

QUALIFYING EXAMINATION, Part 1

Solutions

Problem 1: Mathematical Methods

(a) For $r > 0$ we find

$$\nabla^2 \left(\frac{1}{r} \right) = \frac{1}{r^2} \frac{\partial}{\partial r} r^2 \frac{\partial}{\partial r} \left(\frac{1}{r} \right) = \frac{1}{r^2} \frac{\partial}{\partial r} r^2 \left(-\frac{1}{r^2} \right) = 0$$

However for $r = 0$ we get $\frac{0}{0}$ because of the $\frac{1}{r^2}$ factor in front. To calculate the limit $r \rightarrow 0$, we integrate $\nabla^2 \left(\frac{1}{r} \right)$ over a small sphere of radius a surrounding the origin and use Gauss' theorem (converting a volume integral to a surface integral)

$$\begin{aligned} \int_{r \leq a} d^3 \vec{r} \nabla^2 \left(\frac{1}{r} \right) &= \int_{r \leq a} d^3 \vec{r} \nabla \cdot \nabla \left(\frac{1}{r} \right) \\ &= \int_{r=a} \nabla \left(\frac{1}{r} \right) \cdot d\vec{S} = \int_{r=a} \left(-\frac{1}{a^2} \right) \hat{\mathbf{e}}_r \cdot d\vec{S} = -4\pi . \end{aligned}$$

It follows that

$$\nabla^2 \left(\frac{1}{r} \right) = -4\pi \delta^3(\vec{r}) .$$

Note: $\frac{1}{r}$ is the solution to Poisson's equation with a point charge at the origin with charge $4\pi\epsilon_0$.

(b)

$$\Gamma(1) = \int_0^\infty t^0 e^{-t} dt = -e^{-t} \Big|_0^\infty = 1 .$$

(c)

$$\Gamma(x+1) = \int_0^\infty t^x e^{-t} dt = -t^x e^{-t} \Big|_0^\infty + \int_0^\infty x t^{x-1} e^{-t} dt = 0 + x\Gamma(x) ,$$

where the surface term vanishes at $t = 0$ if $x \geq 0$.

It follows that $\Gamma(n+1) = n\Gamma(n)$ which, given $\Gamma(1) = 1$, means

$$\Gamma(n+1) = n! .$$

(d)

$$\ln n! = \ln [n(n-1)(n-2)\dots 1] = \sum_{m=1}^n \ln m \simeq \int_1^n \ln x dx = n \ln n - n + 1 ,$$

where we can drop the 1 compared to n . Exponentiating, we find

$$n! \approx e^{n \ln n - n} = n^n e^{-n} .$$

(e) We have

$$n! = \Gamma(n+1) = \int_0^\infty t^n e^{-t} dt .$$

We rescale $t = nz$ to obtain

$$n! = n^{n+1} \int_0^\infty z^n e^{-nz} dz = n^{n+1} \int_0^\infty e^{-nz+n \ln z} dz \equiv n^{n+1} \int_0^\infty e^{-nf(z)} dz ,$$

where

$$f(z) = z - \ln z .$$

The saddle point $z = z^*$ is determined by the condition $f'(z) = z - 1/z = 0$, giving $z^* = 1$. We expand $f(z)$ to second order around $z^* = 1$ to find

$$\Gamma(n+1) = n^{n+1} \int_0^\infty e^{-nf(z)} dz \simeq n^{n+1} e^{-nf(z^*)} \int_0^\infty e^{-n \frac{1}{2!} f''(z^*) (z-z^*)^2} dz .$$

Using $f''(z^*) = \frac{1}{z^2} \Big|_{z^*} = 1$, we obtain

$$n! = n^n e^{-n} \sqrt{2\pi n} (1 + \text{corrections}) ,$$

where we have changed variables to $w = z - z^*$ and extended the limits to $w = \pm\infty$ (the integrand is negligible far away from z^*). We can then use the familiar result for a Gaussian integral

$$\int_{-\infty}^\infty e^{-\alpha w^2} dw = \sqrt{\frac{\pi}{\alpha}} .$$

The initial rescaling was done to get a large factor n in the exponent in front of $f(z)$ (to justify the saddle-point approximation). It is also possible to carry out the saddle-point approximation directly for $f(t) = n \ln t - t$, in which case the saddle-point condition $f'(t^*) = 0$ gives $t^* = n$.

