

# QUALIFYING EXAMINATION, Part 1

## Solutions

### Problem 1: Mathematical Methods

(a) Keeping only the lowest power of  $x$  needed, we find

$$\begin{aligned}\frac{1}{x^2} - \frac{1}{\sin^2 x} &= \frac{1}{x^2} - \frac{1}{(x - x^3/6 \dots)^2} \\ &= \frac{1}{x^2} \left( 1 - \frac{1}{1 - 2x^2/3} \right) \\ &= \frac{1}{x^2} [1 - (1 + 2x^2/3!)] \\ &= \frac{1}{x^2} \left( \frac{-2x^2}{3!} \right) = -\frac{1}{3}.\end{aligned}$$

(b) The sum of finite number of terms is finite. For the rest, the ratio of successive terms test gives the condition

$$\begin{aligned}\lim_{n \rightarrow \infty} \left| \frac{\frac{(n+1)!^2 z^{n+1}}{(2n+2)!}}{\frac{(n!)^2 z^n}{(2n)!}} \right| &< 1 \\ \lim_{n \rightarrow \infty} \left| \frac{(n+1)^2 z}{(2n+2)(2n+1)} \right| &< 1 \\ |z| &< 4.\end{aligned}$$

(c) We expand

$$y = \sum_m c_m x^m.$$

Then

$$\begin{aligned}y'' &= \sum_m c_m x^{m-2} m(m-1) \\ &= \sum_m c_{m+2} x^m (m+2)(m+1) \quad (\text{dummy label } m \rightarrow m+2) \\ x^2 y'' &= \sum_m c_m x^2 m(m-1) \\ -2xy' &= -2 \sum_m c_m x^m m \\ 2y &= 2 \sum_m c_m x^m.\end{aligned}$$

Collecting the coefficient of  $x^m$  and setting it to 0, we find

$$\frac{c_{m+2}}{c_m} = -\frac{(m-1)(m-2)}{(m+1)(m+2)}.$$

Choosing  $c_0$  and  $c_1$ , arbitrarily, we find  $c_2 = -1$  and all others 0. Thus the general solution is given by

$$y = c_0(1 - x^2) + c_1x.$$

(d)

$$\begin{aligned} \int_0^\infty \frac{x \sin x}{x^2 + a^2} dx &= \frac{1}{2} \int_{-\infty}^\infty \frac{x \sin x}{x^2 + a^2} dx \quad (\text{even integrand}) \\ &= \Im \left( \frac{1}{2} \int_{-\infty}^\infty \frac{ze^{iz}}{z^2 + a^2} dz \right) \quad (\Im \text{ is imaginary part}) \\ &= \Im \left[ \frac{1}{2} 2\pi i \frac{iae^{-a}}{2ia} \right] \quad (\text{close contour in UHP, pick up residue at } z = ia) \\ &= \frac{\pi}{2} e^{-a}. \end{aligned}$$

## Problem 2: Classical Mechanics

(a) The Hamiltonian is

$$H = \frac{p^2}{2m} - ax^2 + \lambda x^4 .$$

Hamilton's equations of motion are given by

$$\dot{p} = -\frac{\partial H}{\partial x} = 2ax - 4\lambda x^3, \quad \dot{x} = \frac{\partial H}{\partial p} = \frac{p}{m} .$$

(b) Equilibrium requires  $\dot{p} = 0 = x(2a - 4\lambda x^2)$  so that  $x = \pm\sqrt{\frac{a}{2\lambda}}$  and  $x = 0$ .

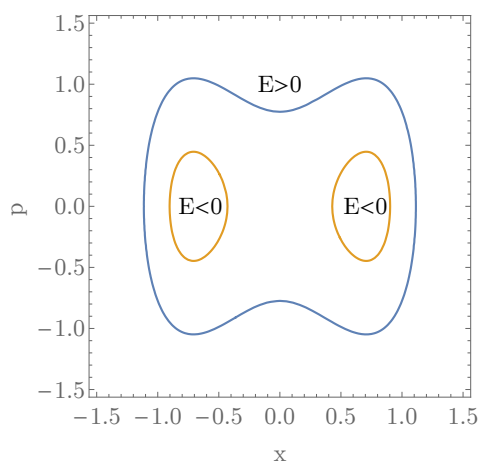
Around  $x = 0$  we have  $U'' = -2a$  so it is unstable since the second derivative is negative.

Around  $x = \pm\sqrt{\frac{a}{2\lambda}}$  we have

$$U(\pm\sqrt{\frac{a}{2\lambda}} + \epsilon) = -a(\pm\sqrt{\frac{a}{2\lambda}} + \epsilon)^2 + \lambda(\pm\sqrt{\frac{a}{2\lambda}} + \epsilon)^4 = -\frac{a^2}{4\lambda} + 2a\epsilon^2 + \dots$$

Thus,  $U'' = 4a > 0$  so the equilibrium points are stable and the effective spring constant for the oscillations is  $k_{eff} = 4a$ . The oscillation frequency is then  $\omega = \sqrt{\frac{k_{eff}}{m}} = \sqrt{\frac{4a}{m}}$ .

(c)



(d) The principle of least action is a variational principle which states that the system evolves in a way which extremizes the classical action  $S$ .

In the Hamiltonian formulation, the classical action for evolution from  $t_1$  to  $t_2$  is  $S = \int_{t_1}^{t_2} (p\dot{x} - H)$  and the principle states that the variation is  $\delta S = 0$  under small changes in the configuration  $\{x(t) \rightarrow x(t) + \delta x(t), p(t) \rightarrow p(t) + \delta p(t)\}$ , where the coordinates are held fixed at the endpoints,  $\delta x(t_1) = \delta x(t_2) = \delta p(t_1) = \delta p(t_2) = 0$ . (We would also accept a formulation of the principle of least action that leads to Lagrange's equations.)

