroscopic superpositions can be produced at will.

Parallelize any computation:



Quantum Computing!



# **Conclusions**

# Conclusions

There is a consistent way to explore non-linear deviations from QM

Locality makes many past tests insensitive — new probes required

NL effects can be experimentally tested by amplifying quantum measurements

Quantum amplification in the history of the universe suppresses access to local non-linearities —> Linear QM is an attractor solution.

If locally diluted, non-zero fields across the wave function could be detected (Earth's magnetic field, cosmological metric)

If NLQM is locally accessible, it will radically change what we can do technologically

**Thank you!**