Problem 2: Classical Mechanics

(a) The first Lagrangian is invariant under translations in the x direction $x \rightarrow x + a$ and therefore p_x (the generator of translations in the x direction) is conserved. It does not depend on time explicitly so $\partial\mathcal{L}/\partial t = -dH/dt = 0$ and the Hamiltonian $H = E$ is conserved. The Lagrangian is not invariant under translations in y or under rotations, and therefore p_y and L_z are not conserved.

The second Lagrangian is not invariant under translations (in x or y) and therefore p_x and p_y are not conserved. It is rotationally invariant and explicitly time independent, and therefore L_z and E are conserved.

The third Lagrangian is invariant under translations in x but not under translations in y or under rotations. Therefore p_x is conserved but p_y and L_z are not conserved. Because of the explicit time dependence in \mathcal{L} , E is not conserved.

(b) The potential is time independent, and hence the total energy is conserved:

$$E = E_0 = \frac{1}{2}mv_0^2 = \text{const.}$$

The potential is central and provides no torque, hence the angular momentum around the scattering center is conserved

$$\vec{L} = \vec{L}_0 = mv_0b\hat{z} = \text{const.}$$

$L_x = L_y = 0$ guaranteeing that the motion is in the x - y plane.

(c) Using the effective potential for the reduced 1D motion in the radial coordinate r , the energy can be written as

$$E = \frac{1}{2}m\dot{r}^2 + U_{\text{eff}}(r) = \frac{1}{2}m\dot{r}^2 + \frac{L^2}{2mr^2} - \frac{k}{r^2} = \frac{1}{2}mv_0^2.$$

At the distance $r = r_{\min}$ of closest approach $\dot{r} = 0$, and we have

$$\frac{(mv_0b)^2}{2mr_{\min}^2} - \frac{k}{r_{\min}^2} = \frac{1}{2}mv_0^2.$$

Therefore

$$r_{\min} = \sqrt{\frac{mv_0^2b^2/2 - k}{mv_0^2/2}} = \sqrt{b^2 - \frac{2k}{mv_0^2}}.$$

Since $k > 0$, $r_{\min} < b$. Qualitatively this is expected since the potential is attractive.

(d) The horizontal coordinate of the point particle is given by $x + X$, while its vertical coordinate is $y = x \tan \alpha$. The kinetic energy T of the system is then

$$T = \frac{1}{2}M\dot{X}^2 + \frac{1}{2}m(\dot{X} + \dot{x})^2 + \frac{1}{2}m\dot{y}^2 = \frac{1}{2}(M + m)\dot{X}^2 + m\dot{X}\dot{x} + \frac{1}{2}m(1 + \tan^2 \alpha)\dot{x}^2$$

The potential energy of the particle is

$$V = mgy = mgx \tan \alpha .$$

The Lagrangian of the system is therefore:

$$\mathcal{L} = T - V = \frac{1}{2}(M + m)\dot{X}^2 + m\dot{X}\dot{x} + \frac{1}{2}m(1 + \tan^2 \alpha)\dot{x}^2 - mgx \tan \alpha .$$

The equations of motion are:

$$\frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \dot{X}} \right) - \frac{\partial \mathcal{L}}{\partial X} = 0 \rightarrow (M + m)\ddot{X} + m\ddot{x} = 0 .$$

$$\frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \dot{x}} \right) - \frac{\partial \mathcal{L}}{\partial x} = 0 \rightarrow m\ddot{X} + m(1 + \tan^2 \alpha)\ddot{x} + mg \tan \alpha = 0 .$$

From the first equation we get

$$\ddot{X} = -\frac{m}{M + m}\ddot{x} .$$

Substituting in the second equation, we find

$$\ddot{x} = -\frac{(M + m)g \sin \alpha \cos \alpha}{M + m \sin^2 \alpha} .$$

We then have

$$\ddot{X} = \frac{mg \sin \alpha \cos \alpha}{M + m \sin^2 \alpha} .$$

Problem 3: Electromagnetism I

(a) Using Coulomb's law

$$V(r, \theta) = \frac{q}{4\pi\epsilon_0} \left(\frac{1}{r_+} + \frac{1}{r_-} - \frac{2}{r} \right)$$

where (applying the law of cosines)

$$r_{\pm} = \left(r^2 + a^2 \mp 2ra \cos \theta \right)^{1/2} .$$

(b) Expanding $1/r_{\pm}$ to order a^2/r^2 (all higher order terms will vanish in the limit), we have

$$\frac{1}{r_{\pm}} \approx \frac{1}{r} \left(1 - \frac{1}{2} \frac{a^2}{r^2} \pm \frac{a}{r} \cos \theta + \frac{3}{2} \frac{a^2}{r^2} \cos^2 \theta \right) .$$

Substituting into the potential, we get

$$V(r, \theta) = \frac{Q}{4\pi\epsilon_0 r^3} \left(3 \cos^2 \theta - 1 \right) = \frac{Q}{2\pi\epsilon_0 r^3} P_2(\cos \theta) ,$$

where we have used the expression for the second Legendre polynomial P_2 .