In the situation asked about in the problem, the classical action is

$$S = \int_0^T (p\dot{x} - H) = - \int_0^T U \left( \pm \sqrt{\frac{a}{2\lambda}} \right) = -T \left( -\frac{a^2}{4\lambda} \right) = \frac{a^2 T}{4\lambda} .$$

### Problem 3: Electromagnetism I

(a) The trace of the second rank tensor  $Q$  is

$$\sum_{i=1}^3 Q_{ii} = \int \sum_i (3x_i'^2 - r'^2) \rho(x') d^3x' = \int (3r'^2 - 3r'^2) \rho(x') d^3x' = 0.$$

$Q_{ij}$  is real, symmetric, and traceless. Therefore it has 5 independent components.

(b)  $q$  and  $\mathbf{p}$  are, respectively, the total charge and dipole moment of the charge distribution. They are given by

$$q = \int \rho(\mathbf{x}') d^3x', \quad \mathbf{p} = \int \rho(\mathbf{x}') \mathbf{r}' d^3x'.$$

(c)

$$q = e - e + e - e = 0.$$

(d)

$$\begin{aligned} p_x &= ea - ea - -ea - ea = 0 \\ p_y &= -ea + ea - ea - -ea = 0 \\ p_z &= 0. \end{aligned}$$

Thus  $\mathbf{p} = \mathbf{0}$ .

(e)

$$\begin{aligned} Q_{xx} &= (3a^2 - 2a^2)(e - e + e - e) = 0 \\ Q_{yy} &= (3a^2 - 2a^2)(-e + e + e - e) = 0 \\ Q_{zz} &= 0 \\ Q_{xy} &= 3a^2(-e) - 3a^2e + 3a^2(-e) - 3a^2e = -12a^2 \\ Q_{xz} &= 0 \\ Q_{yz} &= 0. \end{aligned}$$

Thus

$$Q = \begin{pmatrix} 0 & -12ea^2 & 0 \\ -12ea^2 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}.$$

(f) Since  $q$  and  $\mathbf{p}$  are zero, only the quadrupole term survives. We then have

$$\Phi(\mathbf{x}) = 2 \frac{1}{2} \left( -12ea^2 \frac{xy}{r^5} \right).$$

The electric field is

$$\begin{aligned} \mathbf{E} &= -\nabla \left( -12ea^2 \frac{xy}{r^5} \right) \\ &= \left( \frac{12ea^2y}{r^5} - \frac{60ea^2x^2y}{r^7}, \frac{12ea^2x}{r^5} - \frac{60ea^2xy^2}{r^7}, -\frac{60ea^2xyz}{r^7} \right). \end{aligned}$$

### Problem 4: Electromagnetism II

(a) Define  $k = \omega/c$ . The electric and magnetic fields are now given by

$$\vec{E} = E_0 \sin(kz - \omega t) \hat{x} ,$$

and

$$\vec{B} = B_0 \sin(kz - \omega t) \hat{y} .$$

We can relate  $E_0$  and  $B_0$  by Maxwell's equations; in SI units we have  $B_0 = E_0/c$ , while in Gaussian units, we have  $E_0 = B_0$ .

(b) The Poynting vector is

$$\begin{aligned} \vec{S} &= \frac{1}{\mu_0} \vec{E} \times \vec{B} \text{ (SIunits)} \\ \vec{S} &= \frac{c}{4\pi} \vec{E} \times \vec{B} \text{ (Gaussianunits)} \end{aligned}$$

Recall that the average of  $\sin^2$  over a period is  $1/2$ . Substituting and taking time-averages, we have

$$\begin{aligned} \langle S \rangle &= \frac{E_0^2}{2\mu_0 c} \text{ (SIunits)} \\ &= \frac{cE_0^2}{8\pi} \text{ (Gaussianunits)} \end{aligned}$$

(c) The acceleration of the electron is  $a = eE/m$ ; plugging this into the Larmor formula (and time averaging) we get

$$\begin{aligned} \langle P_{\text{rad}} \rangle &= \frac{1}{4\pi\epsilon_0} \frac{1}{3} \frac{e^4 E_0^2}{m_e^2 c^3} \text{ (SIunits)} \\ &= \frac{1}{3} \frac{e^4 E_0^2}{m_e^2 c^3} \text{ (Gaussianunits)} . \end{aligned}$$

(d) The ratio of the previous two parts is  $\sigma$ . This is simplest in Gaussian units, where we immediately find that

$$\sigma = \frac{8\pi}{3} \frac{e^4}{m_e^2 c^4} = \frac{8\pi}{3} r_e^2 .$$

We get the same result in SI units, using  $\mu_0\epsilon_0 = 1/c^2$ .

(e) The integral of  $d\sigma/d\Omega$  over all angles must equal  $\sigma$ . This implies

$$2\pi A \int_{-1}^1 \sin^2 \theta d \cos \theta = \frac{8\pi}{3} r_e^2 .$$

Using  $\sin^2 \theta = 1 - \cos^2 \theta$ , we have

$$2\pi A \int_{-1}^1 (1 - \cos^2 \theta) d \cos \theta = 2\pi \left(2 - \frac{2}{3}\right) A = \frac{8\pi}{3} r_e^2$$

which yields  $A = r_e^2$ .

(f) For the  $x$  polarized wave  $\theta = \pi/2 - \alpha$ , and so  $d\sigma/d\Omega = r_e^2 \cos^2 \alpha$ .

For the  $y$  polarized wave,  $\theta = \pi/2$  and we have  $d\sigma/d\Omega = r_e^2$ .