(c) Since the conducting sphere is grounded $V(r = R) = 0$. We also know that $V(r \rightarrow \infty) = 0$. The uniqueness of the solution of Laplace's equation with boundary conditions gives us $V(r) = 0$ outside the sphere.

(d) Inside the sphere, we use the general solution of Laplace's equation

$$V_{\text{in}} = \sum \left(A_l r^l + \frac{B_l}{r^{l+1}} \right) P_l(\cos \theta) .$$

The potential is continuous across $r = R$. Using the result in (c), we have

$$V_{\text{in}}(R, \theta) = 0 .$$

Since this is true for all angles θ , the sum must vanish term by term. This gives

$$-A_l R^l = \frac{B_l}{R^{l+1}} .$$

In the limit $r \rightarrow 0$ ($r \ll R$), we match the solution to the result from part. This gives us

$$B_{l \neq 2} = 0 , \quad B_2 = \frac{Q}{2\pi\epsilon_0} .$$

Using the first set of relations between A_l and B_l , we find

$$A_{l \neq 2} = 0, \quad A_2 = -\frac{B_2}{R^5}.$$

Thus the potential inside is given by

$$V_{\text{in}}(r, \theta) = \frac{Q}{2\pi\epsilon_0 r^3} \left(1 - \frac{r^5}{R^5}\right) P_2(\cos \theta).$$

Problem 4: Electromagnetism II

(a) The voltage per turn has to balance the induced electro-motive force (emf)

$$V = -\varepsilon = \frac{d\Phi}{dt} = \pi a^2 \frac{dB}{dt}$$

(b) Using a rectangular Amperian loop outside the solenoid, one can show that the field outside (parallel to the cylinder axis) is independent of the radial distance r . Since the field vanishes at $r \rightarrow \infty$, it must be zero everywhere outside the cylinder.

(c) From symmetry considerations \vec{B} is along the cylinder's axis. Using a rectangular Amperian loop for the field H with one side inside the solenoid and another side outside, gives

$$H = nI \quad \text{or} \quad I = \frac{B(t)}{\mu n} .$$

(d) We use a circular loop C around the cylinder with radius r . By symmetry the magnitude of \vec{A} is constant along the loop and is tangential along the same direction as the current through the curled wire. Using Stokes theorem with $\nabla \times \vec{A} = \vec{B}$ we have

$$\oint_C \vec{A} \cdot d\vec{l} = \int_S \vec{B} \cdot d\vec{S} \equiv \Phi ,$$

where Φ is the magnetic flux through the loop.

We find for $r > a$

$$\begin{aligned} 2\pi r A(r, t) &= B(t) \pi a^2 , \\ A(r, t) &= \frac{a^2}{2r} B(t) , \end{aligned}$$

and for $r < a$

$$\begin{aligned} 2\pi r A(r, t) &= \pi r^2 B(t) , \\ A(r, t) &= \frac{r}{2} B(t) . \end{aligned}$$

(e) The field is uniform inside the torus. Since $\nabla \cdot \vec{B} = 0$, the flux of B is constant around the torus. Since the cross section of the torus is the same everywhere, B must be a constant inside the torus. Since the gap h is small, B in the gap will be near homogeneous (negligible fringe fields) and will have the same value as inside the torus.

To find the value of B , we use a circular Amperian loop of radius Z around the torus. We have

$$\oint_Z \vec{H} d\vec{l} = NI$$

or

$$\frac{B}{\mu}(2\pi Z - h) + \frac{B}{\mu_0}h = NI .$$

This gives

$$\begin{aligned} B &= \frac{\mu\mu_0 NI}{\mu_0(2\pi Z - h) + \mu h} \\ &\approx \frac{\mu\mu_0 NI}{\mu_0 2\pi Z + \mu h} . \end{aligned}$$

The current I in the wire is $I = \frac{V}{R}$, giving the final answer

$$B = \frac{\mu\mu_0 NV}{R(2\mu_0\pi Z + \mu h)} .$$